#### A Journey Back in Time to the Big Bang



- Pulling back from the WMAP spacecraft, we see that we are a tiny speck in our Milky Way Galaxy.
- As the journey continues, we see that the Milky Way is but a tiny speck in the Universe as we pass quasars that were some of the early brightest structures that we can detect with conventional instruments.
- Finally we arrive at the beginning of time as we understand it. The super heated, roiling ionized hydrogen gas of a universe newly made glows throughout the universe. The energy from sonically condensed and rarified ripples in this gas, released as it cooled to 3000°K, is the basis for the data collected by WMAP.
- Credit: NASA / WMAP Science Team, http://map.gsfc.nasa.gov/media/030657/index.html

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# **Current Assignments ...**

#### For today

• Read Sections 11.2 - 11.7

#### For Lecture 19

- Read Chapter 12
- Homework #3

#### Office hours: 3-4 Tuesdays & Thursdays

- Late deadline 11:00 AM, Friday, March 15
  Homework #4
- Posted March 7. Due 11:00 AM, Friday, March 22
  Writing Assignment #2
- Posted Feb. 28. Due 11:00 AM, Thursday, April 4
- **Suggested Conceptual Exercises**
- Ch. 11: 1,3,5,7,9,11,13,15,17,19,21,23,25,27,29,31,33,35,37,39,41,43,45
  Tutorial #8

# **Review of Lecture 17**

#### Textbook, Sections 11.1 - 11.2

- General relativity
- The big bang

*"Matter tells spacetime how to curve, and curved spacetime tells matter how to move".* 

# Plan for Lecture 18

## Textbook, Sections 11.2 - 11.7

- The big bang
- The shape of the universe
- Dark matter and dark energy
- Cosmic inflation

# From L17: Evidence for the Big Bang

- In 1929 it was discovered that the universe was expanding; extrapolating backwards led to the big bang.
- 2. The cosmic microwave background (CMB), left over from the big bang, has been observed and agrees with theoretical predictions.
- **3.** The CMB has been mapped in great detail; its small variations are just as they should be to create galaxies as we see them today.
- 4. Theory predicts just which elements, and in what ratios, should be produced in the big bang; these agree well with observations. PHY1005 (K. Strong) - Lecture 18 - Slide 5

# The Oldest Light in the Universe



NASA's Wilkinson Microwave Anisotropy Probe (WMAP) http://science.nasa.gov/headlines/y2003/11feb\_map.htm Credit: NASA/WMAP Science Team



#### Credit: NASA/WMAP Science Team

http://map.gsfc.nasa.gov/m\_ig/060915/CMB\_Timeline300/jp/gpos (K. Strong) - Lecture 18 - Slide 7

# The Oldest Light in the Universe







http://map.gsfc.nasa.gov/ media/030651/index.html

NASA's Wilkinson Microwave Anisotropy Probe (WMAP)

http://science.nasa.gov/headlines/y2003/11feb\_map.htm Credit: NASA/WMAP Science Team

# The Expanding Universe

- The big bang was not really an explosion.
- It created space and time.
- The expanding universe continues to create spacetime.

## Textbook Figure 11.14

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 It is not expanding into anything.

The expanding surface of a balloon is a 2D analog of 3D space.

# The Shape of the Universe

#### **Three possibilities:**

- Closed like the surface of a sphere. Parallel lines eventually meet.
- Open like a saddle, infinite in extent. Parallel lines diverge.
- Flat no curvature, extends infinitely far in all directions. Parallel lines remain parallel.



http://map.gsfc.nasa.gov/universe/bb\_concepts.html PHY100S (K. Strong) - Lecture 18 - Slide 10

# Which Shape Is the Universe?



http://map.gsfc.nasa.gov/

media/030639/index.html

## Textbook Figure 11.16

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The measured angle is close to  $1^{\circ}$  $\rightarrow$  consistent with a flat universe.

# What is the Universe Made Of?

## Many forms of matter:

- protons, neutrons, electrons (form atoms)
- neutrinos
- black holes
  - → regions of spacetime from which nothing can escape, even light
- dark matter
  - $\rightarrow$  does not interact with EM radiation
  - $\rightarrow$  can be detected due to its gravitational effects
  - -> comprises most of the mass of the universe PHY100S (K. Strong) - Lecture 18 - Slide 12

http://astro.berkeley.edu/~mwhite/darkmatter/dm.html

Not dark matter

> Dark matter



# What Is Dark Matter?

## We don't know! Some possibilities:

- MACHOs (MAssive Compact Halo Objects), including brown dwarfs
  - → Dim objects, intermediate between stars and planets, that are not luminous enough to be directly detectable by telescopes.
- Supermassive black holes
- WIMPs Weakly Interacting Massive Particles
  - New forms of matter, maybe particles produced shortly after the big bang.

# **Detection of Dark Matter - 1**

- By measuring the motions of stars and gas, astronomers can "weigh" galaxies.
- The mass of the galaxies, including the Milky Way, is ~10 times larger than the mass that can be associated with stars, gas and dust.

## Textbook Figure 11.20

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#### → Dark matter provides this mass.

# **Detection of Dark Matter - 2**

 Galaxies can also be "weighed" by measuring how they distort light coming from other galaxies (gravitational lensing)
 → Again, there is missing mass - dark matter.



NASA Hubble Space Telescope image of the rich galaxy cluster, Abell 2218.

The arc-like pattern spread across the picture like a spider web is an illusion caused by the gravitational field of the cluster.

Credits: W.Couch (University of New South Wales), R. Ellis (Cambridge University), and NASA http://hubblesite.org/newscenter/archive/releases/1995/14/image/a/ PHY100S (K. Strong) - Lecture 18 - Slide 15

# **The Accelerating Universe**

- Because of the gravitational attraction between all the matter in the universe, we might expect its expansion to be slowing.
- 1998 Observations of exploding supernovas gave information about distances, speeds, and accelerations across the universe.

→The most distant galaxies were too far away to be explained without acceleration.

• The result: The expansion of the universe is apparently accelerating!

# **Dark Energy**

 This acceleration cannot be driven by any matter or field that we know of

 $\rightarrow$  It must be driven by something new.

- This energy that is slowly pushing the universe apart is called <u>dark energy</u>.
- When the mass of the dark energy and dark matter is added to the luminous and nonluminous matter, the result is just enough for the universe to be flat.

## What is the Universe Made Of?

## Textbook Figure 11.21

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# The Fate of the Universe: endless expansion or big crunch?



• Orange = closed, high density universe which expands for several billion years, then ultimately turns around and collapses under its own weight.

• Green = flat, critical density universe in which the expansion rate continually slows down (the curves becomes ever more horizontal).

• Blue = open, low density universe whose expansion is also slowing down, but not as much as the previous two because the pull of gravity is not as strong.

• Red = universe in which a large fraction of the matter is in a form dubbed "dark energy" which is causing the expansion of the universe to speed up (accelerate).

• There is growing evidence that our universe is following the red curve.

http://map.gsfc.nasa.gov/universe/bb\_concepts\_exp.html http://map.gsfc.nasa.gov/universe/uni\_fate.html PH

# The Fate of the Universe

- Universe 1 accelerated expansion with dark energy
- http://snap.lbl.gov/multimedia/animations/ accel\_full.avi
- Universe 2 the universal expansion, stopping and reversing
- http://snap.lbl.gov/multimedia/animations/ Expand&contract\_full.avi
- Universe 3 constant expansion with no dark energy
- http://snap.lbl.gov/multimedia/animations/ Expanding\_full.avi





