Lorentz (or Length) Contraction

Let's return to our example, and consider just the one-way travel of the spaceship from Earth to the star. In the Earth-star reference frame A, the spaceship travels 10 ly in 12.5 years at speed 0.8c. So the distance between Earth and the star, as seen in A is:

$$\Delta x_{one-way} = v\Delta t_{one-way} = v\frac{L}{\beta} = cL$$
 (m) = L (lightyears) = 10 ly

In the spaceship reference frame A', we already found that:

$$\Delta t'_{\text{one-way}} = \frac{\Delta t}{1/\sqrt{1 - v^2/c^2}} = 7.5 \text{ years}$$

Because the Earth and star move past the spaceship at speed v = 0.8c, the distance between Earth and the star, as seen in A', is

$$\Delta x'_{one-way} = v \Delta t'_{one-way} = v \frac{\Delta t}{1/\sqrt{1 - v^2/c^2}} = \Delta x \sqrt{1 - v^2/c^2} = (10 \text{ ly})(\sqrt{1 - 0.8^2}) = 6.0 \text{ ly}$$

This indicates that the distance between any two points is always greatest in a frame (the Earth-star frame A in our example) fixed with respect to those points. In any other reference frame is smaller. This is called the <u>Lorentz contraction</u>.

In general: $\Delta x' = \Delta x \sqrt{1 - \frac{v^2}{c^2}} = \frac{\Delta x}{\gamma}$ where A' is moving w.r.t. the two points.

Note: an object's length in a frame in which it is at rest is called its proper length.

Lorentz Velocity Addition

Let's say that object O has velocity u' in reference frame A'.

What will the velocity of the object be as measured by an observer in A?



Galilean Approach: $v_A^O = v_{A'}^O + v_A^{A'}$ $\therefore u = u' + v$ Does this still hold? No!e.g., if v = 0.75 c and u' = 0.80 c, then u = 1.55 c > c !

STR Approach: Apply the Lorentz transformations and take the differentials. $x = \gamma(x'+vt') \rightarrow dx = \gamma(dx'+vdt')$

$$t = \gamma \left(t' + \frac{v}{c^2} x' \right) \rightarrow dt = \gamma \left(dt' + \frac{v}{c^2} dx' \right)$$

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 $u = \frac{dx}{dt} = \frac{dx' + vdt'}{dt' + \frac{v}{c^2} dx'} = \frac{\frac{dx'}{dt'} + v}{1 + \frac{v}{c^2} \frac{dx'}{dt'}}$ Thus $u = \frac{u' + v}{1 + \frac{v}{c^2} u'}$ and $u' = \frac{u - v}{1 - \frac{v}{c^2} u}$ Special cases: (1) u' << c and v << c: $u \equiv u' + v$ and $u' \equiv u - v$ i.e., we get the Galilean velocity transforms (2) u' \approx c and v \approx c: Let's say: $u' = (1 - \delta_1)c: \delta_1 \approx 0, \delta_1 > 0$ $v = (1 - \delta_2)c \delta_2 \approx 0, \delta_2 > 0$ Then: $u = \frac{c(2 - \delta_1 - \delta_2)}{1 + (1 - \delta_1)(1 - \delta_2)} = \frac{c(2 - \delta_1 - \delta_2)}{2 - \delta_1 - \delta_2 + \delta_1\delta_2} < c!!$

So u is still less than c. Note: it would be 2c using the Galilean velocity transform.

Now, let's make the x component of u and u' explicit: $u_x = \frac{u'_x + v}{1 + \frac{v}{c^2}u'_x}$

where v is the relative velocity of A' to A along the <u>x axis</u>. What happens to the velocity perpendicular to the direction of motion of the two reference frames?

Galilean Approach: $u_v = u'_v$

With STR, we still have y = y', but $t \neq t'$. Again apply the Lorentz transformations and take the differentials.

$$y = y' \rightarrow dy = dy'$$

$$t = \gamma \left(t' + \frac{v}{c^2} x' \right) \rightarrow dt = \gamma \left(dt' + \frac{v}{c^2} dx' \right)$$

$$dv'$$

Thus: $u_y = \frac{dy}{dt} = \frac{dy'}{\gamma\left(dt' + \frac{v}{c^2}dx'\right)} = \frac{\frac{dy'}{dt'}}{\gamma\left(1 + \frac{v}{c^2}\frac{dx'}{dt'}\right)} = \frac{u'_y}{\gamma\left(1 + \frac{v}{c^2}u'_x\right)}$ v in x,x' direction

So $u_y \neq u'_y$ Note: even if $u'_x = 0$, so that the object only has a y' velocity component in reference frame A', we still have $u_y \neq u'_y$. In this case: $u_y = u'_y / \gamma$. We can derive a similar equation for the z component: $u_z = \frac{u'_z}{\gamma \left(1 + \frac{v}{c^2} u'_x\right)}$