Stratospheric dynamics, ozone and climate

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Outline

- The wave-driven meridional circulation
- Stratospheric planetary waves
 - \rightarrow Charney-Drazin theorem
- Polar vortex dynamics
 - \rightarrow Stratospheric sudden warmings
- Brewer-Dobson circulation
 - → Seasonal cycle in extratropical ozone
- Chemistry-climate coupling
- Long-term changes in ozone
- Climate change in the stratosphere

The wave-driven diabatic circulation

- Radiation tends to relax the atmosphere towards temperature $T_{\rm rad}$ and zonal wind $u_{\rm rad}$, in thermal-wind balance
 - \rightarrow In troposphere, $T_{\rm rad}/u_{\rm rad}$ is convectively and baroclinically unstable; leads to vigorous motion
 - \rightarrow In middle atmosphere, $T_{\rm rad}/u_{\rm rad}$ is dynamically stable, so represents a possible quasi-steady state
- Being isolated from the Earth's surface, the middle atmosphere has little thermal inertia
 - \rightarrow Radiative relaxation time is at most a few weeks
 - → Hence the flow tends to be far more zonally symmetric than the tropospheric flow, making the zonal average more physically meaningful
 - → One expects the lowest temperatures over the winter pole and the highest temperatures over the summer pole
 - \rightarrow Thus eastward flow in the winter hemisphere, westward flow in the summer hemisphere
- Zonal motion is "free", but rotation and stratification restrict meridional and vertical motion, respectively
 - → Persistent meridional motion requires a torque
 - → Persistent vertical motion requires diabatic heating

- In fact, a persistent meridional circulation is observed
 - \rightarrow Why is it there?
- In the steady limit, the angular momentum balance is $f\bar{v}^* \approx -\nabla \cdot \pmb{F} \quad \text{(EP flux convergence; a.k.a. "wave drag")}$
 - → Eastward (positive) torque drives equatorward flow, while westward (negative) torque drives poleward flow
 - → Angular momentum is *conserved* (even with viscosity)
 - $\rightarrow \nabla \cdot \boldsymbol{F}$ associated with waves; not a property of the medium
- In the steady limit, the thermodynamic balance is

$$\frac{N^2 T_0}{g} \, \bar{w}^* \approx Q \approx -r(T - T_{\rm rad})$$

- \rightarrow Downwelling implies $T > T_{\rm rad}$, upwelling implies $T < T_{\rm rad}$
- → Atmosphere is energetically open
- \bar{v}^* and \bar{w}^* constrained by mass continuity, which couples $\nabla \cdot \boldsymbol{F}$ and $T T_{\text{rad}}$; but which causes which?
- Without wave drag, there would be no radiative heating/cooling (apart from transient effects) and $T \approx T_{\rm rad}$
 - → Wave drag is a "gyroscopic pump"
 - → Radiation doesn't drive the meridional circulation, it accommodates it (it's not that "hot air rises")

- Since middle atmosphere is dynamically stable, $\nabla \cdot \boldsymbol{F}$ is associated with waves propagating up from troposphere
 - → Unlike troposphere, middle atmosphere is a "refrigerator"
- In the steady limit, can consider $T = \underbrace{T_{\text{rad}}}_{\text{radiation}} + \underbrace{(T T_{\text{rad}})}_{\text{wave drag}}$
 - \rightarrow In global mean, $\langle T \rangle \approx \langle T_{\rm rad} \rangle$ at every pressure level

Stratospheric planetary waves

• Consider (for simplicity) the 2-D (barotropic) β -plane eqns $\frac{\partial \nabla^2 \psi}{\partial t} + \boldsymbol{v} \cdot \boldsymbol{\nabla} (\nabla^2 \psi + \beta y) = 0 \quad \text{where} \quad \boldsymbol{v} = \left(-\frac{\partial \psi}{\partial y}, \frac{\partial \psi}{\partial x} \right),$ linearized about a slowly-varying zonal flow \bar{u} :

$$\left(\frac{\partial}{\partial t} + \bar{u}\frac{\partial}{\partial x}\right)\nabla^2\psi' + \beta\frac{\partial\psi'}{\partial x} = 0 \quad (\nabla^2\psi + \beta y = \text{absolute vorticity})$$

- \rightarrow One obtains the dispersion relation $c = \bar{u} \frac{\beta}{k^2 + \ell^2}$
- \rightarrow These are westward propagating waves (relative to \bar{u})

$$\beta \sim 10^{-11} \text{m}^{-1} s^{-1}, \quad k \sim \ell \sim 10^{-6} \text{m}^{-1} \quad (\lambda \sim 6000 \, \text{km})$$

$$\Rightarrow \frac{\beta}{k^2 + \ell^2} \sim 10 \,\mathrm{m\,s^{-1}}$$
, comparable to $\bar{u} \sim 10 \,\mathrm{m\,s^{-1}}$

→ Hence large-scale stationary forcing in an eastward flow efficiently generates planetary-scale Rossby waves

- Now, $\operatorname{sgn}(\overline{u'v'}) = -\operatorname{sgn}(k\ell)$ and $c_{g(y)} = \frac{2\beta k\ell}{(k^2 + \ell^2)^2}$
 - ⇒ meridional momentum flux opposite to energy flux
 - ⇒ Rossby waves carry negative momentum
- There is another way of seeing this
 - → Consider the 2-D zonal-mean momentum equation
 - → Pressure-gradient and Coriolis terms drop out, leaving

$$\frac{\partial \bar{u}}{\partial t} = -\frac{\partial}{\partial y} (\overline{u'v'}) = \overline{\psi'_{yy}\psi'_{x}} + \overline{\psi'_{y}\psi'_{xy}} = \overline{\nabla^{2}\psi'\psi'_{x}} - \overline{\psi'_{xx}\psi'_{x}}$$

$$= -\frac{1}{\beta} \overline{\nabla^{2}\psi'\left(\frac{\partial}{\partial t} + \bar{u}\frac{\partial}{\partial x}\right)\nabla^{2}\psi'} = \frac{\partial}{\partial t} \left(-\underbrace{\frac{1}{2\beta} \overline{(\nabla^{2}\psi')^{2}}}_{= 4}\right)$$

- → So as Rossby waves enter a region, they decelerate the flow; as they leave it, they accelerate it
- \rightarrow This effect is illustrated in a laboratory experiment by Whitehead (1975 *Tellus*)
- All this goes through for 3-D (baroclinic) flow, where A is the *Eliassen-Palm wave activity* (a.k.a. negative of the *pseudomomentum*), and for stationary waves

$$\bar{u} = \frac{\beta}{k^2 + \ell^2 + (f^2 m^2 / N^2)}$$

- \rightarrow No propagation is possible in the summer stratosphere where $\bar{u}<0$
- \rightarrow For $\bar{u}(y,z)$ (slowly varying), ℓ and m will evolve as a wave packet propagates, while k is fixed, but clearly

$$\bar{u} < U_c(k) \equiv \frac{\beta}{k^2}$$

- \rightarrow Hence if u becomes more positive with altitude, as in the winter stratosphere, the higher-k waves become evanescent
- ightharpoonup Zonal wave 2: $k \sim \frac{2\pi}{12\,000\,\mathrm{km}} \quad \Rightarrow \quad U_c \sim 45\,\mathrm{m\,s^{-1}}$
- ightarrow Zonal wave 3: $k \sim \frac{2\pi}{8\,000\,\mathrm{km}} \quad \Rightarrow \quad U_c \sim 20\,\mathrm{m\,s^{-1}}$
- → Only planetary waves 1 and 2 reach the stratosphere (Charney-Drazin theorem)
- We are now in a position to understand the stratospheric meridional circulation
 - → Persistent planetary-wave forcing gives a negative angular momentum forcing in the winter hemisphere
 - \rightarrow Negative angular momentum forcing must translate into a poleward mass flux

- → Circulation must return equatorward somewhere, but needs an opposite-signed momentum forcing to do so
- → To the extent that wave momentum transfer is given, this can only happen in the planetary boundary layer ("downward control")
- \rightarrow This drives upwelling in the tropics, and downwelling in the extratropics, below the forcing level
- This understanding of the meridional circulation explains why the circulation is dominantly in the winter hemisphere
- It also explains the NH-SH asymmetry
 - \rightarrow The NH generates stronger planetary Rossby waves than does the SH
 - \rightarrow The lack of large continental land masses in the SH leads to much lower amplitude planetary waves
 - \rightarrow This leads to a stronger poleward mass flux in the NH than in the SH
 - → It also implies more tropical upwelling, and lower tropical lower stratospheric temperatures, in NH winter as compared with SH winter
 - → Accounts for the "tropical tape recorder" in water vapour

Polar vortex dynamics

- Polar downwelling from planetary-wave drag warms and weakens the winter polar vortices from radiative equilibrium
 - → More so in the NH than in the SH
 - \rightarrow Stronger downwelling implies higher polar T, so a weaker meridional gradient of T, so a weaker eastward circumpolar flow
- The stronger Arctic planetary-wave forcing leads to greater variability in the Arctic; there is a nonlinear feedback
 - → Virtually no variability in the quiescent summertime
- Sometimes the forcing is so strong that polar temperatures rise by several tens of degrees K in a few days, and the vortex becomes westward
 - \rightarrow Called a "stratospheric sudden warming"
 - → Cannot be explained by radiation or horizontal advection
 - \rightarrow Rare in the SH; first ever observed was in 2002
- Sudden warmings can also be understood as the balanced response to negative angular momentum deposition by planetary waves

- \rightarrow Need to consider the transient rather than the steadystate response
- Where planetary waves are dissipated they exert a westward torque
 - → The torque cannot go entirely into a westward acceleration, because this would violate thermal-wind balance
 - → The response goes partly into zonal-wind deceleration, and partly into an instantaneous meridional circulation
 - → Alternatively, the torque goes partly into the relative and partly into the planetary angular momentum
 - → The meridional circulation induces a temperature tendency through adiabatic heating/cooling which is in thermal-wind balance with the zonal-wind tendency
 - \rightarrow For fixed forcing this leads to a steady response:

$$\frac{\partial T}{\partial t} + r(T - T_{\text{rad}}) = -\frac{N^2 T_0}{g} \, \bar{w}^* \equiv W = \text{const.}$$

$$\implies T = T_{\text{rad}} + \frac{W}{r} (1 - e^{-rt})$$

• Normally, stationary planetary waves propagate towards the equator and dissipate in midlatitudes

- → Sudden warmings occur when the wave dissipation manages to penetrate into the Arctic vortex, and the induced downwelling is focused over the pole
- → When the vortex becomes westward, wave propagation shuts down and the vortex recovers radiatively
- There is another interpretation of sudden warmings in terms of 2-D vortex dynamics (corresponding to an isentropic layer)
 - \rightarrow We know that absorption of Rossby waves must lead to a reduction in angular momentum
 - → In planar geometry, angular momentum is given by

$$\iint \hat{\boldsymbol{z}} \cdot (\boldsymbol{r} \times \boldsymbol{v}) \, dx dy = -\frac{1}{2} \iint r^2 \omega \, dx dy \quad (\omega = \text{vorticity})$$

→ In spherical geometry, it is given by

$$\iint \sin \phi \, q \, d(\sin \phi) d\lambda \quad (q = \text{potential vorticity});$$

note that $\sin \phi \approx 1 - \frac{1}{2}(\phi')^2$ near the pole, $\phi' \equiv \frac{\pi}{2} - \phi$

- \rightarrow Since either ω or q are materially conserved, a reduction in angular momentum with the same ω or q requires a deformation of the vortex
- → And indeed sudden warmings are associated with a break-up of the polar vortex

Brewer-Dobson circulation

- Contrast in mixing timescales between troposphere and stratosphere ⇒ tropopause is a distinct chemical boundary for long-lived species
 - → Analogous to top of the planetary boundary layer
- Chemical measurements have always played a crucial role in understanding the stratospheric circulation
 - → Within the stratosphere, the mean diabatic circulation also transports chemical species
 - → Poleward sense of circulation originally inferred from water vapour (Brewer 1949) and ozone (Dobson 1956)
 - \rightarrow Clearly visible in long-lived species, e.g. CH₄ and N₂O
- Implies that the tropics are the entry point for the stratosphere, and the extratropics the departure point
 - → Diabatic (mass) circulation represents the "advective" part of the flux of species
 - → Has overturning time scale of several years
 - \rightarrow Acts to steepen meridional gradients of long-lived species

- There is also a contribution due to mixing (with no net mass flux)
 - \rightarrow For planetary waves, it is closely coupled to transport since $\nabla \cdot F = \overline{v'q'}$ (q is potential vorticity)
 - → Acts to flatten meridional gradients of long-lived species
 - \rightarrow Observed distributions represent a balance between these two effects
- Because of mixing, there is a distribution of transit times to any particular region of the stratosphere
 - → Mean of the distribution is the "mean age"; oldest in polar upper stratosphere
- Horizontal mixing is strongest in the "surf zone" of the winter hemisphere, associated with breaking planetary waves
 - \rightarrow Surf zone q stirring drives the Brewer-Dobson circulation
 - \rightarrow Deformation of q is associated with absorption of the negative angular momentum of the planetary waves, as with the vortex break-up in sudden warmings
 - → Horizontal inhomogeneity of mixing leads to "mixing barriers": polar vortex edge and tropical pipe

- Seasonality of stratospheric planetary-wave drag implies a seasonality in the Brewer-Dobson circulation
 - \rightarrow Leads to spring buildup of extratropical column O₃ (more so in NH than in SH)
 - → Ozone decays photochemically in the quiescent summer
 - → Summertime ozone more similar between NH and SH
- Interannual variability of planetary-wave drag implies interannual variability in the wintertime ozone buildup
 - \rightarrow Variations in tropical and extratropical O₃ anti-correlated
 - \rightarrow More O_3 buildup in spring implies more chemical O_3 loss in summer you can only destroy what you have
 - → There is a remarkable seasonal persistence in the interannual ozone anomalies
 - → Midlatitude and polar variations are highly correlated

Chemistry-climate coupling

- Relationship between diabatic circulation and transport, via wave drag, is crucial to understanding cause and effect
 - \rightarrow Important for attribution of observed record
 - → Important for diagnosing causes of model errors

An example: If stratospheric wave drag decreased then:

- → Diabatic and Brewer-Dobson circulations would weaken
- \rightarrow Tropical tropopause would warm, hence more stratospheric H_2O
- \rightarrow Air would be older, hence also more stratospheric H₂O (more CH₄ oxidation) and less CH₄
- \rightarrow More HO_x , hence more chemical ozone loss (except in middle stratosphere)
- \rightarrow Less diabatic descent in midlatitudes and poles, hence colder temperatures
- \rightarrow Less ozone transport to midlatitudes and poles, hence still colder temperatures
- → Stronger polar vortex, hence more of a barrier to ozone transport, and less polar ozone
- \rightarrow Increased PSC chemistry, hence more chemical ozone loss, and still less polar ozone
- ullet On long timescales, changes in atmospheric composition can change both $T_{\rm rad}$ and wave drag

- But since radiative and chemical effects are dissipative (stable), unforced natural variability arises solely from dynamics, through variability in wave drag
 - → Variability in tropospheric forcing of planetary waves
 - → Variability internal to the stratosphere (e.g. QBO)
- There are feedbacks, which tend to be positive:
 - \rightarrow More transport \Rightarrow more $O_3 \Rightarrow$ still more heating
 - \rightarrow Higher T \Rightarrow less heterogeneous chemistry \Rightarrow still more O_3
 - \rightarrow Weaker vortex \Rightarrow more transport \Rightarrow still more O_3
- So one can think of chemistry and radiation as amplifiers of the dynamical variability, but the variability originates in the dynamics (apart from volcanic eruptions)
- Link between diabatic and Brewer-Dobson circulations leads to positive O₃-T correlation in the lower stratosphere which is, in itself, purely coincidental
 - → Extent of coupling, and causality of statistical relationships, need to be assessed on a case-by-case basis

Long-term changes in ozone

- Chlorine loading over last few decades has led to Antarctic ozone hole
 - → Antarctic is always cold enough for PSC formation
 - → Indeed even before the chlorine build-up and chemical ozone loss, the Antarctic exhibited a transport-induced, purely dynamical, "ozone hole"
- No ozone hole in Arctic because vortex is too warm
 - → When the Arctic has a cold winter, then chemical ozone loss occurs
 - \rightarrow However low T from less downwelling \Rightarrow less O₃ transport
 - \rightarrow Low O₃ in recent cold Arctic winters comes half from chemical loss, half from reduced transport (WMO 2003)
- In Arctic, ozone levels controlled by meteorology more than by halogen loading (but effects are coupled)
 - \rightarrow One has to look at effects of halogen loading on ozone variability
 - \rightarrow Observed record is just one realization from an ensemble of possibilities

- Chemistry climate models reproduce development of Antarctic ozone hole, but don't simulate the Arctic record particularly well (whereas CTMs do, since driven by obs)
 - → But does this mean they are wrong?
- There have also been ozone decreases in midlatitudes
 - → Can be simulated by 2D models, but there are issues

We can only understand the ozone record (outside of Antarctica) in the context of the dynamical record

- There have been long-term changes in various stratospheric climate indicators over past decades
 - → Arctic wintertime vortex has gotten colder and stronger, and more persistent
 - → Stratopheric air has become "older"
 - → Stratospheric PWD has weakened
- These results are all qualitatively consistent with each other
- However they are quite sensitive to the months and time period considered
 - → Arctic exhibited warming (and more PWD) in early winter, and cooling (and less PWD) in late winter

- → Note that this cooling cannot be explained by ozone loss
- A long-term decrease in stratospheric PWD implies a long-term decrease in the winter-to-summer O_3
 - \rightarrow Over 1979-2000, estimate is about 20-30% of midlatitude total O₃ changes in January-March (WMO 2003)
 - → Is there a chemical amplification, as in the Arctic?

Climate change in the stratosphere

- WMGHGs will increase as halogens decrease
 - → Climate change and ozone recovery problem are coupled
 - → Furthermore ozone is a GHG, which interacts strongly with temperature especially in the upper stratosphere
- Increasing CO₂ will cool the stratosphere radiatively, but mainly the upper stratosphere
 - \rightarrow Increased IR emission exceeds increased IR absorption
 - \rightarrow Decreases O_3 loss rates, hence enhances O_3 recovery
- In the lower stratosphere, cooling could potentially decrease O_3 , but direct radiative effect is weak

- → The most significant effects would arise from dynamical feedbacks (meaning planetary wave drag)
- GCM predictions of greenhouse-gas-induced changes in NH stratospheric wave drag diverge widely
 - → Older results tended to suggest a reduction in NH stratospheric wave drag (Shindell et al. 1998 Nature)
 - \rightarrow Was argued by some to explain observed Arctic trends in 1990s (and to imply a future Arctic O₃ hole)
 - but not after 1998!
 - \rightarrow More recent studies tend to suggest an *increase* in stratospheric wave drag
- We probably cannot say anything with any confidence
 - → This aspect of GCM behaviour is highly sensitive
 - → In most cases, results are not statistically significant
 - → Certainly GCMs do not explain the past NH behaviour
 - \rightarrow Long-term records show no particular trend
- Natural variability may well be the dominant factor in the evolution of Arctic ozone over the next few decades
 - \rightarrow However the Antarctic ozone hole should recover by ~ 2050 (will take a decade or so to see turnover)