

Chapter 11

The Radial Equation

(10-82, P198) \rightarrow

$$\left(\frac{d^2}{dr^2} + \frac{2}{r} \frac{d}{dr} \right) R_{nlm}(r) - \frac{2\mu}{\hbar^2} \left[V(r) + \frac{l(l+1)\hbar^2}{2\mu r^2} \right] R_{nlm}(r) + \frac{2\mu E}{\hbar^2} R_{nlm}(r) = 0 \quad (1)$$

We will examine the sds. to this equ. for a variety of potentials, restricted by the cond. that:

$$V(r) \rightarrow 0 \quad \text{as } r \rightarrow \infty \quad \text{faster than } \frac{1}{r} \quad \left(\begin{array}{l} \text{except for the} \\ \text{Coulomb pot.} \end{array} \right)$$

(see eq. 12)

Also we will assume $\lim_{r \rightarrow 0} r^2 V(r) = 0$ (i.e. not singular as $\frac{1}{r^2}$ at the origin)

(2) (see eq. 8)

For convenience; $U_{nlm}(r) = r R_{nlm}(r)$ (3)

Since $\left(\frac{d^2}{dr^2} + \frac{2}{r} \frac{d}{dr} \right) \frac{U_{nlm}(r)}{r} = \frac{1}{r} \frac{d^2}{dr^2} U_{nlm}(r)$ (4)

$$\rightarrow \frac{d^2 U_{nlm}(r)}{dr^2} + \frac{2\mu}{\hbar^2} \left[E - V(r) - \frac{l(l+1)\hbar^2}{2\mu r^2} \right] U_{nlm}(r) = 0 \quad (5)$$

This looks like a one-dim. equ. except that:

a) $V(r) \rightarrow V(r) + \underbrace{\frac{l(l+1)\hbar^2}{2\mu r^2}}_{\text{Repulsive centrifugal barrier}}$ (6)

b) the def. of $U_{\text{nem}}(r)$ and the finiteness of the wave func. at the origin require that:

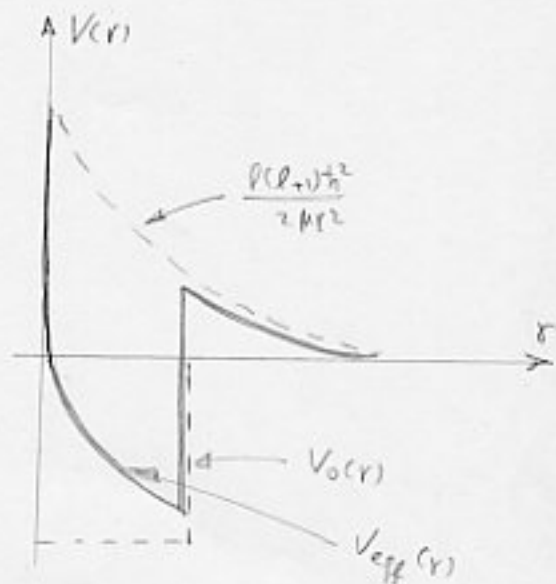
$$U_{\text{nem}}(0) = 0 \quad (7)$$

which makes it more like the one-dim. prob. for which

$V = +\infty$ in the left-hand region.

Consider the radial eq. near the origin. Taking into account the eq. (2);

$$\rightarrow \frac{d^2 U}{dr^2} - \frac{l(l+1)}{r^2} U \approx 0 \quad (8)$$



If we assume $U(r) \sim r^S$ (9)

$$(8) \rightarrow S(S-1)r^{S-2} - \frac{l(l+1)}{r^2} r^S \approx 0$$

$$\rightarrow S(S-1) - l(l+1) = 0 \quad (10)$$

$$\rightarrow \begin{cases} S = l+1 \\ S = -l \end{cases}$$

$$\rightarrow \begin{cases} U(r) \sim r^{l+1} & \text{regular sol. } (U(0) = 0) \\ U(r) \sim r^{-l} & \text{irregular sol. } (U(0) \neq 0) \end{cases}$$

(11)

For $r \rightarrow \infty$

$$\frac{d^2 u}{dr^2} + \frac{2ME}{\hbar^2} u \approx 0 \quad (12)$$

Since $\int |Y(r)|^2 dr^3 = 1$

$$\rightarrow \int_0^\infty r^2 dr \int d\Omega |R_{n\ell m}(r) Y_{\ell}^m(\theta, \phi)|^2 = \int_0^\infty r^2 dr |R_{n\ell m}(r)|^2 = 1$$

$$\rightarrow \int_0^\infty dr |u_{n\ell m}(r)|^2 = 1 \quad (13)$$

so that $\rightarrow u(r) \rightarrow 0$ as $r \rightarrow \infty$

a) If $E < 0$

$$\frac{2ME}{\hbar^2} = -\alpha^2 \quad (14)$$

The asymptotic sol. $u(r) \sim e^{-\alpha r}$ ($r \rightarrow \infty$) (15)

b) If $E > 0$, we have sol. that are normalizable in a

box. With $\frac{2ME}{\hbar^2} = k^2$ (16)

sol. $\rightarrow a e^{ikr} + b e^{-ikr}$ (17)

A - The Free Particle

$$V(r) = 0$$

$$\left[\frac{d^2}{dr^2} + \frac{2}{r} \frac{d}{dr} - \frac{l(l+1)}{r^2} \right] R(r) + k^2 R(r) = 0 \quad (18)$$

where $k^2 = \frac{2mE}{\hbar^2}$

$\rho \equiv kr$ change of variable

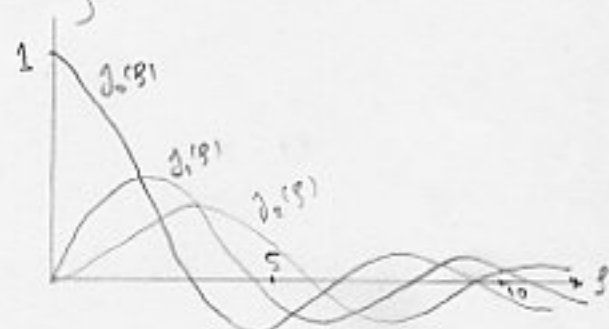
$$\rightarrow \frac{d^2 R}{d\rho^2} + \frac{2}{\rho} \frac{dR}{d\rho} - \frac{l(l+1)}{\rho^2} R + R = 0 \quad (19)$$

sols. \rightarrow spherical Bessel funcs.

$$\begin{cases} j_l(\rho) = (-\rho)^l \left(\frac{1}{\rho} \frac{d}{d\rho} \right)^l \left(\frac{\sin \rho}{\rho} \right) & \text{(regular)} \\ n_l(\rho) = -(-\rho)^l \left(\frac{1}{\rho} \frac{d}{d\rho} \right)^l \left(\frac{\cos \rho}{\rho} \right) & \text{(irregular)} \end{cases} \quad (20)$$

$$\begin{cases} j_0(\rho) = \frac{\sin \rho}{\rho} \\ n_0(\rho) = -\frac{\cos \rho}{\rho} \end{cases} \quad \begin{cases} j_1(\rho) = \frac{\sin \rho}{\rho^2} - \frac{\cos \rho}{\rho} \\ n_1(\rho) = -\frac{\cos \rho}{\rho^2} - \frac{\sin \rho}{\rho} \end{cases}$$

$$\begin{cases} j_2(\rho) = \left(\frac{3}{\rho^3} - \frac{1}{\rho} \right) \sin \rho - \frac{3}{\rho^2} \cos \rho \\ n_2(\rho) = -\left(\frac{3}{\rho^3} - \frac{1}{\rho} \right) \cos \rho - \frac{3}{\rho^2} \sin \rho \end{cases} \quad (21)$$



The combinations that will be of interest for large ρ are the spherical Hankel funcs.



$$h_e^{(1)} = J_e(s) + i\eta_e(s)$$

$$h_e^{(2)} = J_e(s) - i\eta_e(s) = (h_e^{(1)}(s))^* \quad (22)$$

$$h_0^{(1)}(s) = \frac{e^{is}}{is}, \quad h_1^{(1)}(s) = -\frac{e^{is}}{s} \left(1 + \frac{i}{s}\right), \quad h_2^{(1)} = \frac{ie^{is}}{s} \left(1 + \frac{3i}{s} - \frac{3}{s^2}\right) \quad (23)$$

a) The behavior near the origin,

$$s \rightarrow 0 \quad s \ll \ell$$

$$J_e(s) \approx \frac{s^\ell}{1 \cdot 3 \cdot 5 \cdots (2\ell+1)} = \frac{s^\ell}{(2\ell+1)!!}, \quad \eta_e \approx \frac{-(2\ell-1)!!}{s^{\ell+1}} \quad (24)$$

b) For large s -asymptotic limit;

$$s \rightarrow \text{large} \quad s \gg \ell$$

$$J_e(s) \approx \frac{1}{s} \Sigma \left(s - \frac{\ell n}{2}\right) = \frac{1}{s} \cos \left(s - \frac{(\ell+1)n}{2}\right)$$

$$\eta_e(s) = -\frac{1}{s} \cos \left(s - \frac{\ell n}{2}\right) = \frac{1}{s} \Sigma \left(s - \frac{(\ell+1)n}{2}\right) \quad (25)$$

so that;
$$h_e^{(1)}(s) \approx -\frac{i}{s} e^{i(s - \ell n/2)} = \frac{1}{s} e^{i(s - (\ell+1)n/2)}$$

The sol., regular at the origin:

$$R_\ell(r) = j_\ell(kr) \quad (26)$$

$$(25) \rightarrow \text{Its asymptotic form; } R_\ell(kr) \approx \frac{1}{2ikr} \left[e^{-i(kr - \ell\pi/2)} - e^{i(kr - \ell\pi/2)} \right]$$

$$\left\{ \begin{array}{l} \text{Remark: } j_0(0) = 1, j_1(0) = 0 \dots \\ \eta_0(0) = \infty, \eta_1(0) = \infty \dots \end{array} \right.$$

incoming
spherical
wave

outgoing
spherical
wave

(26)

Now, the generalization of the one-dim. flux:

$$\vec{j} = \frac{\hbar}{2i\mu} \left[\psi^*(\vec{r}) \nabla \psi(\vec{r}) - (\nabla \psi^*(\vec{r})) \psi(\vec{r}) \right] \quad (27)$$

We shall see that it is only the flux in the radial direction that is of interest for large r .

$$\int dr \hat{e}_r \cdot \vec{j}(r) = \frac{\hbar}{2i\mu} \int dr \left(\psi^* \frac{\partial \psi}{\partial r} - \frac{\partial \psi^*}{\partial r} \psi \right) \quad \text{radial flux integrated over all angles} \quad (28)$$

$$\text{Remark: } \nabla \psi = \hat{e}_r \frac{\partial \psi}{\partial r} + \hat{e}_\theta \frac{1}{r} \frac{\partial \psi}{\partial \theta} + \hat{e}_\varphi \frac{1}{r \sin \theta} \frac{\partial \psi}{\partial \varphi}$$

$$\text{For a sol. of the form; } \psi(\vec{r}) = C \frac{e^{\pm ikr}}{r} Y_\ell^m(\theta, \varphi)$$

$$\int dr j_r = \frac{\hbar}{2i\mu} |C|^2 \left[\frac{e^{\mp ikr}}{r} \left(\pm ik \frac{e^{\pm ikr}}{r} - \frac{e^{\pm ikr}}{r} \right) - \text{Complex conjugate} \right]$$

$$= \pm \frac{\hbar k |C|^2}{\mu} \frac{1}{r^2} \quad \left(\begin{array}{l} + \rightarrow \text{outgoing flux} \\ - \rightarrow \text{incoming} \end{array} \right) \quad (29)$$

where we have used, $\int d\Omega |Y_{\ell}^m(\theta, \varphi)|^2 = 1$

Remark: $\int d\Omega Y_{\ell}^{m'}(\theta, \varphi) Y_{\ell}^m(\theta, \varphi) = \delta_{\ell\ell'} \delta_{mm'}$ (30)

(29) $\rightarrow \int r^2 d\Omega j_r = \text{indep. of } r$ (31) (total flux going through the spherical surface at radius r)

For our sol. (26)

$$\int r^2 d\Omega j_r = -\frac{\hbar k}{\mu} \left| \frac{r}{2ik} e^{i\ell\Omega/2} \right|^2 = -\frac{\hbar k}{\mu} \frac{1}{4k^2} \quad \text{incoming flux}$$

$$\int r^2 d\Omega j_r = +\frac{\hbar k}{\mu} \frac{1}{4k^2} \quad \text{outgoing flux} \quad (40)$$

Total flux = 0 (Since there is no source of flux)

In general by the flux conservation, for a sol. of the form

$$R_{\ell}(r) \underset{r \rightarrow \text{large}}{\approx} -\frac{1}{2ikr} \left[e^{-i(kr - \ell\Omega/2)} - S_{\ell}(k) e^{i(kr - \ell\Omega/2)} \right] \quad (V(r) \neq 0) \quad (41)$$

We must have $|S_{\ell}(k)|^2 = 1$ (42)

$S_{\ell}(k)$ can be written in the form, $S_{\ell}(k) = e^{2i\delta_{\ell}(k)}$ (43)

(41) $\rightarrow R_{\ell}(r) e^{-2i\delta_{\ell}(k)} \approx -\frac{1}{2ikr} \left[e^{-i(kr - \ell\Omega/2 + \delta_{\ell}(k))} - e^{i(kr - \ell\Omega/2 + \delta_{\ell}(k))} \right]$

$$\rightarrow R_\ell(r) \approx e^{i\delta_\ell(k)} \frac{S_\ell [kr - \ell\pi/2 + \delta_\ell(k)]}{kr} \quad (44)$$

$\rightarrow \delta_\ell(k)$: phase shift

Note that for free particle sol. we had:

$$J_\ell(kr) \xrightarrow{\text{asymptotic form}} \frac{S_\ell [kr - \ell\pi/2]}{kr} \quad (45)$$

Now, note that the flux in the \hat{e}_θ -dir involves,

$$\hat{e}_\theta \cdot \mathbf{J} = \frac{\hbar}{2im} \left(\psi^* \frac{1}{r} \frac{\partial}{\partial \theta} \psi - c.c. \right) \sim \frac{1}{r^3}$$

$$\hat{e}_\theta \cdot \mathbf{J} (r^2 dr) \sim \frac{1}{r} \quad (46)$$

\rightarrow So for $r \rightarrow \infty$ the dominant term is only radial flux.

B. The Square Well, Bound State

$$V(r) = \begin{cases} -V_0 & r < a \\ 0 & r > a \end{cases} \quad (47)$$

The radial eqn:

$$\frac{d^2 R}{dr^2} + \frac{2}{r} \frac{dR}{dr} - \frac{l(l+1)}{r^2} R + \frac{2M}{\hbar^2} (V_0 + E) R = 0 \quad r < a \quad (48)$$

$$\frac{d^2 R}{dr^2} + \frac{2}{r} \frac{dR}{dr} - \frac{l(l+1)}{r^2} R + \frac{2\mu}{\hbar^2} E R = 0 \quad r > a$$

We look for bound state sols. for which $E < 0$

$$\frac{2\mu}{\hbar^2} (V_0 + E) \equiv k^2, \quad \frac{2\mu}{\hbar^2} E = -\alpha^2 \quad (49)$$

For $r < a$:

$$R(r) = A j_l(kr) \quad \text{regular at the origin}$$

For $r > a$

$$R(r) \rightarrow 0 \quad \text{as } r \rightarrow \infty \quad (50)$$

The sol.: spherical Bessel func.

$$R(r) = B h_e^{(1)}(i\alpha r)$$

where k is replaced by $i\alpha$ in equ. (18)

Remark: $h_e^{(1)}(\rho) \approx -\frac{i}{\rho} e^{i(\rho - l\pi/2)}$
 $\rho \rightarrow \text{large}$
 $\rho \gg l$

In our case $\rho \rightarrow i\alpha r$, $\Rightarrow h_e^{(1)}(i\alpha r) \approx -\frac{1}{2r} \underbrace{e^{-\alpha r}}_{\text{decreasing}} e^{i\pi/2}$
 satisfying equ. (50)
 (51)

i) The two sols. must match at $r=a$

ii) and so must the derivatives

$$\rightarrow [A j_l(kr)]_{r=a} = [B h_l^{(1)}(iar)]_{r=a}$$

$$\left[\frac{d}{dr} (A j_l(kr)) \right]_{r=a} = \left[\frac{d}{dr} (B h_l^{(1)}(iar)) \right]_{r=a}$$

$$\rightarrow R \left[\frac{d}{d(kr)} (A j_l(kr)) \right]_{r=a} = i\alpha \left[\frac{d}{d(iar)} (B h_l^{(1)}(iar)) \right]_{r=a}$$

$$\rightarrow R \left[\frac{d j_l(\rho)/d\rho}{j_l(\rho)} \right]_{\rho=ka} = i\alpha \left[\frac{d h_l^{(1)}(\rho)/d\rho}{h_l^{(1)}(\rho)} \right]_{\rho=iaa} \quad (52)$$

This is a very complicated transcendental equ. involving $l, V_0,$ and E .

Special Case: $l=0$, we use the func. $u(r) = rR(r)$
(for simplification);

$$\left[\frac{u(r)}{r} \right]_{\text{left}}^{r=a} = \left[\frac{u(r)}{r} \right]_{\text{right}}^{r=a} \rightarrow [u(r)]_L^{r=a} = [u(r)]_R^{r=0} \quad (53)$$

$$\left[\frac{d}{dr} \left(\frac{u}{r} \right) \right]_L^{r=a} = \left[\frac{d}{dr} \left(\frac{u}{r} \right) \right]_R^{r=a} \quad (54)$$

$$\rightarrow \left[\frac{1}{r} \frac{du}{dr} - \frac{u}{r^2} \right]_L = \left[\frac{1}{r} \frac{du}{dr} - \frac{u}{r^2} \right]_R \quad (55)$$

$$(53) (54) \rightarrow \left[\frac{1}{r} \frac{du}{dr} \right]_L = \left[\frac{1}{r} \frac{du}{dr} \right]_R$$

$$\rightarrow \left[\frac{du}{dr} \right]_L = \left[\frac{du}{dr} \right]_R \quad (56)$$

$$(21) \rightarrow U(r)_L = A r J_0(kr) = \frac{A}{k} \Sigma_1(kr)$$

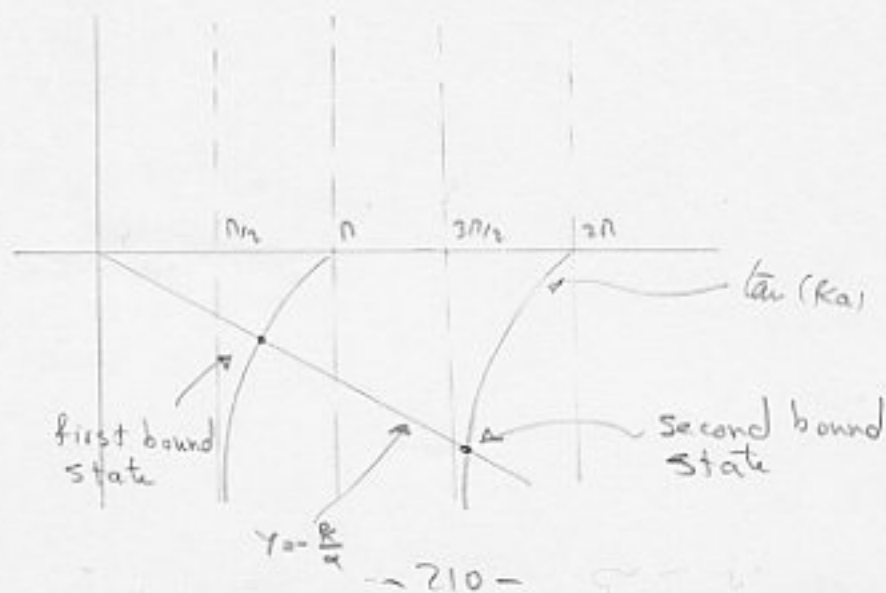
$$(23) \rightarrow U(r)_R = B r h_0^{(1)}(\alpha r) = -\frac{B}{\alpha} e^{-\alpha r}$$

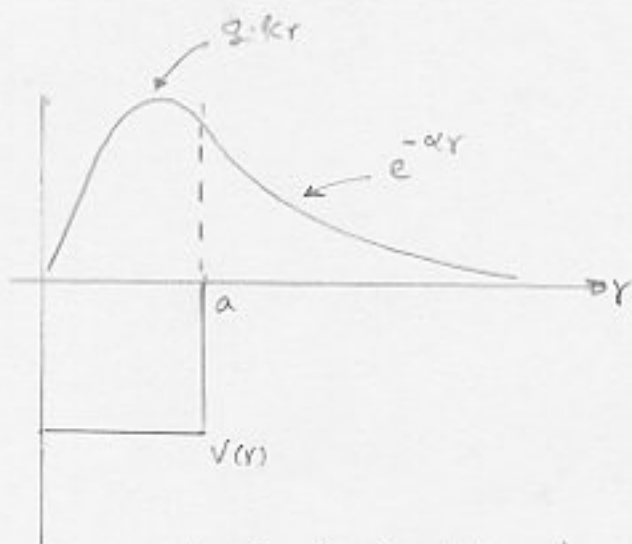
$$\left[\frac{A}{k} \Sigma_1(kr) \right]_{r=a} = \left[-\frac{B}{\alpha} e^{-\alpha r} \right]_{r=a}$$

$$\left[\frac{A R}{k} G_1(kr) \right]_{r=a} = \left[+\frac{B \alpha}{\alpha} e^{-\alpha r} \right]_{r=a}$$

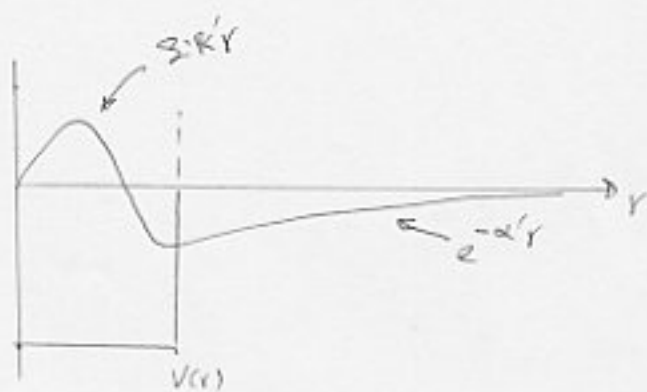
$$\rightarrow \frac{\Sigma_1(ka)}{R G_1(ka)} = \frac{-e^{-\alpha a}}{\alpha e^{-\alpha a}} \rightarrow \tan(ka) = -\frac{k}{\alpha}$$

$$\text{Since } k > 0, \alpha > 0 \rightarrow -\frac{k}{\alpha} < 0$$





Sol. for the first bound state ($U(r) = R(r)$, $l=0$)



Sol. for the second bound state

Now, for very deep pot. for which $Ka \gg l$, we are justified in using the asymptotic form of $J_l(x)$ (equ. 25);

$$(52) \rightarrow -\frac{1}{a} + k \cot\left(Ka - \frac{l\pi}{2}\right) = \text{R.H.S.} \quad (57)$$

does not contain V_0

The R.H.S. does not contain V_0 (which is large) and if $|E| \ll V_0$,

$$\cot\left(Ka - \frac{l\pi}{2}\right) = \frac{\text{R.H.S.}}{k} + \frac{1}{Ka}$$

\rightarrow the largeness of Ka implies that $\cot\left(Ka - \frac{l\pi}{2}\right) \approx 0$

$$\rightarrow Ka - \frac{l\pi}{2} \approx \left(n + \frac{1}{2}\right)\pi \quad (58)$$

Since $|E| \ll V_0$, (49) $\rightarrow k \approx k_0 \left(1 + \frac{E}{2V_0}\right)$ (59)

where $k_0 = \frac{2\mu V_0}{\hbar^2}$ (60)

$$(58) (5a) \Rightarrow \frac{E}{2V_0} = -1 + \frac{[n + (l+1)/2] \pi}{k_0 a} \quad (61)$$

For $n \rightarrow$ large (far from the bottom of the well)

for all $l \ll k_0 a$, the energy levels are approximately equally spaced;

$$\frac{\Delta E}{2V_0} \approx \frac{\pi}{k_0 a} \quad (62)$$

A related Prob.: The infinite box in 3-dim:

$$V(r) = \begin{cases} 0 & r < a \\ \infty & r > a \end{cases} \quad (63)$$

$$\frac{2AE}{\hbar^2} = k^2 \quad (64)$$

$$R(r) = A j_l(kr) \quad \text{regular at } r=0 \quad (r < a)$$

$$R(r) = 0 \quad r \geq a \quad (65)$$

$$\rightarrow j_l(ka) = 0 \quad (66)$$

		Roots						
ka	l	0	1	2	3	4	5	
		3.14	4.49	5.76	6.99	8.18	9.36	$n=1$ first root
		6.28	7.73	9.10	10.42			$n=2$ second \rightarrow
		9.42						$n=3$ - - -

Using the spectroscopic notation;

$$S : l=0$$

$$P : l=1$$

$$D : l=2$$

$$F : l=3$$

$$G : l=4$$

(67)

Then the order of the levels are;

1S, 1P, 1D, 2S, 1F, 2P, 1G, 2D, 1H, 3S

A model of nucleus:

Protons and Neutrons inside such an infinite box.

For spin $\frac{1}{2}$ particles (fermions) each level is occupied only with 2-particles -

Consider Protons;

$$1S : \quad 2 \text{ - protons} \quad (l=0)$$

$$1P : \quad 6 \text{ - " } \quad (l=1, m=-1, 0, +1)$$

$$1D : \quad 10 \text{ - " } \quad (l=2 \text{ - - - })$$

$$2S : \quad 2 \text{ - " } \quad (l=0)$$

$$1F : \quad 14 \text{ - " } \quad (l=3)$$

$$2P : \quad 6 \text{ - " } \quad (l=1)$$

$$1G : \quad 18 \text{ - " } \quad (l=4)$$

$$2D : \quad 10 \text{ - " } \quad (l=2)$$

$$1H : \quad 22 \text{ - " } \quad (l=5)$$

$$3S : \quad 2 \text{ - " } \quad (l=0)$$

(68)

Thus the levels will be filled when the number of protons is:

$$2, 8 (= 2+6), 18 (= 2+6+10), 20 (= 18+2), 34 (= 20+14), \\ 40, 58, 68, 90, 92, 106 \quad (69)$$

and similarly for neutrons.

In real nuclei; the magic numbers are:

$$2, 8, 20, 28, 50, 82, 126, \dots \quad (70)$$

The nuclei having these number of protons or neutrons exhibit special characteristics.

C. The Square Well Continuum Sols.

With $E > 0$, we write $\frac{2\mu E}{\hbar^2} = K^2 \quad (71)$

For $r > a$, the sol. will be a combination of the regular and irregular sols. of the free field equ.

$$R_e(r) = B j_e(kr) + C n_e(kr) \quad (72)$$

$$\text{For } r < a \quad R_l(r) = A j_l(kr) \quad (73)$$

$$\text{where } k^2 = \frac{2\mu(E+V_0)}{\hbar^2} \quad (74) \quad (V_0 > 0 \text{ for attractive p.t.})$$

$$\text{Matching: } R \left[\frac{d j_l(s)/ds}{j_l(s)} \right]_{s=ka} = K \left[\frac{B \frac{d j_l(s)/ds}{j_l(s)} + C \frac{d n_l(s)/ds}{n_l(s)}}{B j_l(s) + C n_l(s)} \right]_{s=ka} \quad (75)$$

The asymptotic form of (72);

$$R_l(r) \approx \frac{B}{kr} \left[\sin\left(kr - \frac{l\pi}{2}\right) - \frac{C}{B} \cos\left(kr - \frac{l\pi}{2}\right) \right] \quad (76)$$

$$(44) \rightarrow R_l(r) \approx \frac{e^{i\delta_l(k)}}{kr} \left[\sin\left(kr - \frac{l\pi}{2}\right) \cos \delta_l(k) + \cos\left(kr - \frac{l\pi}{2}\right) \sin \delta_l(k) \right] \quad (77)$$

$$\text{With the assumption } e^{i\delta_l(k)} \approx 1 \quad (\text{for } \delta_l(k) \approx \text{small}) \quad (78)$$

$$(76)(77) \rightarrow \frac{C}{B} = -\tan \delta_l(k) \quad (79)$$

The actual computation of C/B from (75) is tedious, except for $l=0$

Once again $u(r) = r R(r)$

for $l=0$

$$\begin{aligned} r R(r) &= B \sin kr + C \cos kr & r > a \\ r R(r) &= A \sin kr & r < a \end{aligned} \quad (80)$$

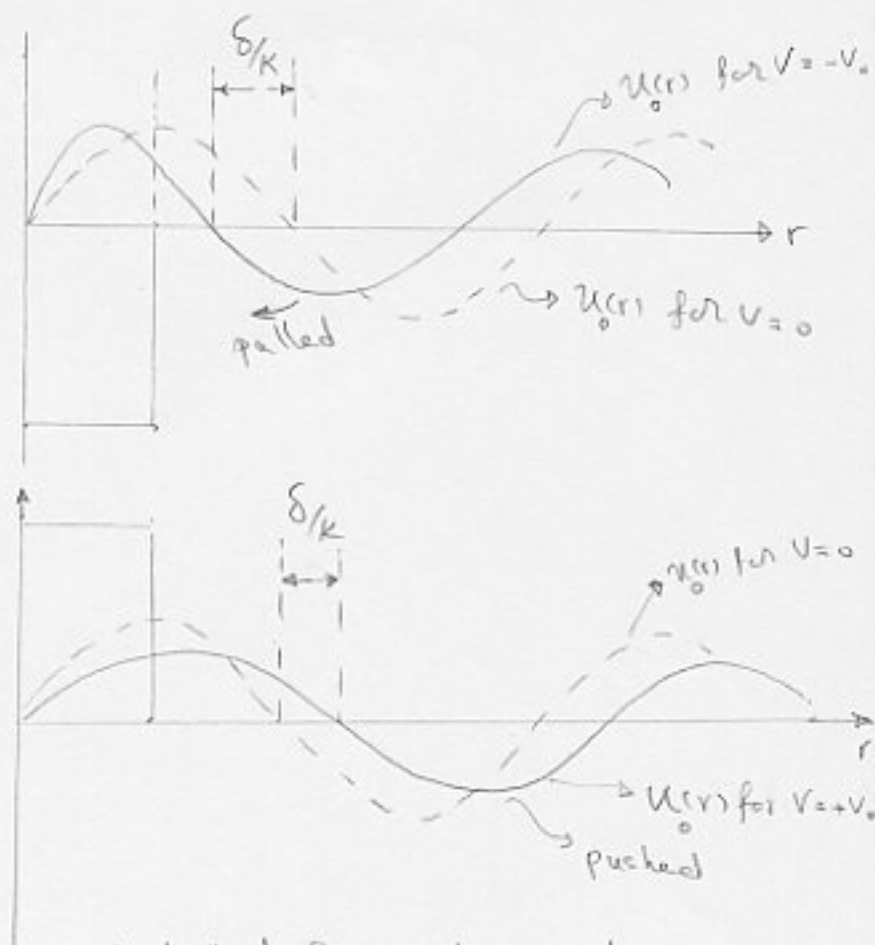
$$(53)(56) \rightarrow \begin{cases} A \sin ka = B \sin ka + C \cos ka \\ R A \cos ka = B K \cos ka - C k \sin ka \end{cases} \quad (81)$$

$$\rightarrow \frac{1}{R} \tan ka = \frac{B \sin ka + C \cos ka}{B K \cos ka - C k \sin ka} \quad (82)$$

Using (79) and (82) $\delta_{e=0}$ can be obtained

Indeed in general

$$(75)(79) \rightarrow \tan \delta_e(k) = \frac{k j_e'(ka) j_e(Ra) - R j_e(ka) j_e'(Ra)}{k \eta_e'(ka) j_e(Ra) - R \eta_e(ka) j_e'(Ra)} \quad (83)$$



Note that for repulsive pot. equ. (74) must be changed to $k^2 = 2\mu(E - V_0)$, where $V_0 > 0$ and k may or may not be real.

Important relation :

By superposition principle we have the following sol. for the free-particle equ. (using equ. (26));

$$\Psi(\vec{r}) = \sum_{\ell=0}^{\infty} \sum_{m=-\ell}^{\ell} A_{\ell m} J_{\ell}(kr) Y_{\ell m}(\theta, \varphi) \quad (84)$$

Another sol. of the free-particle equ., which reads

$$(\nabla^2 + k^2) \Psi(r) = 0 \quad (85)$$

before the separation into angular and radial part is made, is the plane wave;

$$\Psi(r) = e^{ik \cdot r} \quad (86)$$

Now if we define the z-axis by the dir. of \vec{k} ;

$$\Rightarrow e^{ik \cdot r} = e^{ikr \cos \theta} \quad (87)$$

$\rightarrow \Psi(\vec{r})$ has no φ -dependence

$\rightarrow m=0$ only in (84)

$$\text{Using } Y_{\ell}^0(\theta, \varphi) = \left(\frac{2\ell+1}{4\pi}\right)^{1/2} P_{\ell}(\cos \theta) \quad (88)$$

\rightarrow Legendre polynomials

$$\rightarrow e^{ikr\cos\theta} = \sum_{\ell=0}^{\infty} \left(\frac{2\ell+1}{4\pi}\right)^{1/2} A_{\ell} j_{\ell}(kr) P_{\ell}(\cos\theta) \quad (89)$$

Now using the $\frac{1}{2} \int_{-1}^1 d(\cos\theta) P_{\ell}(\cos\theta) P_{\ell'}(\cos\theta) = \frac{\delta_{\ell\ell'}}{2\ell+1}$ (90)

$$\rightarrow A_{\ell} j_{\ell}(kr) = \frac{1}{2} [4\pi(2\ell+1)]^{1/2} \int_{-1}^1 dz P_{\ell}(z) e^{ikrz} \quad (91)$$

Now using: $j_{\ell}(s) = \frac{1}{2i^{\ell}} \int_{-1}^1 ds e^{is} P_{\ell}(s)$ (92)

$$\rightarrow A_{\ell} j_{\ell}(kr) = \frac{1}{2} [4\pi(2\ell+1)]^{1/2} (2i^{\ell} j_{\ell}(kr)) \quad (92)$$

$$\rightarrow A_{\ell} = i^{\ell} [4\pi(2\ell+1)]^{1/2} \quad (93)$$

(86) (88) (87) (93) $\rightarrow e^{ikr\cos\theta} = \sum_{\ell=0}^{\infty} (2\ell+1) i^{\ell} j_{\ell}(kr) P_{\ell}(\cos\theta)$ (94)