

# Chapter 13

## Interaction of Electrons with Electromagnetic Field

### Classical Theory;

Maxwell's equs. in Gaussian units in the vacuum

$$\nabla \cdot \mathbf{B}(\vec{r}, t) = 0 \quad (a)$$

$$\nabla \times \mathbf{E}(\vec{r}, t) + \frac{1}{c} \frac{\partial \mathbf{B}(\vec{r}, t)}{\partial t} = 0 \quad (b)$$

$$\nabla \cdot \mathbf{E}(\vec{r}, t) = 4\pi \rho(\vec{r}, t) \quad (c) \quad (1)$$

$$\nabla \times \mathbf{B}(\vec{r}, t) - \frac{1}{c} \frac{\partial \mathbf{E}(\vec{r}, t)}{\partial t} = \frac{4\pi}{c} \mathbf{J}(\vec{r}, t) \quad (d)$$

$\rho(\vec{r}, t)$ : charge density

$\mathbf{J}(\vec{r}, t)$ : current

: The sources of electromag. fields  
 $\mathbf{E}(\vec{r}, t)$ ,  $\mathbf{B}(\vec{r}, t)$

$$(1) \rightarrow \frac{\partial \rho(\vec{r}, t)}{\partial t} + \nabla \cdot \mathbf{J}(\vec{r}, t) = 0 \quad (2) \text{ Charge conservation}$$

In terms of scalar and vector pot.:

$$\mathbf{B}(\vec{r}, t) = \nabla \times \mathbf{A}(\vec{r}, t)$$

$$\mathbf{E}(\vec{r}, t) = -\frac{1}{c} \frac{\partial \mathbf{A}(\vec{r}, t)}{\partial t} - \nabla \varphi(\vec{r}, t) \quad (3)$$

In order to satisfy the first two of (1) .

$$\left. \begin{array}{l} \text{Remark:} \\ \nabla \cdot \nabla \times \mathbf{F} = 0 \\ \nabla \times \nabla \varphi = 0 \end{array} \right\}$$

The fields  $E$  and  $B$  do not determine  $\mathcal{A}$  and  $\mathcal{A}$  uniquely.

New pots. given by

$$\begin{aligned} A'(\vec{r}, t) &= A(\vec{r}, t) - \nabla f(\vec{r}, t) \\ \varphi'(\vec{r}, t) &= \varphi(\vec{r}, t) + \frac{1}{c} \frac{\partial f(\vec{r}, t)}{\partial t} \end{aligned} \quad (4)$$

yield the same  $E$  and  $B$ .

$$\begin{aligned} \begin{cases} A \\ \varphi \end{cases} &\xrightarrow{\text{Gauge tr.}} \begin{cases} A' \\ \varphi' \end{cases} \\ \begin{cases} E \\ B \end{cases} &\xrightarrow[\text{invariance}]{} \begin{cases} E \\ B \end{cases} \end{aligned} \quad (5)$$

→ we can choose  $f(\vec{r}, t)$  in the most convenient way.

$$(1-c)(1-d)(3) \rightarrow \begin{cases} -\nabla^2 \varphi(\vec{r}, t) - \frac{1}{c} \frac{\partial}{\partial t} (\nabla \cdot A) = 4\pi \rho(\vec{r}, t) & (a) \\ \nabla \times (\nabla \times A) + \frac{1}{c^2} \frac{\partial^2 A}{\partial t^2} + \frac{1}{c} \frac{\partial}{\partial t} \nabla \varphi = \frac{4\pi}{c} \mathcal{J}(\vec{r}, t) & (b) \end{cases} \quad (6)$$

The source-dip. equs.

Since  $\nabla \times (\nabla \times A) = -\nabla^2 A + \nabla(\nabla \cdot A)$  (only valid in Cartesian coord.)

$$(6-b) \rightarrow -\nabla^2 A + \frac{1}{c^2} \frac{\partial^2 A}{\partial t^2} + \nabla \left( \nabla \cdot A + \frac{1}{c} \frac{\partial \varphi}{\partial t} \right) = \frac{4\pi}{c} \mathcal{J} \quad (7)$$

If  $\frac{\partial \rho}{\partial t} = 0$  (static charge dist.), it is convenient.

to choose the gauge such that

$$\nabla \cdot A(\vec{r}, t) = 0 \quad (8) \quad \text{Coulomb gauge}$$

(with a suitable choice of  $f(\vec{r}, t)$  (eqn. (4))),

In that case: (3)  $\rightarrow -\nabla^2 \varphi(\vec{r}) = 4\pi\rho(r)$  (9)  
 (1-c)  
 (8)  $t$ -indep.

$$(7)(9) \rightarrow -\nabla^2 A(\vec{r}, t) + \frac{1}{c^2} \frac{\partial^2 A(\vec{r}, t)}{\partial t^2} = \frac{4\pi}{c} J(\vec{r}, t) \quad (10)$$

(since  $\frac{\partial \varphi}{\partial t} = 0$ )

When  $\frac{\partial \rho}{\partial t} \neq 0$ , it is convenient to choose Lorentz gauge;

$$\nabla \cdot A(\vec{r}, t) + \frac{1}{c} \frac{\partial \varphi(\vec{r}, t)}{\partial t} = 0 \quad (11)$$

This leaves the equ. (6-b) unaltered,

$$(11) (7) \rightarrow -\nabla^2 A(\vec{r}, t) + \frac{1}{c^2} \frac{\partial^2 A(\vec{r}, t)}{\partial t^2} = \frac{4\pi}{c} J(\vec{r}, t) \quad (12)$$

$$(11) (6-a) \rightarrow -\nabla^2 \varphi(\vec{r}, t) + \frac{1}{c^2} \frac{\partial^2 \varphi(\vec{r}, t)}{\partial t^2} = 4\pi\rho(\vec{r}, t) \quad (13)$$

Now, the int. of a point electron of mass  $\mu$  and charge  $(-e)$  with an elmag. field is the classical Lorentz force equ.;

$$\mu \frac{d^2 \vec{r}}{dt^2} = -e \left[ \vec{E}(\vec{r}, t) + \frac{\vec{v}}{c} \times \vec{B}(\vec{r}, t) \right] \quad (14)$$

Remember:  $\frac{d}{dt} \left( \frac{\partial L}{\partial \dot{q}_i} \right) - \frac{\partial L}{\partial q_i} = 0 \quad (15)$

$$p_i = \frac{\partial L(q_i, \dot{q}_i, t)}{\partial \dot{q}_i} \quad (16)$$

$$L = T - U = \frac{1}{2} m v^2 - \left( q\phi - \frac{q}{c} \mathbf{A} \cdot \mathbf{v} \right) \quad (17) \quad (\text{Goldstein})$$

$\uparrow$   
 generalized  
 pot

$$(16)/(17) \rightarrow p_i = m \dot{x}_i + \frac{q}{c} A_i \quad (18)$$

$$\rightarrow m \dot{x}_i = p_i - \frac{q}{c} A_i$$

$$H = \frac{(m \dot{x}_i)^2}{2m} + q\phi = \frac{1}{2m} \left( \bar{p} - \frac{q}{c} \bar{A} \right)^2 + q\phi \quad (19)$$

$$\text{Remember: } \dot{q}_i = \frac{\partial H}{\partial p_i}, \quad -\dot{p}_i = \frac{\partial H}{\partial q_i}, \quad \frac{\partial L}{\partial t} = \frac{\partial H}{\partial t} \quad (20)$$

Using (20) and (19), the equ. (16) can be proved for each component.

So in the presence of the vector and the scalar fields  $A(\vec{r}, t)$  and  $\varphi(\vec{r}, t)$ ; the Hamiltonian for an electron having charge  $(-e)$  and mass  $\mu$  is given by:

$$H = \frac{1}{2\mu} \left[ P + \frac{e}{c} A(\vec{r}, t) \right]^2 - e\varphi(\vec{r}) \quad (21)$$

↑ we are dealing with static scalar pots.

The corresponding Schrödinger equation:

$$\text{with, } P \longrightarrow P - \frac{q}{c} A(\vec{r}, t) \quad (22)$$

$$\text{and } V = q\varphi(\vec{r}) \quad (23)$$

$$-\frac{1}{2\mu} \left( \frac{\hbar}{i} \nabla - \frac{-e}{c} A \right)^2 \psi(\vec{r}, t) + (-e)\varphi(\vec{r}) \psi(\vec{r}, t) = E \psi(\vec{r}, t) \quad (24)$$

Now;

$$\begin{aligned} & \frac{1}{2\mu} \left( \frac{\hbar}{i} \nabla + \frac{e}{c} A \right) \cdot \left( \frac{\hbar}{i} \nabla \psi + \frac{e}{c} A \psi \right) = \\ & = -\frac{\hbar^2}{2\mu} \nabla^2 \psi - \frac{2ie\hbar}{\mu c} A \cdot \nabla \psi - \frac{2ie\hbar}{2\mu c} (\nabla \cdot A) \psi + \frac{e^2}{2\mu c^2} A^2 \psi \\ & = -\frac{\hbar^2}{2\mu} \nabla^2 \psi - \frac{ie\hbar}{\mu c} A \cdot \nabla \psi + \frac{e^2}{2\mu c^2} A^2 \psi \quad (25) \end{aligned}$$

where we have used;

$$\begin{aligned} \nabla \cdot (A \psi) &= (\nabla \cdot A) \psi + A \cdot \nabla \psi \\ \text{and } \nabla \cdot A &= 0 \end{aligned} \quad (26)$$

For a const. uniform mag. field,  $\vec{B}$ , we may take

$$A = -\frac{1}{2} r \times B \quad (27)$$

Note that this choice is not unique, because

$$A = -\frac{1}{2} r \times B + \nabla g \quad (28) \quad g: \text{arbitrary func.}$$

gives the same  $\vec{B}$ .

Now 
$$A = -\frac{1}{2} (yB_z - zB_y, zB_x - xB_z, xB_y - yB_x)$$

$$\nabla \times A = \left( \frac{1}{2} B_x + \frac{1}{2} B_x, B_y, B_z \right) = \vec{B} \quad (29)$$

The second term in (25)

$$\begin{aligned} + \frac{ze\hbar}{2\mu c} r \times B \cdot \nabla \psi &= -\frac{ze\hbar}{2\mu c} B \cdot r \times \nabla \psi = \frac{e}{2\mu c} B \cdot r \times \left( \frac{\hbar}{i} \nabla \psi \right) \\ &= \frac{e}{2\mu c} B \cdot L \psi \end{aligned}$$

where we have used

$$(A \times B) \cdot C = B \cdot (A \times C)$$

and the third term, taking  $\vec{B} = B \hat{z}$

$$\frac{e^2}{8\mu c^2} (r \times B)^2 \psi = \frac{e^2}{8\mu c^2} [r^2 B^2 - (r \cdot B)^2] \psi = \frac{e^2 B^2}{8\mu c^2} (x^2 + y^2) \psi \quad (31)$$

similar to two-dim.  
Harmonic osc. pot.

When we have used

$$(A \times B) \cdot (A \times B) = A^2 B^2 - (A \cdot B)^2$$

Let us compare the magnitude of the two terms;

$$\begin{aligned} \frac{\langle \frac{e^2 B^2}{8\mu c^2} (x^2 + y^2) \rangle}{\langle \frac{e}{2\mu c} B \cdot L \rangle} &= \frac{\frac{e^2 B^2}{8\mu c^2} \langle (x^2 + y^2) \rangle}{\frac{e}{2\mu c} B \cdot \langle L \rangle} = \frac{\frac{e^2 B^2}{8\mu c^2} \langle (x^2 + y^2) \rangle}{\frac{e}{2\mu c} B \langle L \rangle} \\ &\approx \frac{\frac{e^2 B^2}{8\mu c^2} a_0^2}{\frac{e}{2\mu c} B (\hbar)} \approx \frac{1}{4} \frac{e^2}{\hbar c} \frac{B}{e/a_0} \approx \frac{1}{548} \frac{B}{e/a_0} \\ &= \frac{B}{548 (4.8 \times 10^{-10}) / (0.5 \times 10^{-8})^2} = \frac{B}{9 \times 10^9 \text{ gauss}} \quad (32) \end{aligned}$$

Then for atomic systems, with  $B \leq 10^6$  gauss (available in the laboratory) the quadratic term is negligible.

Also

$$\begin{aligned} \frac{\text{Linear term}}{\text{Coulomb pot. energy}} &= \frac{\langle \frac{e}{2\mu c} B \cdot L \rangle}{\frac{e^2}{r}} \approx \frac{\frac{e}{2\mu c} \hbar B}{\frac{e^2}{r}} \approx \frac{1}{2} \frac{\hbar/\mu c}{e/a_0} B \\ &\approx \frac{1}{274} \frac{B}{e/a_0} = \frac{B}{5 \times 10^9 \text{ gauss}} \quad (33) \end{aligned}$$

So that the linear term will only slightly perturb the atomic energy levels.

The quadratic is important in two conditions:

{ i -  $B \rightarrow$  large (on the surface of neutron stars  $B \sim 10^{12}$  gauss)  
ii - macroscopic motion of an electron (out of the atom)  
 $\rightarrow (x^2 + y^2) \approx$  large) in an external field.

The linear term is

$$H_1 = \frac{e}{2\mu c} B L_z = \omega_L L_z \quad (34)$$

where  $\omega_L = \frac{eB}{2\mu c}$  Larmor frequency

If the energy eigenstates are simultaneous eigenstates of  $L^2$  and  $L_z$ ;

$$\text{i.e. } H \psi = E \psi, \quad L^2 \psi = l(l+1)\hbar^2 \psi, \quad L_z \psi = m\hbar \psi \quad (35)$$

$$\rightarrow H_1 \psi = \omega_L (m\hbar) \psi$$

$$\text{or } H_1 U_{nlm}(\vec{r}) = \hbar \omega_L m U_{nlm}(\vec{r}) \quad (36)$$

$$-l \leq m \leq l$$



$$E = -\frac{1}{2} \mu c^2 \left( \frac{Z\alpha}{n^2} \right)^2 + \hbar \omega_L m \quad (37)$$

→  $(2l+1)$ -fold degeneracy has been disappeared -

The size of splitting:

$$\begin{aligned} \hbar \omega_L &= \frac{eB\hbar}{2\mu c} = \frac{e\hbar}{2\mu c} \left( \frac{B}{e/a_0^2} \right) \frac{e}{a_0^2} = \frac{e^2\hbar}{2\mu c} \left( \frac{\mu c a_0}{\hbar} \right)^2 \left( \frac{B}{e/a_0^2} \right) \\ &= \left( \frac{1}{2} \alpha^2 \mu c^2 \right) \frac{B}{e/a_0^2} = \frac{B}{2.4 \times 10^9} (13.6) \text{ eV} \quad (38) \end{aligned}$$

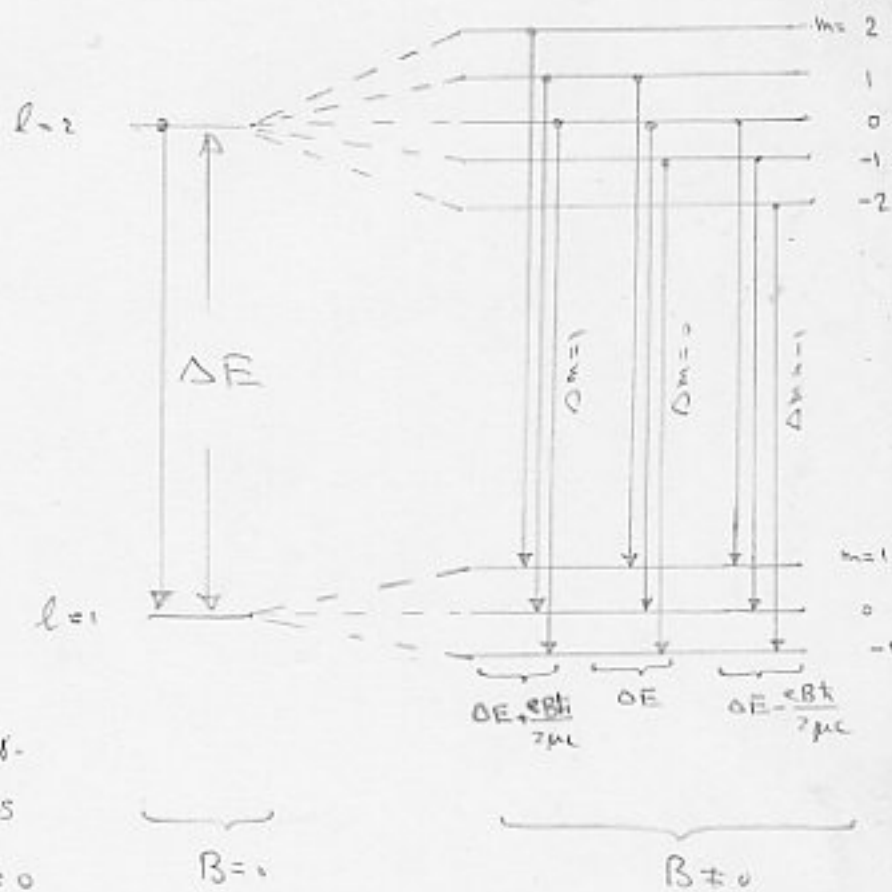
So, the energy levels split to  $2l+1$  distinct levels (in the presence of  $B \neq 0$ ); This is the normal Zeeman effect.

The anomalous Zeeman effect takes into account the spin of the electron.

Taking into account the selection rule (to be discussed later)

$$\Delta m = 0, \pm 1$$

A single tr. line → 3-tr. lines  
 $B = 0$                        $B \neq 0$



Another approximation;  
under conditions where

$\left\{ \begin{array}{l} \text{quadratic term } (\sim B^2) \text{ is not negligible} \\ \text{Coulomb term } (-e\varphi(\vec{r})) \text{ is negligible} \end{array} \right.$

The Schrödinger equ.:

$$-\frac{\hbar^2}{2\mu} \nabla^2 \psi + \frac{eB}{2\mu c} L_z \psi + \frac{e^2 B^2}{8\mu c^2} (x^2 + y^2) \psi = E \psi \quad (39)$$

The presence of the potential  $\sim (x^2 + y^2)$  suggests the use of cylindrical coordinates for the separation of the variables;

$$\begin{cases} x = \rho \cos \varphi \\ y = \rho \sin \varphi \end{cases} \rightarrow \begin{cases} d\rho = \cos \varphi dx - \sin \varphi dy \\ d\varphi = \frac{1}{\rho} (-\sin \varphi dx + \cos \varphi dy) \end{cases} \quad (40)$$

$$\frac{\partial}{\partial x} = \frac{\partial}{\partial \rho} \frac{\partial \rho}{\partial x} + \frac{\partial}{\partial \varphi} \frac{\partial \varphi}{\partial x} = \cos \varphi \frac{\partial}{\partial \rho} - \frac{\sin \varphi}{\rho} \frac{\partial}{\partial \varphi}$$

$$\frac{\partial}{\partial y} = \frac{\partial}{\partial \rho} \frac{\partial \rho}{\partial y} + \frac{\partial}{\partial \varphi} \frac{\partial \varphi}{\partial y} = \sin \varphi \frac{\partial}{\partial \rho} + \frac{\cos \varphi}{\rho} \frac{\partial}{\partial \varphi} \quad (41)$$

$$\rightarrow \nabla^2 = \frac{\partial^2}{\partial z^2} + \frac{\partial^2}{\partial \rho^2} + \frac{1}{\rho} \frac{\partial}{\partial \rho} + \frac{1}{\rho^2} \frac{\partial^2}{\partial \varphi^2} \quad (42)$$

If we now write  $\psi(\vec{r}) = u_m(\rho) e^{im\varphi} e^{ikz}$

$$\rightarrow \frac{d^2 u}{d\rho^2} + \frac{1}{\rho} \frac{du}{d\rho} - \frac{m^2}{\rho^2} u - \frac{e^2 B^2}{4\hbar^2 c^2} \rho^2 u + \left( \frac{2kE}{\hbar^2} - \frac{eB\hbar m}{\hbar^2 c} - k^2 \right) u = 0$$

No let  $u = \sqrt{\frac{eB}{2\hbar c}} \rho$  (43)

$$\rightarrow \frac{d^2 u}{d\rho^2} + \frac{1}{\rho} \frac{du}{d\rho} - \frac{m^2}{\rho^2} u - \rho^2 u + \lambda u = 0 \quad (44)$$

where  $\lambda = \frac{4\mu c}{eB\hbar} \left( E - \frac{\hbar^2 k^2}{2\mu} \right) - 2m$  (45)

i) The behavior of  $u(\rho)$  at infinity;

$$(44) \rightarrow \frac{d^2 u}{d\rho^2} - \rho^2 u \approx 0 \quad (46)$$

$\rho \rightarrow \infty$

$$\rightarrow u(\rho) \sim e^{-\rho^2/2} \quad (47)$$

ii) The behavior of  $u(\rho)$  near  $\rho \approx 0$

$$(44) \rightarrow \frac{d^2 u}{d\rho^2} + \frac{1}{\rho} \frac{du}{d\rho} - \frac{m^2}{\rho^2} u \approx 0 \quad (48)$$

$$\rightarrow u(\rho) \sim \rho^{|m|} \quad (49) \quad \left( \rho^{|m|} \text{ is singular near } \rho \rightarrow \infty \right)$$

$$\rightarrow u(\rho) = \rho^{|m|} e^{-\rho^2/2} G(\rho) \quad (50)$$

$$(44) (50) \rightarrow \frac{d^2 G}{dx^2} + \left( \frac{2|m|+1}{x} - 2x \right) \frac{dG}{dx} + (\lambda - 2 - 2|m|) G = 0 \quad (51)$$

Now let  $y \equiv x^2$  (52)

$$\rightarrow \frac{d^2 G}{dy^2} + \left( \frac{|m|+1}{y} - 1 \right) \frac{dG}{dy} + \frac{\lambda - 2 - 2|m|}{4y} G = 0 \quad (53)$$

This is of the form of equ. for  $H$  discussed in chap. 12, (equ. 12-35);

Comparing (53) with (12-35, p229); we must have;

$$\frac{\lambda}{4} - \frac{1+|m|}{2} = n_r \quad (54)$$

where  $n_r = 0, 1, 2, \dots$  (55)

$$\stackrel{(45)}{(54)} \rightarrow E = -\frac{\hbar^2 k^2}{2\mu} = \frac{eB\hbar}{2\mu c} (2n_r + 1 + |m| + m) \quad (56)$$

and  $G(y) = L_{n_r}^{|m|}(y)$  (57)

Classical limit;

Given equ. (21) for  $H$ , without the scalar pot. term;

we have  $\mu \dot{X} = P - \left( \frac{e}{c} A \right) \rightarrow V = \frac{P + \frac{e}{c} A}{\mu}$  (58)

and with  $A = -\frac{1}{2} r \times B$

$$\mu(r \times v) = r \times p + \frac{e}{c} r \times \left(-\frac{1}{2} r \times B\right) = \bar{L} - \frac{e}{2c} [\bar{r}(r \cdot B) - r^2 \bar{B}] \quad (59)$$

where we have used the identity;

$$a \times (b \times c) = b(a \cdot c) - c(a \cdot b) \quad (60)$$

The z-component of (59)

$$[\mu(r \times v)]_z = L_z + \frac{e}{2c} B (x^2 + y^2) \quad (61)$$

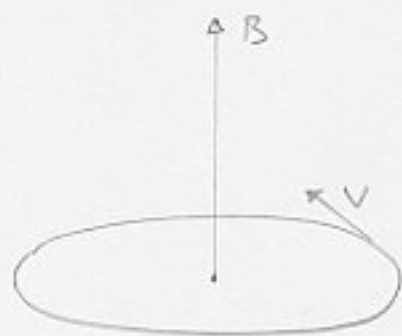
that is;  $\mu g v = L_z + \frac{eB}{2c} S^2 \quad (62)$

The Lorentz force on the electron;

$$\bar{F} = -\frac{e}{c} \bar{v} \times \bar{B} + 0 \quad (63)$$

$$\rightarrow \frac{\mu v^2}{g} = \frac{e v B}{c} \quad (64)$$

for circular motion



$$\begin{cases} (62) \rightarrow \frac{1}{2} \mu v^2 = \frac{eB}{\mu c} L_z & (65) \\ (64) \rightarrow S = \left[ \frac{2c}{eB} L_z \right]^{1/2} & (66) \end{cases}$$

Now return to the expression for  $E$ , (56),

Since  $\hbar =$  very small;

$E$  can be of macroscopic size for reasonable  $B$ , if

$$(2n_r + 1 + |m| + m) = \text{very large} \quad (66')$$

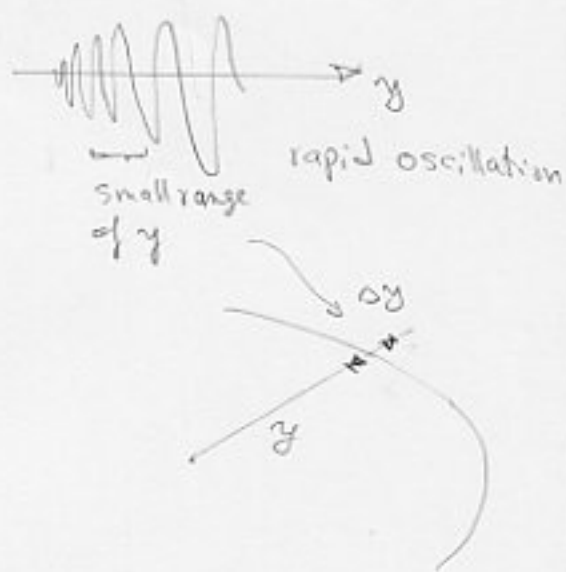
We have two cases:

a)  $m < 0 \xrightarrow{(66')} n_r = \text{very large}$

$n_r$ : degree of the polynomial  $L_{n_r}^{2|m|}(y)$  (number of zeros)

Now if  $n_r = \text{large} \rightarrow$

$\rightarrow$  the func. cannot be large  
for some small range of  
 $y$  where the classical orbit  
would be located



b)  $m > 0$  the coeff.  $= (2n_r + 1 + 2m)$

This can be large with  $n_r = \text{small}$ , provided  $m = \text{large}$

$$\rightarrow E - \frac{\hbar^2 k^2}{2\mu} \approx \frac{cB}{\mu c} \hbar m \quad (67)$$

This is in agreement with the classical result (65)

Note that  $L_z = m\hbar$  (68)

is positive, as expected.

We can also show that the radius of the orbit, as determined by the peaking of the radial probability dist., corresponds to the classical value.

Let us take  $n_r = 0 \rightarrow \int_{n_r=0}^{|m|} \psi(r) = \text{const.}$

(50)  $\rightarrow P(x) = |U(m)|^2 \sim x^{2|m|} e^{-x^2}$  (69)

$$\frac{dP}{dx} = (2|m| x^{2|m|-1} - 2x^{2|m|+1}) e^{-x^2} = 0 \quad (70)$$

has a max at

$$x = \sqrt{|m|} \quad (71)$$

$$r = \left( \frac{2c}{eB} \hbar m \right)^{1/2} \quad (72)$$

which is in agreement with (66)

This prob. is a beautiful illustration of the correspondence principle.

There are several interesting quantum mechanical effects connected with the interaction with a mag. field that we now turn to.

The Schrödinger equ. (24) appears to violate the principle of gauge invariance, since it is  $A(\vec{r}, t)$  that appears in the equ., and under the tr.;

$$A \longrightarrow A + \nabla f(\vec{r}, t) \quad (73)$$

the Hamiltonian is changed acc. to;

$$\frac{1}{2\mu} \left( \frac{\hbar}{i} \nabla + e A \right)^2 \longrightarrow \frac{1}{2\mu} \left( \frac{\hbar}{i} \nabla + e A + e \nabla f \right)^2 \quad (74)$$

It is possible to save gauge invariance by using the fact that a change of the wave func. by a phase factor, which may depend on  $\vec{r}$ , has no physical consequences.

Thus if we require that (73) must be accompanied by the tr.

$$\psi(\vec{r}, t) \longrightarrow e^{iA(\vec{r}, t)} \psi(\vec{r}, t) \quad (75)$$

The first term of the left hand side of (24) becomes;



$$\begin{aligned}
& \frac{1}{2\mu} \left( \frac{\hbar}{i} \nabla + \frac{e}{c} A + \frac{e}{c} \nabla f \right) \cdot \left( \frac{\hbar}{i} \nabla + \frac{e}{c} A + \frac{e}{c} \nabla f \right) e^{i\Lambda} \Psi = \\
& = \frac{1}{2\mu} \left( \frac{\hbar}{i} \nabla + \frac{e}{c} A + \frac{e}{c} \nabla f \right) \cdot \left[ e^{i\Lambda} \left( \frac{\hbar}{i} \nabla \Psi + \frac{e}{c} A \Psi + \frac{e}{c} \nabla f \Psi \right. \right. \\
& \quad \left. \left. + \hbar \nabla \Lambda \Psi \right) \right] \\
& = \frac{1}{2\mu} e^{i\Lambda} \left( \frac{\hbar}{i} \nabla + \frac{e}{c} A + \frac{e}{c} \nabla f + \hbar \nabla \Lambda \right)^2 \Psi \quad (76)
\end{aligned}$$

Thus with the choice;  $\Lambda = -\frac{e}{\hbar c} f$  (77)

That is, with the tr. law,

$$\Psi(\vec{r}, t) \longrightarrow e^{-\frac{i e}{\hbar c} f(\vec{r}, t)} \Psi(\vec{r}, t) \quad (78)$$

gauge invariance is restored.

In the field-free region,  $B=0$

$$\text{i.e. } \nabla \times A = 0 \quad (79)$$

$$\rightarrow A = \nabla f \quad (80)$$

In this case we may describe the motion of an electron in two ways:

$$\text{i)} \quad \left[ \frac{1}{2\mu} \left( \frac{\hbar}{i} \nabla \right)^2 + V(\vec{r}) \right] \Psi = E \Psi \quad (81)$$

$$\text{ii)} \quad \frac{1}{2\mu} \left( \frac{\hbar}{i} \nabla + \frac{e}{c} A \right)^2 \Psi' + V(\vec{r}) \Psi' = E \Psi' \quad (82)$$

where  $A$  is given by (80)

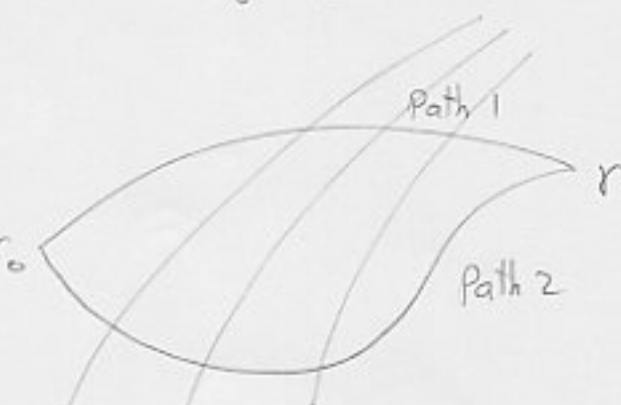
and take  $\psi' = e^{-\frac{i e}{\hbar c} \int \psi}$  (83)

(80)  $\rightarrow \int_{r_0}^r \vec{dr}' \cdot \vec{A}(\vec{r}', t)$  (84)

↑  
arbitrary point  
the origin or infinity

The integral only makes sense if  $B = 0$  (field-free region)

i.e.

$$\int_1 \vec{dr}' \cdot \vec{A}(\vec{r}', t) - \int_2 \vec{dr}' \cdot \vec{A}(\vec{r}', t) = \oint \vec{dr}' \cdot \vec{A}(\vec{r}', t)$$


$$= \int_S \nabla' \times \vec{A}(\vec{r}', t) \cdot d\vec{s} = \int_S B \cdot d\vec{s} = \Phi$$
 (85)

where we have used Stokes' theorem.

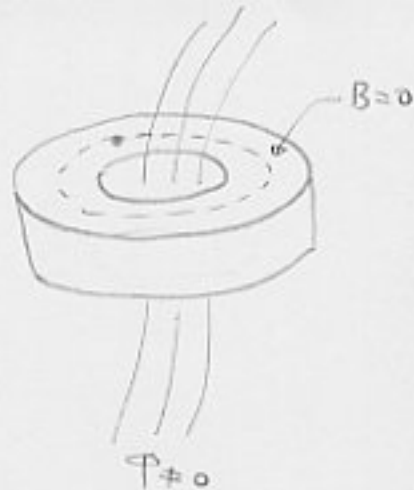
$$\rightarrow \int_1 \vec{dr}' \cdot \vec{A}(\vec{r}', t) \neq \int_2 \vec{dr}' \cdot \vec{A}(\vec{r}', t) \quad \text{if } B \neq 0 \text{ in general}$$

Thus if  $\Phi = 0 \rightarrow$  the phase factor will be indep. of the path

Such independence is required if we insist that the wave func. be single-valued.

An interesting consequence:

An electron moves in a field-free region surrounds a hole containing flux  $\Phi$ .



Upon completing a circuit the electron acquires an additional phase factor  $e^{ie\Phi/\hbar c}$ .

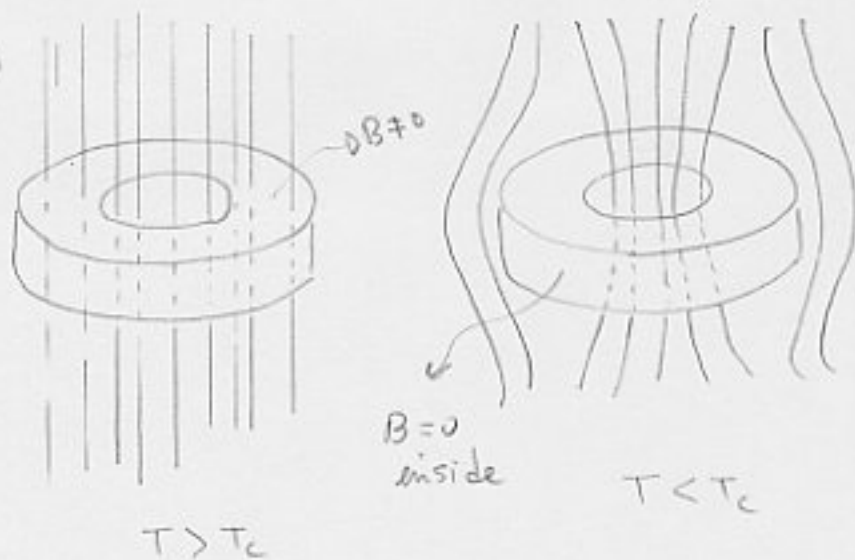
Since the electron wave func. is single-valued;

$$\rightarrow e^{ie\Phi/\hbar c} = 1 \quad \rightarrow \Phi = \frac{2\pi\hbar c}{e} n \quad (86)$$

$$n = 0, \pm 1, \pm 2, \dots$$

An ingenious measurement of the flux shows that (86) holds with the modification that

$$\Phi = \frac{2\pi\hbar c}{2e} n \quad (87)$$



$\rightarrow$  The electrons move in pair, (correlated states).

A superconductor at  $T > T_c$  behaves like any other metal. When  $T < T_c$  it becomes superconductor expelling the mag. field.

## Another Example: Aharonov and Bohm

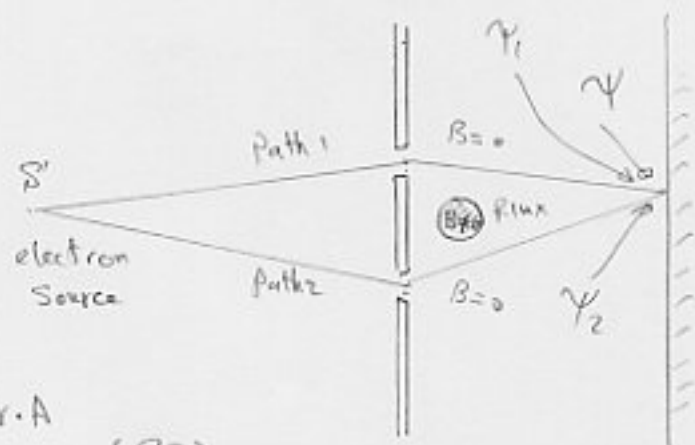
Interference experiment of the electrons;

In the absence of solenoid;

$$\psi = \psi_1 + \psi_2$$

In the presence of solenoid;

$$\begin{aligned}\psi &= \psi_1 e^{\frac{ie}{\hbar c} \int_1 dr \cdot A} + \psi_2 e^{\frac{ie}{\hbar c} \int_2 dr \cdot A} \\ &= (\psi_1 e^{\frac{ie\Phi}{\hbar c}} + \psi_2) e^{\frac{ie}{\hbar c} \int_2 dr \cdot A} \quad (88)\end{aligned}$$



→ Thus the flux causes a relative change in phase between  $\psi_1$  and  $\psi_2$  (i.e.  $e^{\frac{ie\Phi}{\hbar c}}$ )

→ This will change the interference pattern.