EFFECTS OF EPP-NOX AND SOLAR UV VARIATIONS ON OZONE IN THE POLAR STRATOSPHERE: CORRELATIVE STUDIES USING UARS HALOE AND SBUV(/2) DATA

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1992

1988

Figure 7: Same as Figure 6 but for the Northern Hemisphere.

1984

2000

2004



1 hPa SH Spring Polar NOx Time Series: **OBJECTIVES:** tic Ap Index vs. HALOE NO., 55°S to 75°S, SON 160 -160. ent (a) HALOE NO., 1 hPa, 55°S-75°S -(a) HALOE/ACE NO_ 1 ht • To evaluate the relative importance of er er solar proton events (SPE's) and ral Ap Index, JJASON magnetospheric electron precipitation R = 0.90- 0.86 (EEP) in producing downward descending index 10 NO_x in the springtime polar stratosphere (c) Solar Wind Plasma Speed, JJASO Pa JJASON on interannual timescales. 10 Proton Ion Prod., 1.3 hPa Proton Ion Prod., 1.3 hPa R = 0.77R = 0.54Ē • To evaluate the relative importance of 50 energetic particle precipitation induced Index Annual Mea Solar UV Index Annual changes in NO_x (EPP-NO_x) and changes toin solar ultraviolet (UV) spectral irradiance on polar ozone as a function raladadarbudi 1992 1994 of season. 1996 1998 2000 2002 1996 1998 2000 2002 2004 2006 .5 Correlation Coefficient Figure 1: HALOE NO + NO2 sunset data for Figure 2: HALOE NO + NO2 sunset data for Figure 3: Correlation vs. Ap as a function SH spring (SON) SON supplemented by 3 years of ACE data** of pressure level. HALOE NO_x and O₃ Data (1 hPa, SH Polar, Spring): SBUV (/2) O₃ (1 hPa, Polar, Monthly Averages): set) 55°S to 75°S SON (a) SBUV (/2) 1 hPa 0, Anomaly, 70°S-90°S (a) HALOE 03, 1 hPa, 60°S-70°S, SO (b) HALOE NOz, 1 hPa, 60°S-70°S. So hPa (c) Auroral Ap Index, SON R = -0.7120 15 Core-10 50 1988 1992 1996 2000 2004 1992 1998 1994 1996 2000 2002 2004 100 Figure 6: Southern Hemisphere. Note positive correlation during the Figure 4: HALOE O3 and NO + NO2 sunset data for SH summer season when polar ozone is a minimum. spring (SON). A clear negative correlation is evident. Also shown is the Ap index, which is closely correlated with the -.5 0 Correlation Coefficient (/2) 1 1 NO. data Figure 5: Correlation of HALOE NO, vs. O3 as a function of pressure level. ** ACE data provided by Cora Randall, P. Bernath, and ACE team. Solar Proton Ion production rates provided by C. Jackman j.

CONCLUSIONS:

• In the Southern Hemisphere where correlations are most significant, results show that the high-latitude spring HALOE NO, data for the upper stratosphere and lower mesosphere correlate best (up to 0.86 at 1 hPa) with the Ap index, a measure of magnetospheric electron precipitation. Therefore, although SPE's can dominate over short time periods or even a complete season, EEP appears to mainly determine interannual NO. variability over the whole 12vear interval analyzed here. Correlation coefficients continue to be significant down to about 10 hPa.

Also in the SH. a negative correlation is obtained between HALOE ozone and NO, (sunset) data at high latitudes during spring. This negative correlation is also significant down to about the 10 hPa level and is interpreted to be a consequence of direct photochemical destruction of ozone by downward descending EPP-NO_x. However, during the summer season, evidence is obtained for a positive solar cycle variation of SBUV ozone at high latitudes near the stratopause. This summertime ozone variation at high altitudes is interpreted to be a consequence of solar UVinduced changes in the ozone production rate.