
Jet Formation in the Equatorial Oceans Through Barotropic and Inertial Instabilities

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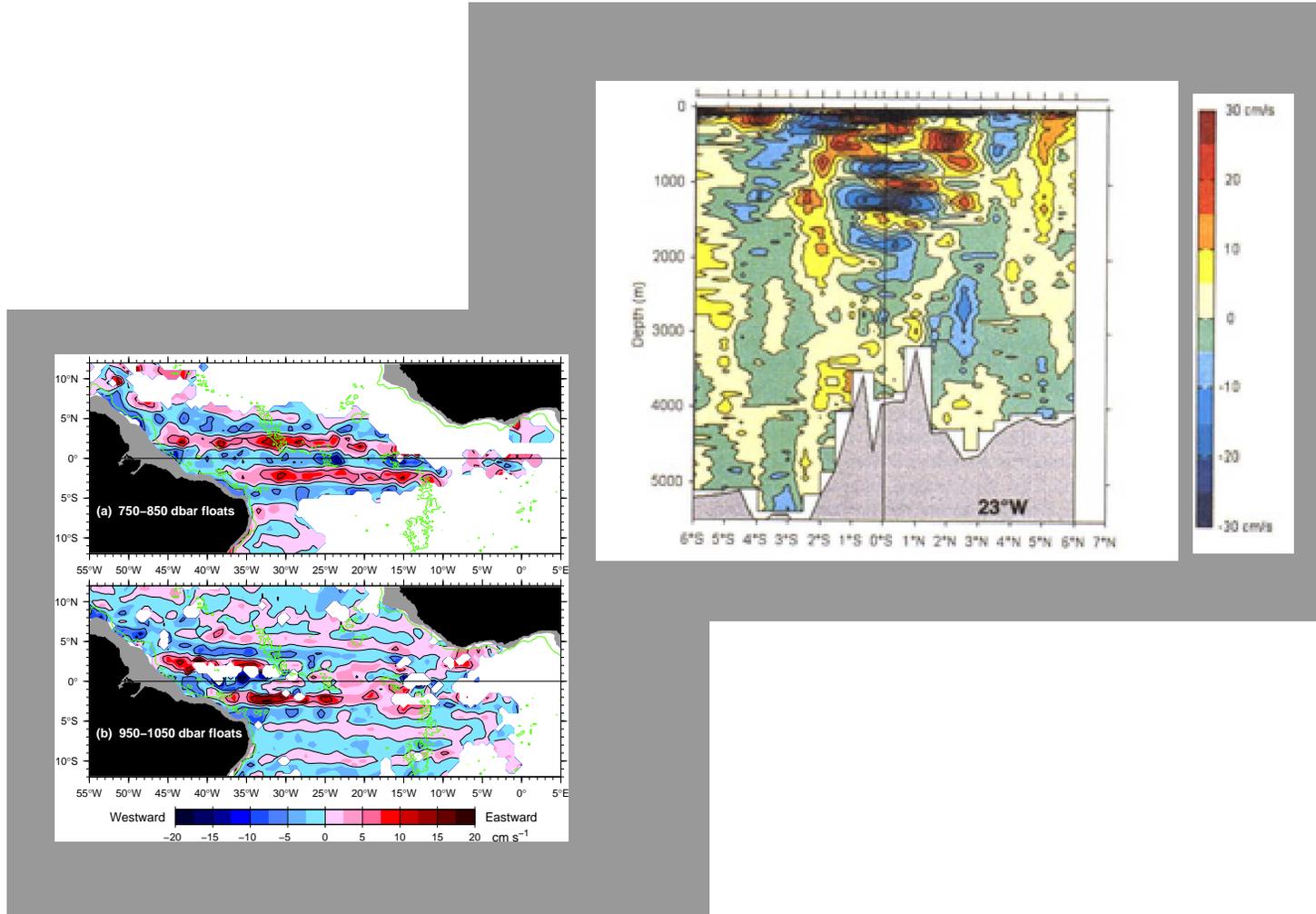
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- applications

Motivation

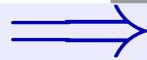
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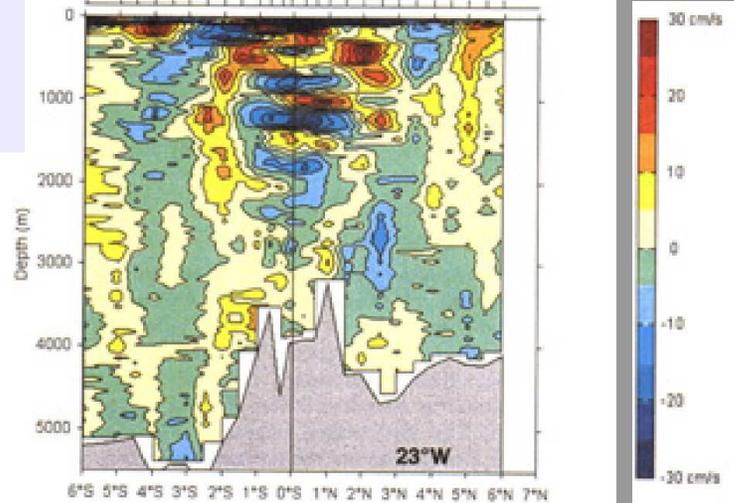
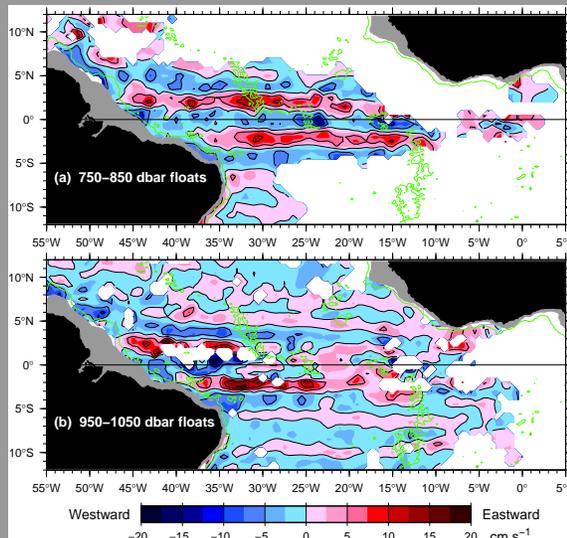
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Observations of the equatorial Atlantic:

Gouriou et al. 2001



- instantaneous merid. sect. 23° W
- strongly barotropic jets



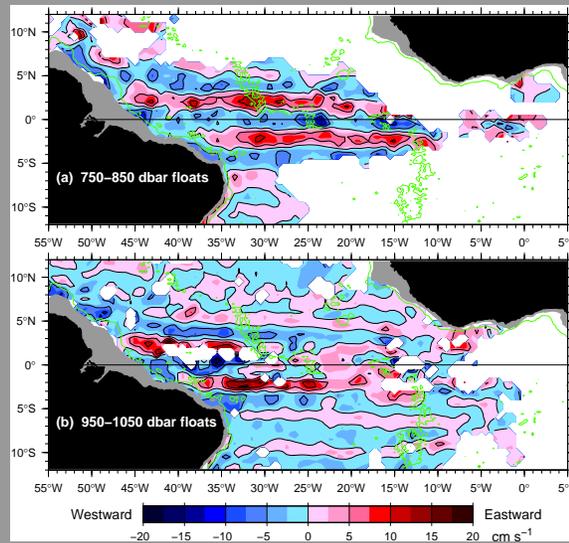
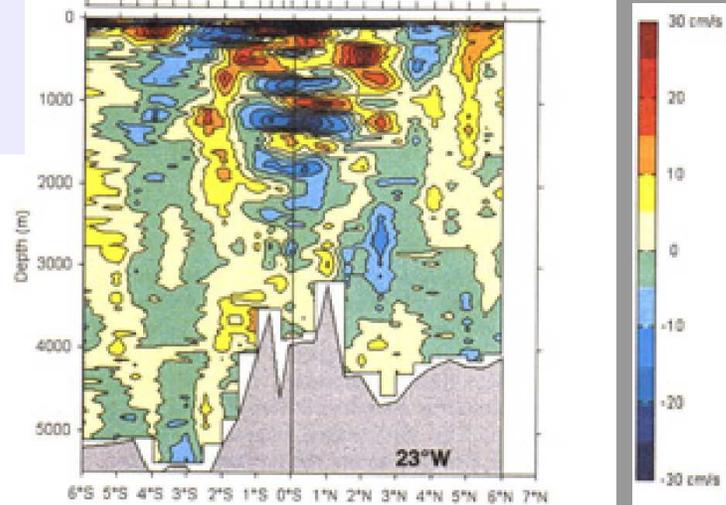
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Ollitrault et al. 2006

- zonal currents near 1000 m depth
- multiple, equally spaced jets

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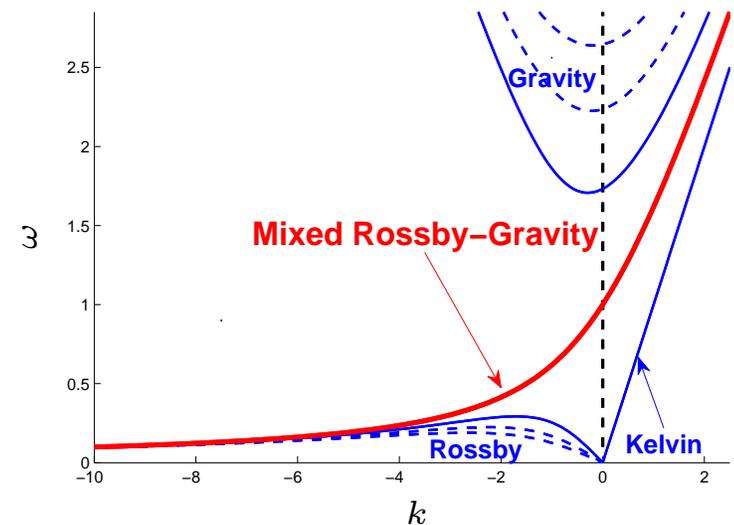
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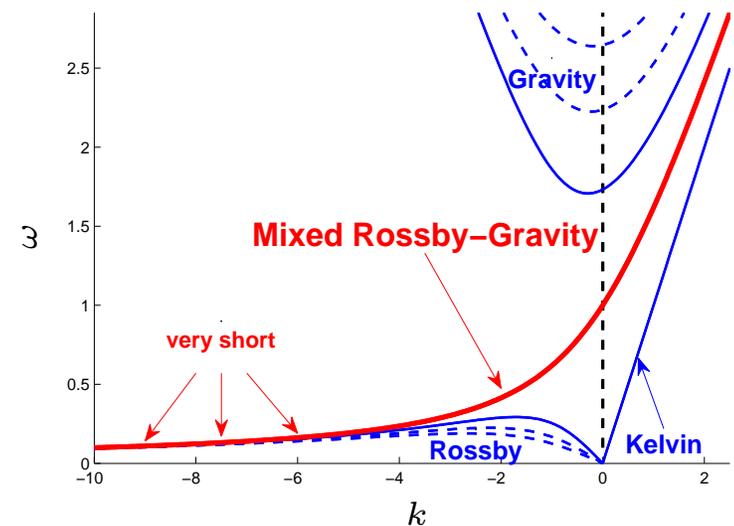
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- Dispersion diagram for meridionally confined equatorial waves:
 - Short wavelength Rossby and mixed Rossby-gravity waves have low-frequency and eastward group velocity.
 - Variability near western boundary could excite such waves which would then propagate into the interior.



Jet formation through wave instability

- In the short wavelength limit, the instability of barotropic Rossby waves (Lorenz, 1972; Gill, 1974; Lee and Smith, 2003) leads to the formation of zonal jets (Manfroi and Young, 1999).

Jet formation through wave instability

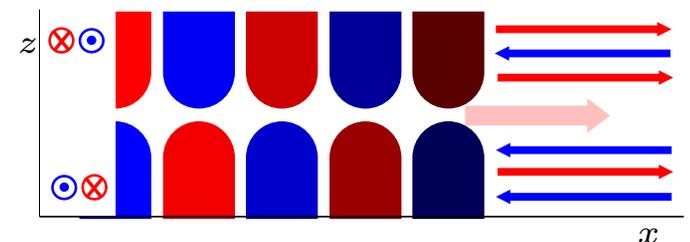
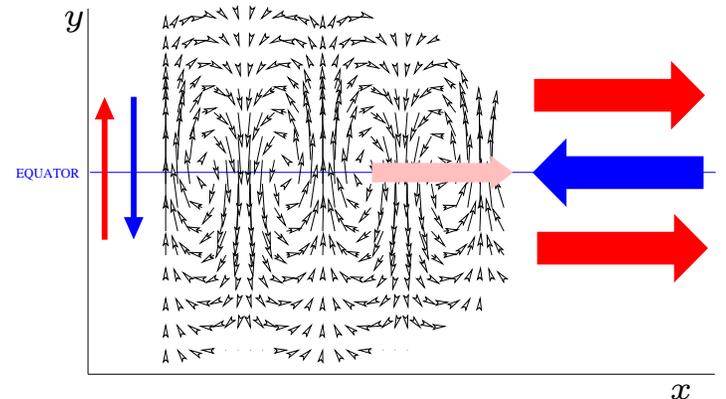
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- Numerical simulations of MRG waves excited from the western boundary show that they can lead to zonal jets:

- Baroclinic equatorial deep jets (d'Orgeville et al., 2007; Hua et al., 2008)
- Extra-equatorial barotropic jets (Ménèsquen et al., submitted)



Simulations of low frequency MRG waves

- To understand the basic mechanism performed high-resolution Primitive Equations simulations using the [Regional Ocean Modeling System \(ROMS\)](#) ([Shchepetkin and McWilliams, 2005](#)) initialized with a [mixed Rossby-gravity](#) wave in a zonally-periodic reentrant β -plane channel.

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- Parameters varied were the amplitude of meridional velocity V_0^* and the zonal wavelength λ_x , or in nondimensional terms:
 - **Froude number** $V_0 \equiv \frac{V_0^*}{c}$, **zonal wavenumber** $k = 2\pi \frac{L_D}{\lambda_x}$

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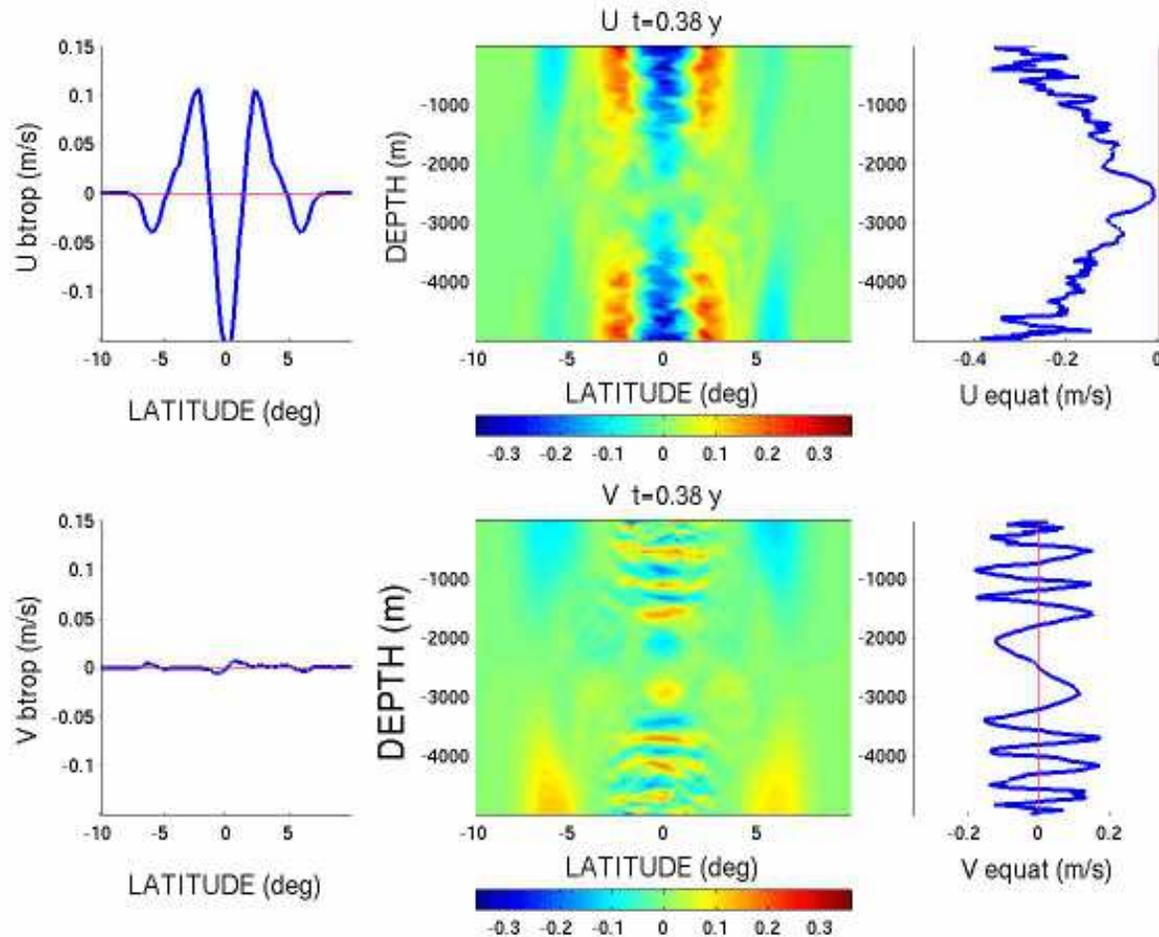
⇒ zonal shear in meridional velocity dominates the β effect (rotation) and ω (wave propagation).

Simulations of low frequency MRG waves

- Simulation with $V_0 = 0.11$ (0.36 cm amplitude) and $k = -6.3$ (350 km wavelength)
 - at $0.1^\circ \times 0.1^\circ$ horizontal resolution ($1^\circ \approx 100\text{km}$)
 - 100-200 vertical levels (25-50 m vertical resolution)

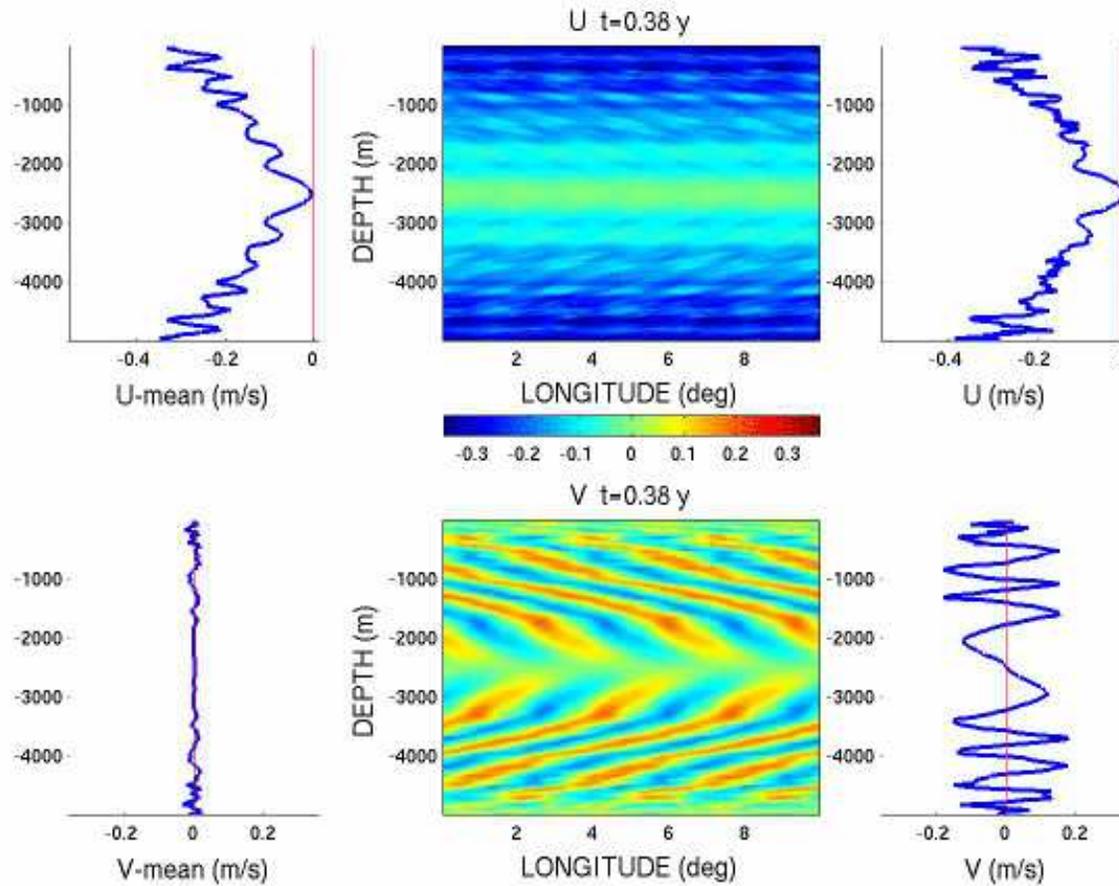
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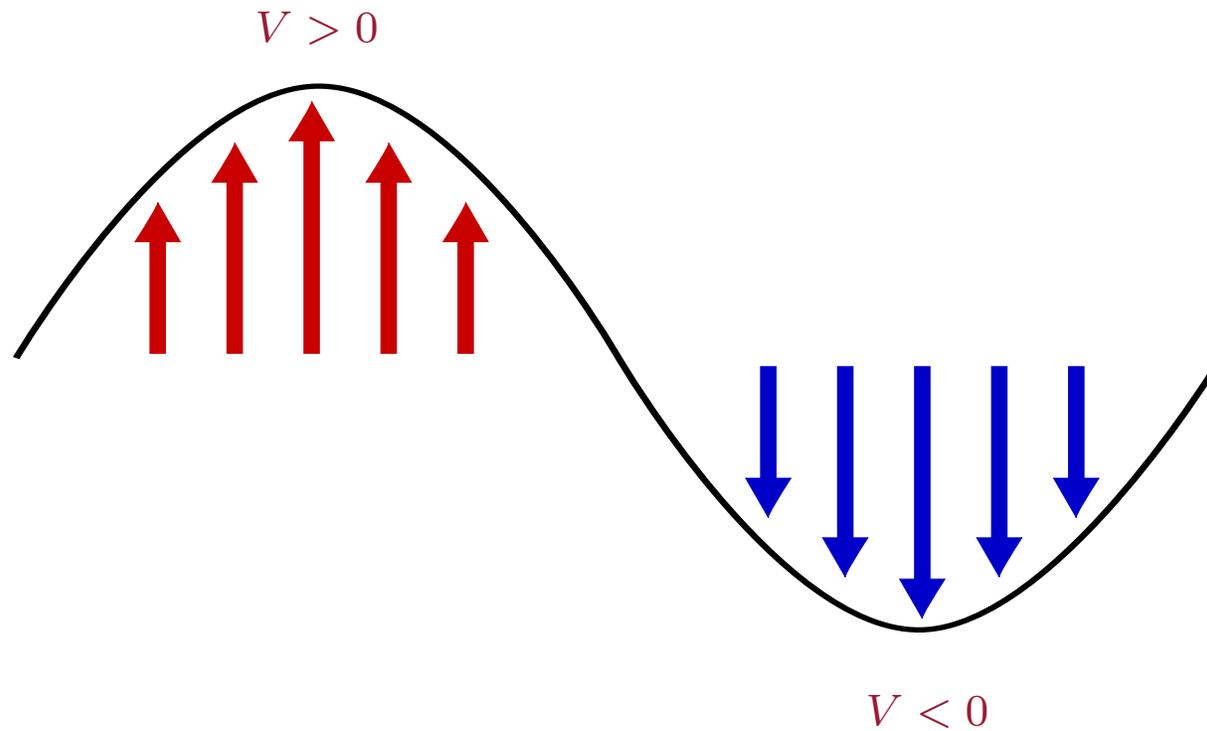


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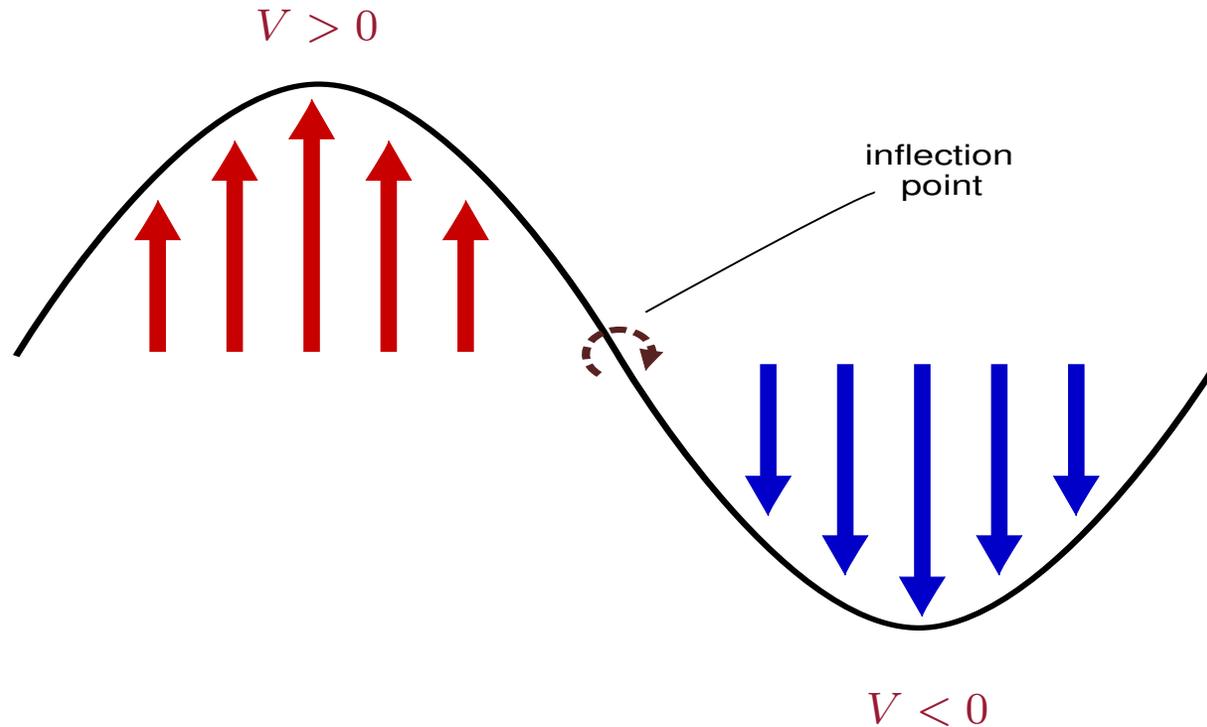


Barotropic Instability



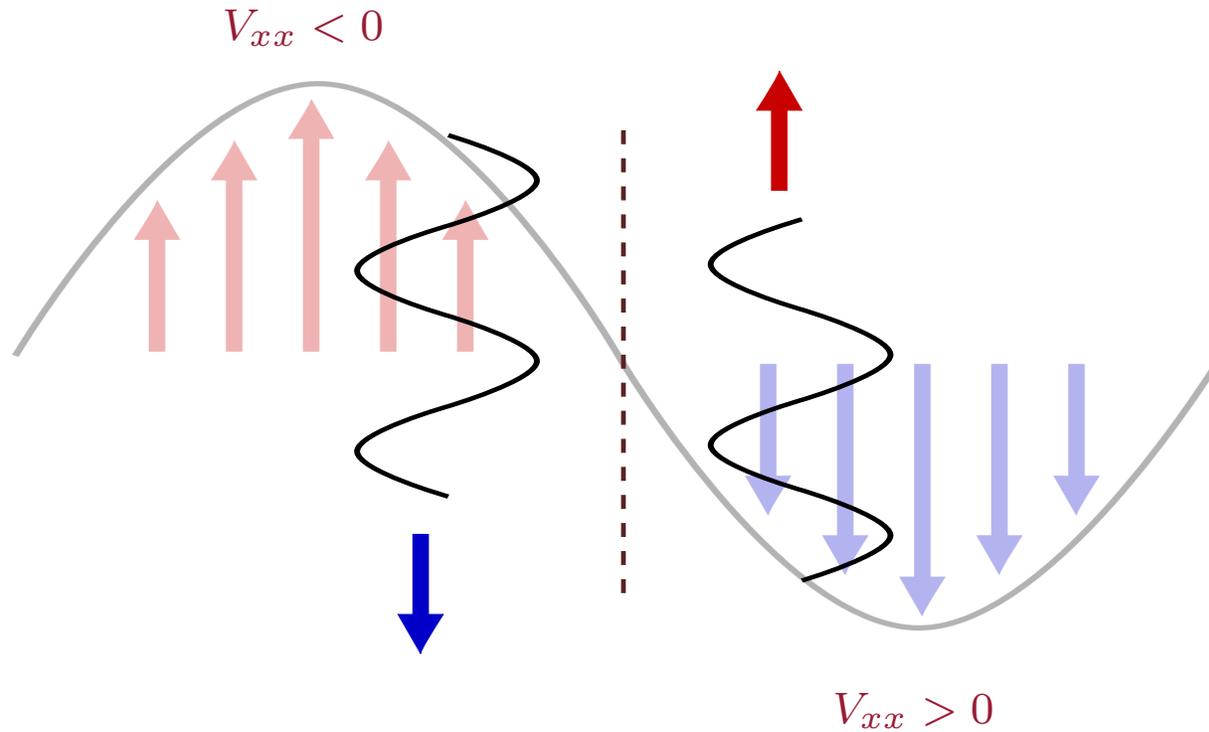
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- A **necessary condition for instability** is that the flow $V(x)$ have an inflection point (**Rayleigh**), and furthermore, that it be a maximum of absolute vorticity (**Fjortoft**).

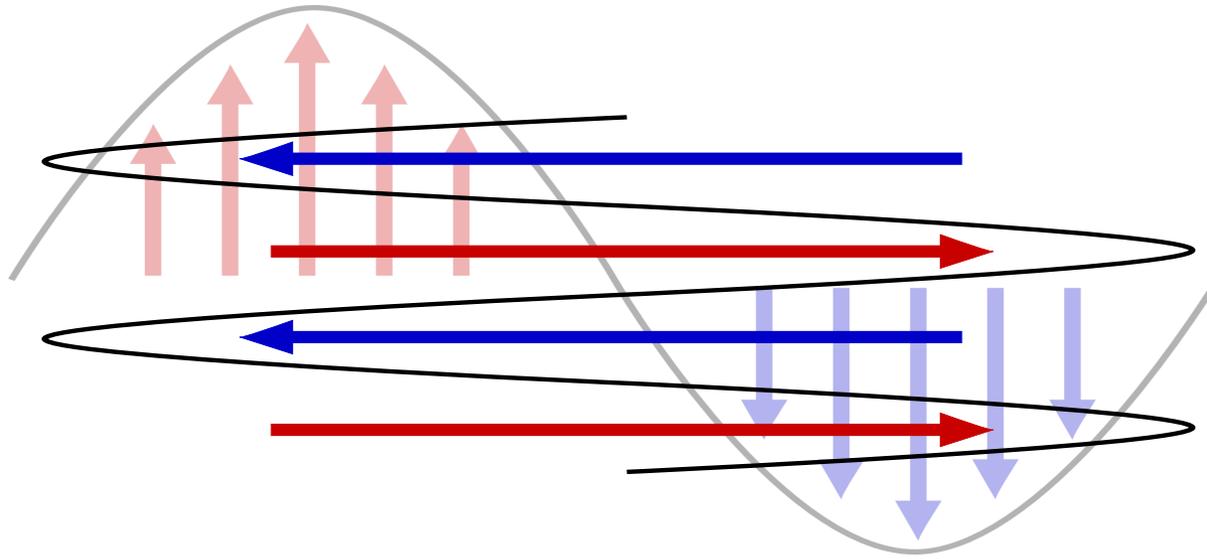
Barotropic Instability



- “Rossby” waves propagate in the positive y direction relative to background flow where $V_{xx} > 0$, and in the negative y direction where $V_{xx} < 0$.

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- For small enough values of l , waves can **phase lock** and grow exponentially.

Linear theory in $k \ll -1$ limit

● Scale length and time like $(x, y) = k^{-1}(\xi, \eta)$ and $t = (kV_0)^{-1}\tau$. Basic state wave is:

$$V \sim V_0 \exp\left(\frac{-\eta^2}{2k^2}\right) \cos(z) \cos(\xi),$$

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- Next order in k^{-1} : time dep. of MRG wave, advection by U , β effect.

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- Look for horizontally non-divergent perturbations $u' = -\psi'_\eta$, $v' = \psi'_\xi$. Then

$$\nabla^2 \psi'_\tau + \cos(z) \exp\left(\frac{-\eta^2}{2k^2}\right) \cos(\xi) (\nabla^2 \psi'_\eta + \psi'_\eta) = \frac{V_0}{k} S[\cos(2\xi), \sin(2\xi), y, z]$$

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- Seek solution with periodicity of coefficients (Floquet theorem with exponent $n_0 \equiv 0$).

$$\psi'(\xi, \eta, z, t) = \Re \left\{ e^{in_0\xi} \sum_{n=-\infty}^{\infty} \hat{\psi}_n(\eta, t; z) e^{in\xi} \right\}$$

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- A good approximation is obtained with a three term truncation:

$$\psi' \approx \Re \left\{ \hat{\psi}_{-1}(\eta, z, t) e^{-i\xi} + \hat{\psi}_0(\eta, z, t) + \hat{\psi}_1(\eta, z, t) e^{i\xi} \right\}$$

Linear theory in $k \ll -1$ limit

- Substituting the truncated series into the PDE, setting the coefficients of $e^{-i\xi}$, e^0 , and $e^{i\xi}$ equal to zero, and discretizing the differential operators in η yields:

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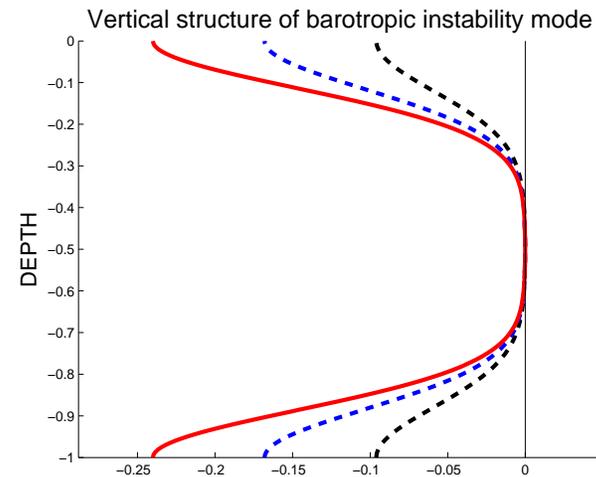
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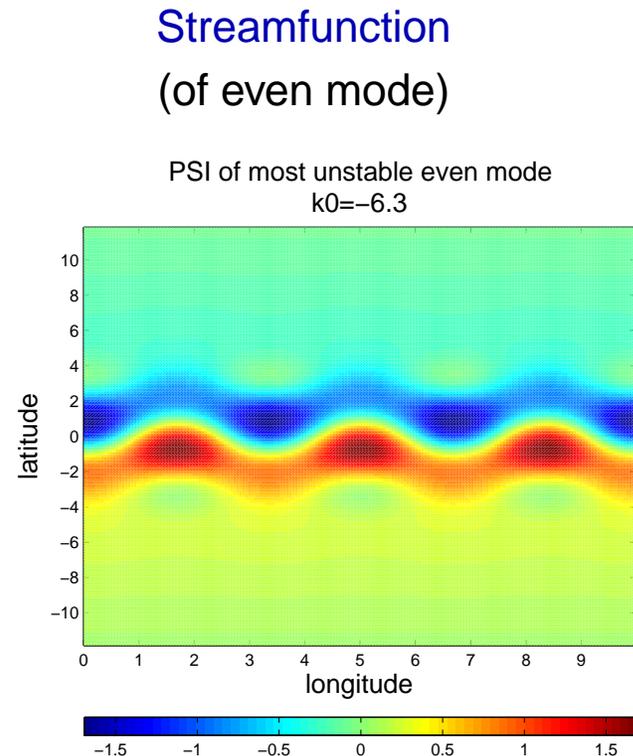
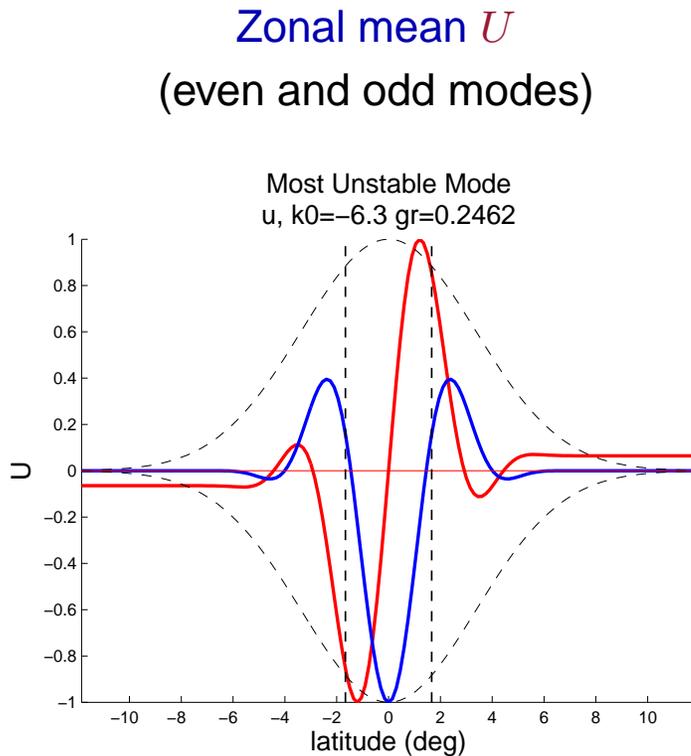
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- Right hand side \mathbf{S} acts like an initial seed for the perturbation.
- Growth rate is proportional to $\cos(z)$.



Linear theory in $k \ll -1$ limit

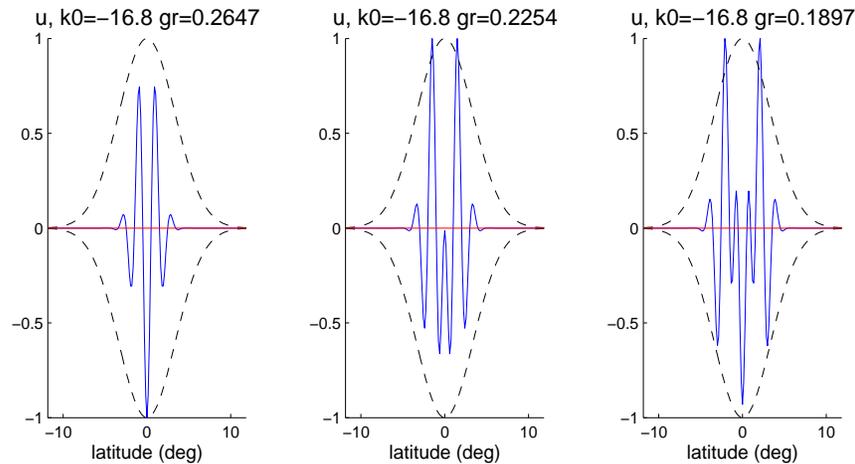
- Most unstable mode for $k = -6.3$:



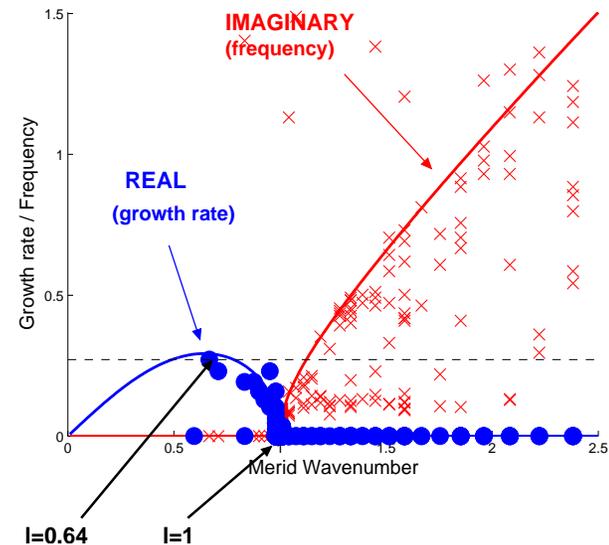
- even and odd modes come in pairs, with **odd modes** having slightly higher growth rate.

Linear theory in $k \ll -1$ limit

$u_0(y)$ for fastest growing even eigenmodes

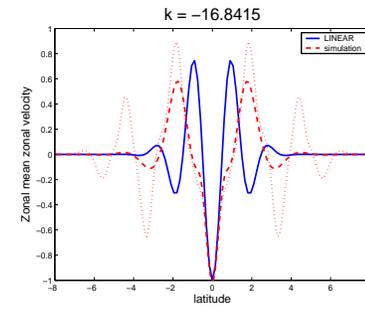
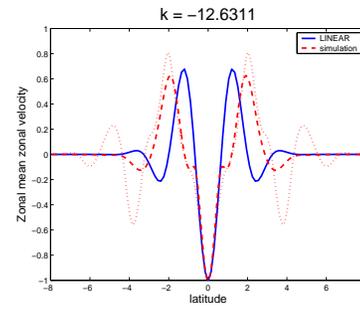
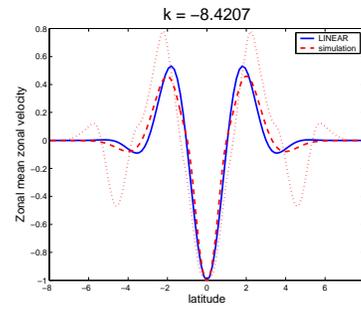
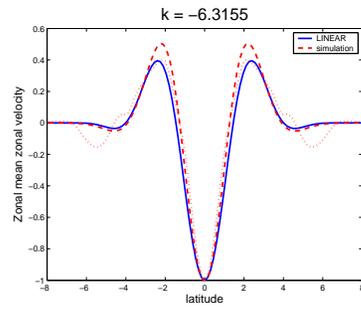


Growth rate vs. meridional wavenumber

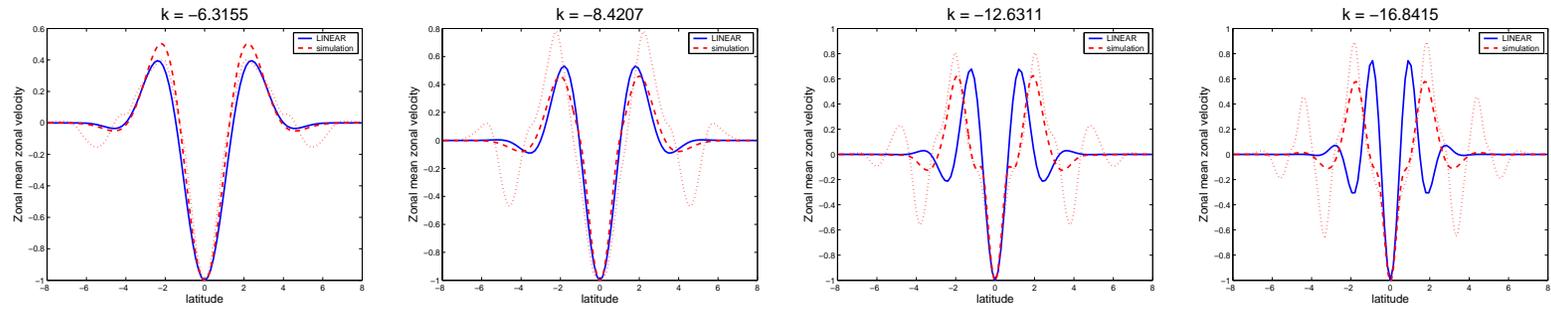


- Jet spacing near equator matches barotropic Rossby wave case (Gill, 1974).
- Jet spacing wider than wavelength of MRG wave $l < 1$.
- Fastest growing jet has width 1.6 times zonal scale of basic state wave.

Comparison with simulations

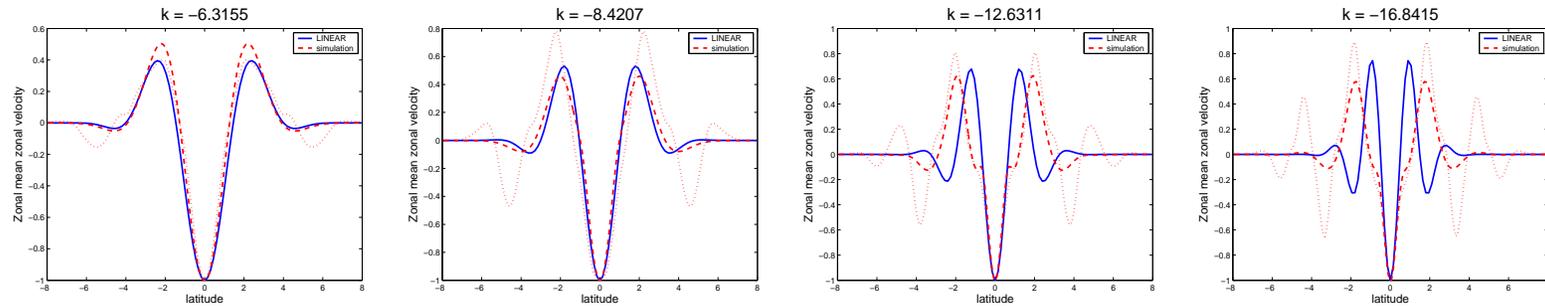


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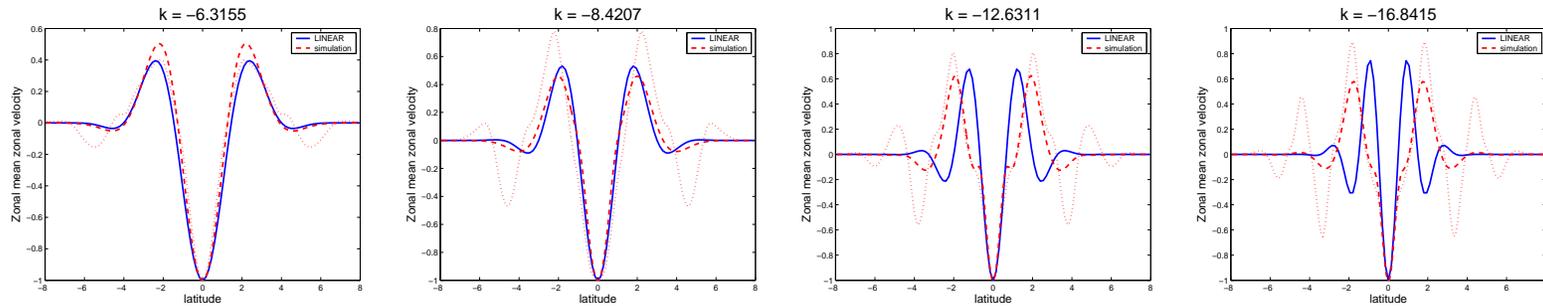
 U even about equator since odd modes inertially unstable

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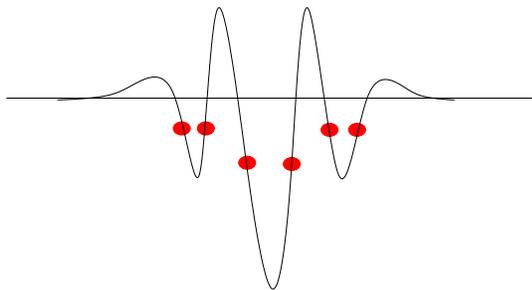


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- Extra-equatorial jet positioning poleward of that in most unstable linear mode: barotropic instability?



barotropic instability possible when $U_{yy} = \beta$
 (does not lead to meridional jets due to Coriolis effect).

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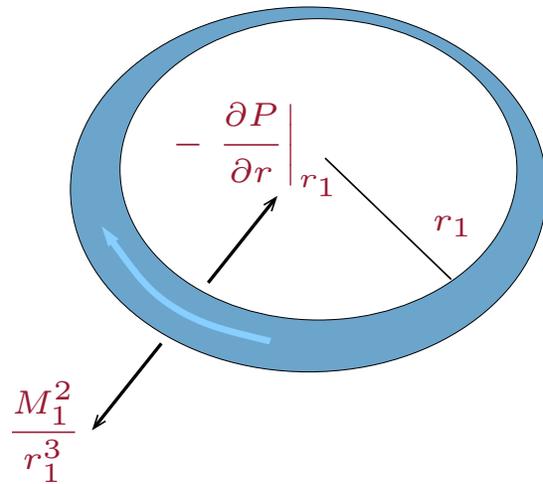
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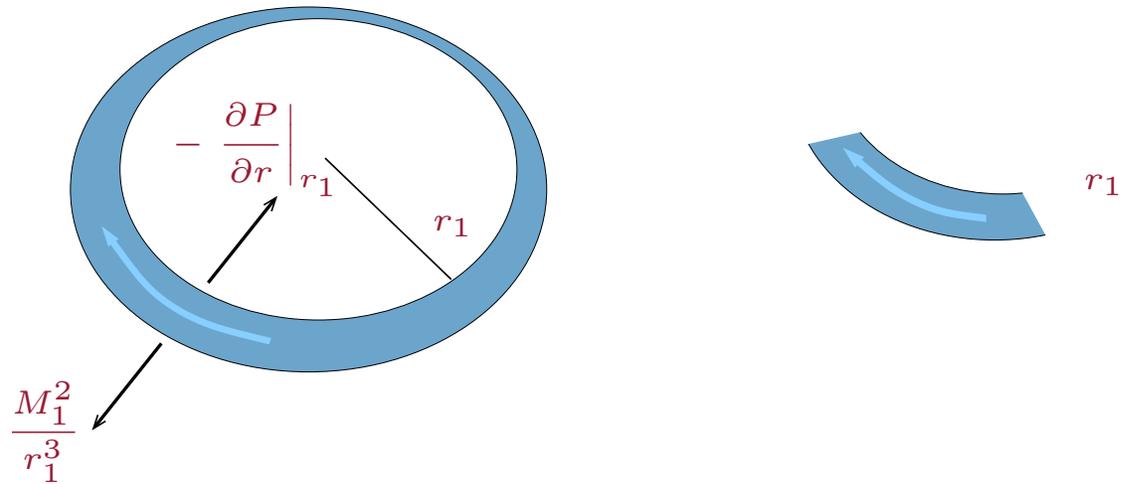
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Inertial Instability



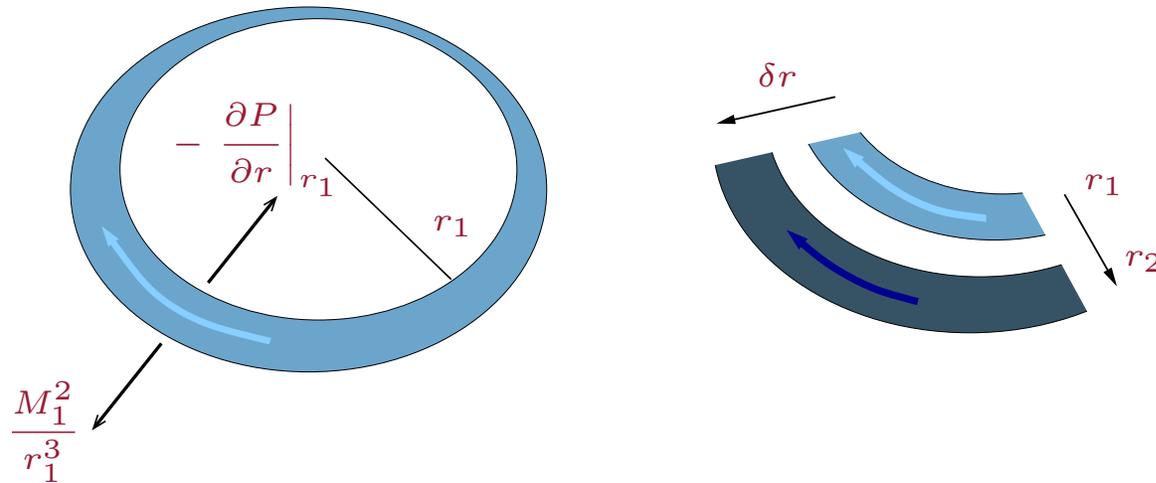
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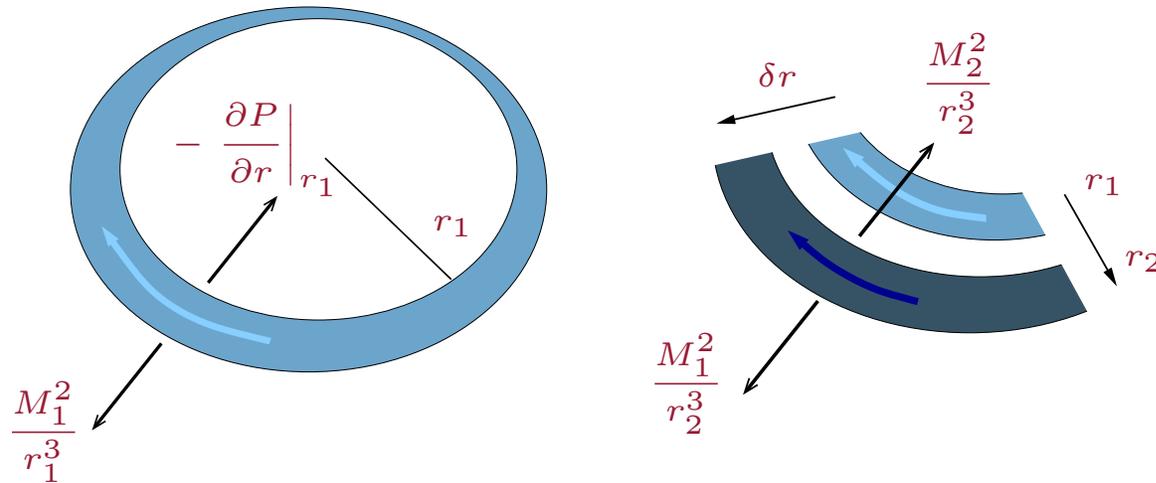
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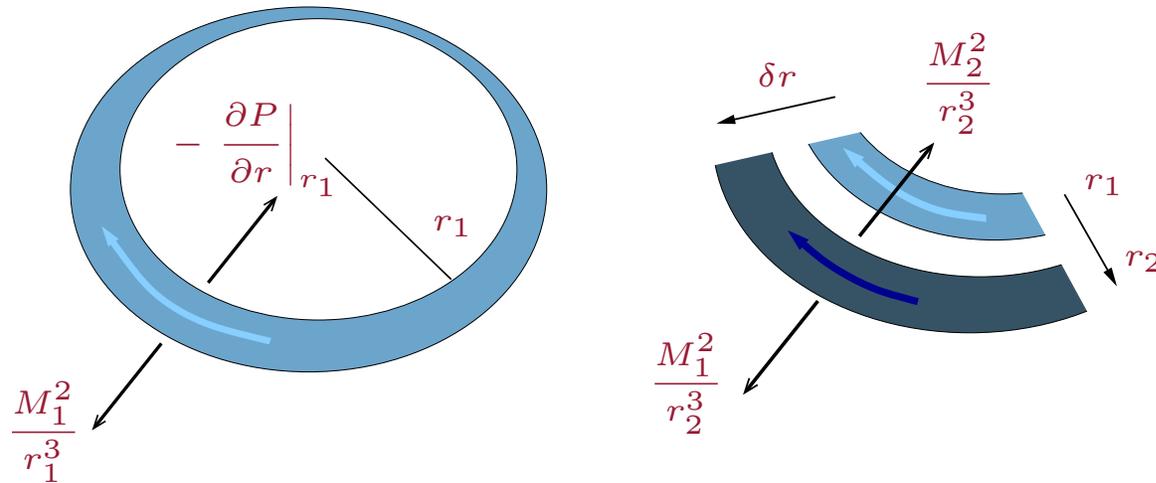
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$$\frac{M_1^2}{r_2^3} - \frac{\partial P}{\partial r} \Big|_{r_2} = \frac{M_1^2 - M_2^2}{r_2^3} \approx -\frac{1}{r_2^3} \frac{\partial (M^2)}{\partial r} \Big|_{r_2} \delta r$$

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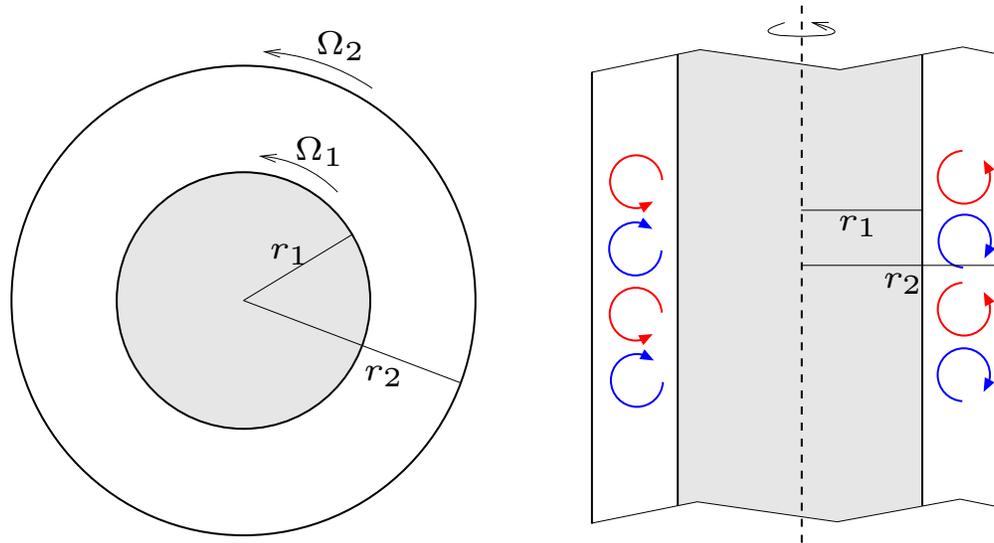


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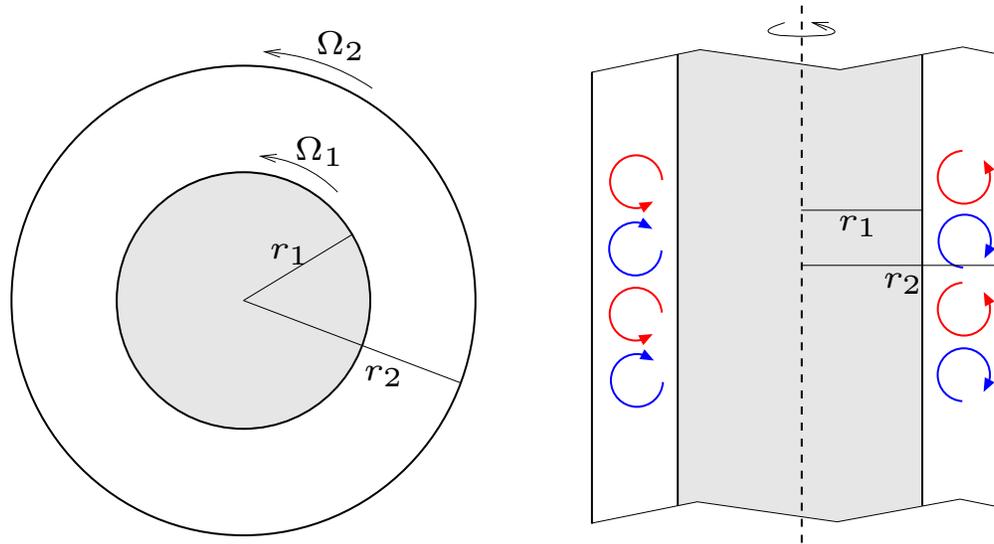
⇒ An axially symmetric flow is **inertially stable** if $\frac{\partial (M^2)}{\partial r} > 0$ (Rayleigh, 1917).

Inertial Instability



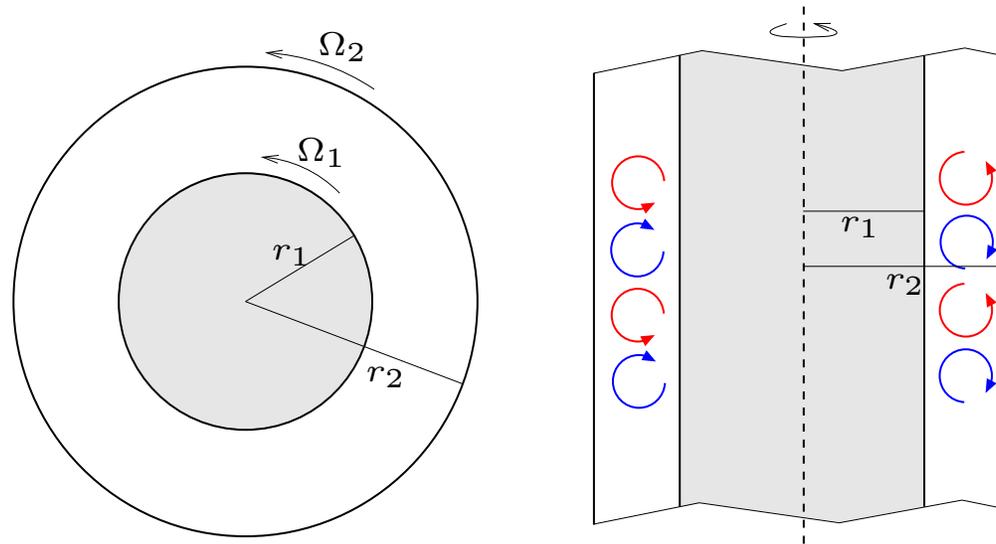
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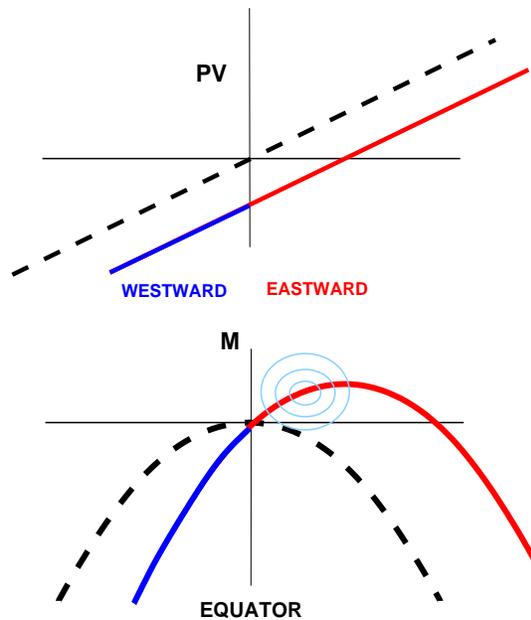


- A famous example of inertial instability is the instability of laminar flow in the [Taylor-Couette experiment](#), when outer cylinder rotates too fast relative to inner cylinder (Taylor (1923)).
- First bifurcation leads to [Taylor vortices](#) in the radial-vertical plane.
- For inertial stability on the earth, maximum angular momentum must be [at the equator](#).
- Combined with effects of [stratification](#), the condition for stability is that [Potential Vorticity](#) have the sign of latitude (“[symmetric stability](#)”).

Inertial Instability

- For stability, maximum angular momentum $M \approx U - \frac{1}{2}\beta y^2$ must be at the equator:

Equatorial shear inertial instability

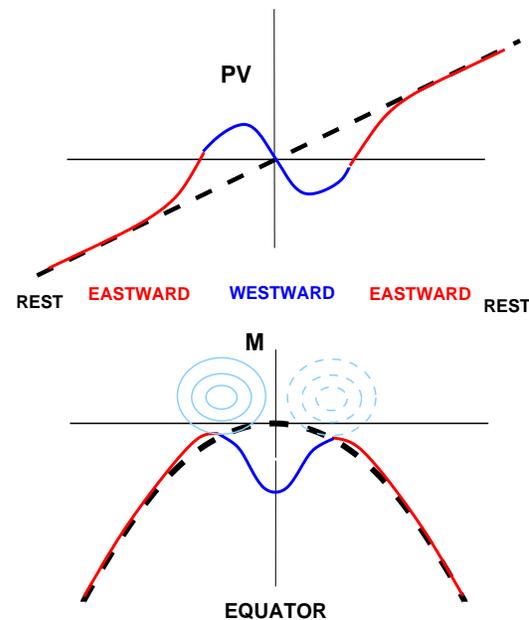


(Dunkerton, 1981)

$$U = \lambda y, \text{ unstable for all } \lambda \neq 0$$

Adjustment mixes in latitude and depth.

“Curvature” inertial instability



$$U = (-U_{00} + \frac{1}{2}by^2)e^{-\frac{1}{2}y^2},$$

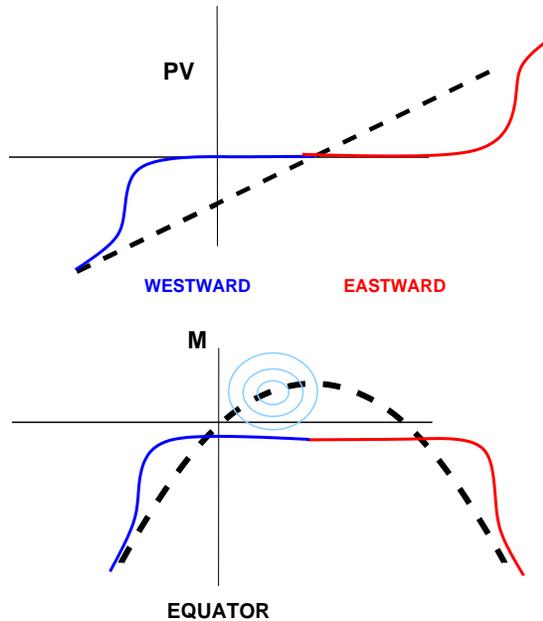
stable for $b > \beta$

un-

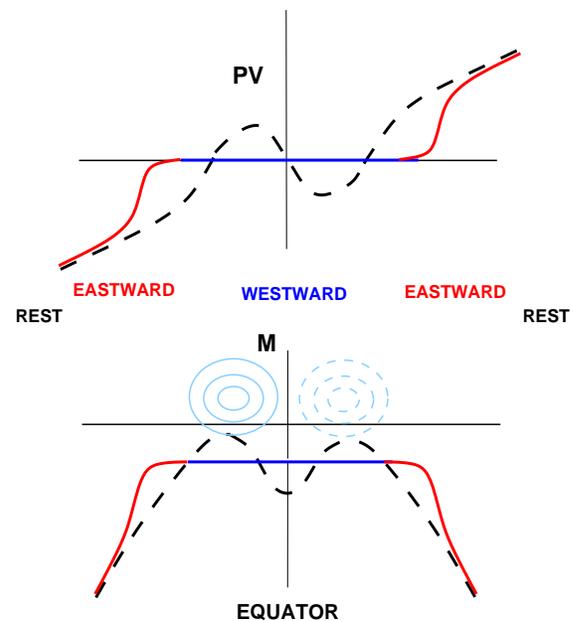
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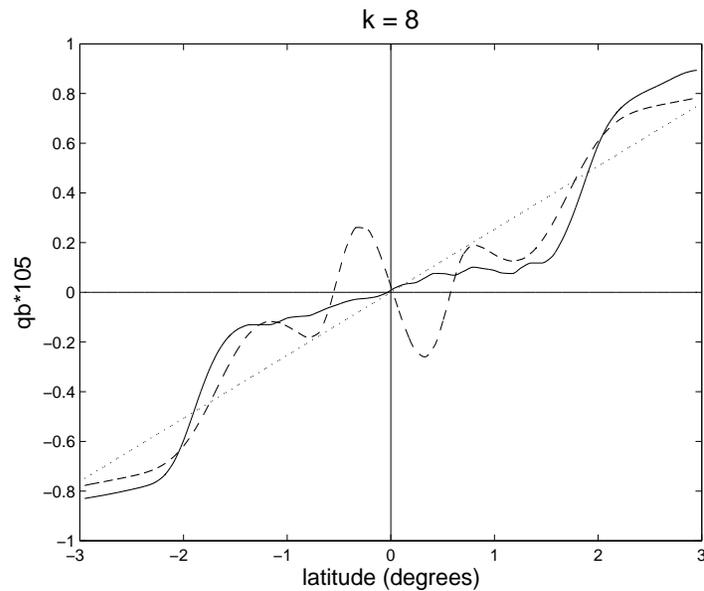


- inertial adjustment takes place over wider latitude interval than covered by initial instability (Hua et al., 1997, Griffiths, 2003).
- Results in wide area of zero PV .

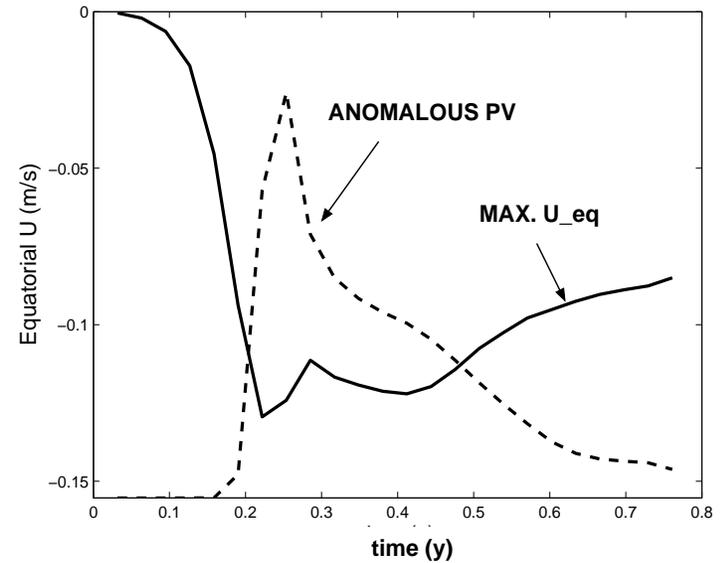
Inertial instability in simulations

- Evidence of inertial instability in the simulations:

Absolute vorticity vs. latitude



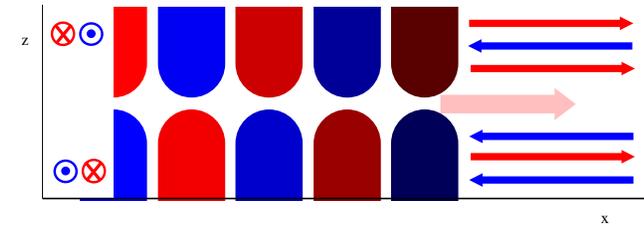
Jet amplitude and anomalous PV



- Also, energy increase in zonally symmetric components of v and w fields.

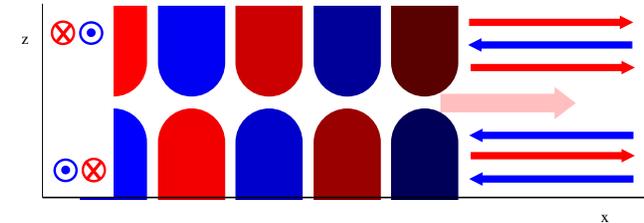
Baroclinic equatorial deep jets

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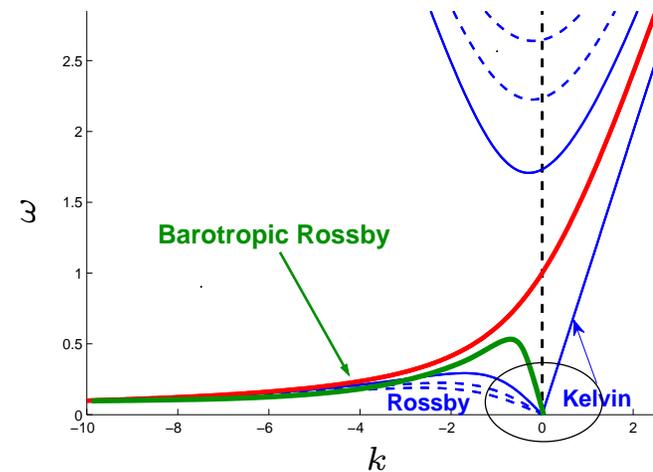
- Hua et al., 2008:

- Since the destabilization is taking place mostly in the western edge of the basin, what is observed in the remainder of the basin is the projection of the jets pattern onto low-frequency, long zonal wavelength waves with eastward group velocities.

- The only such wave is the equatorial Kelvin wave, which has

$$u(y, z) \propto \cos(mz) \exp \left[- \left(\frac{\beta m}{2NH} \right) y^2 \right]$$

- The vertical mode m of the observed jets corresponds to the vertical mode of the Kelvin wave with meridional width $\sqrt{NH/\beta m}$ comparable to the width of the most unstable barotropic jet mode, namely $m \propto k^2 \propto T^2$.

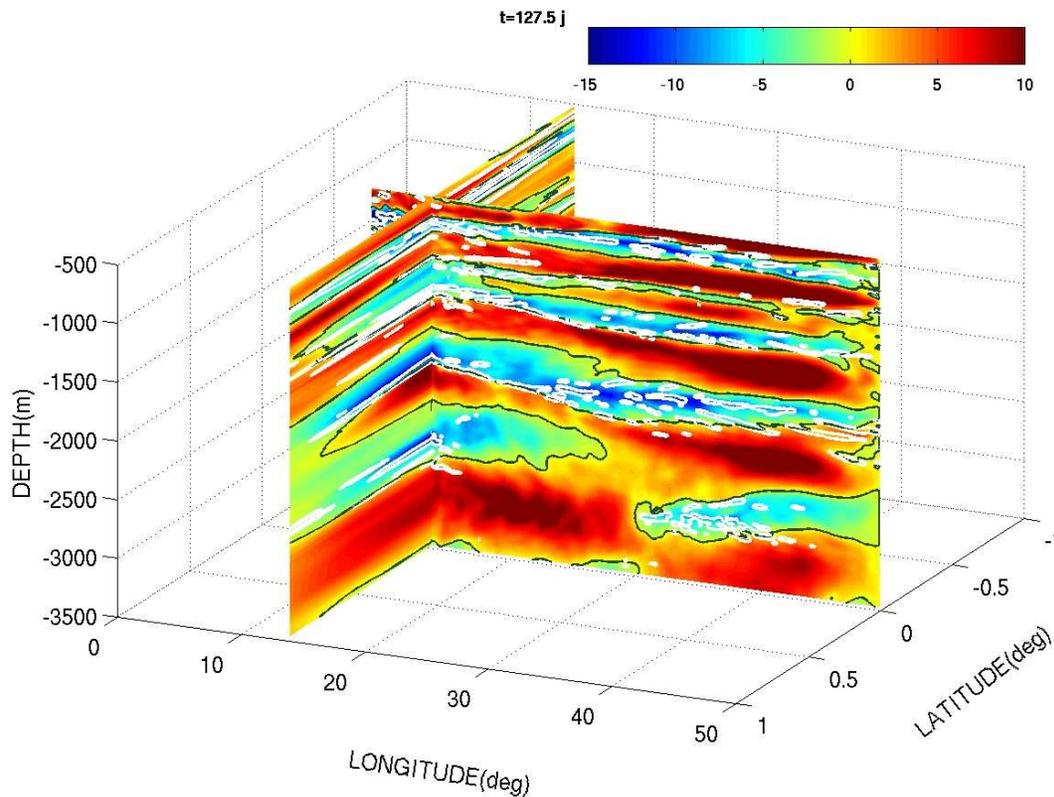


Layering, zonal jets, and inertial instability

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(Ménèsquen et al., submitted)

- Regions of vertical mixing are concentrated in **westward jets**, and are strongly correlated with the criterion for marginal inertial instability.

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Continuing Work

- Very short wavelength waves lead to **local super-rotation**. Application to planetary atmospheres?
- Baroclinic effects of depth-dependent triggering of inertial instability.
- The effect of the **non-traditional Coriolis force**? Simulations show strong vertical symmetry breaking. Evidence of global conservation of angular momentum for wave momentum flux?

