

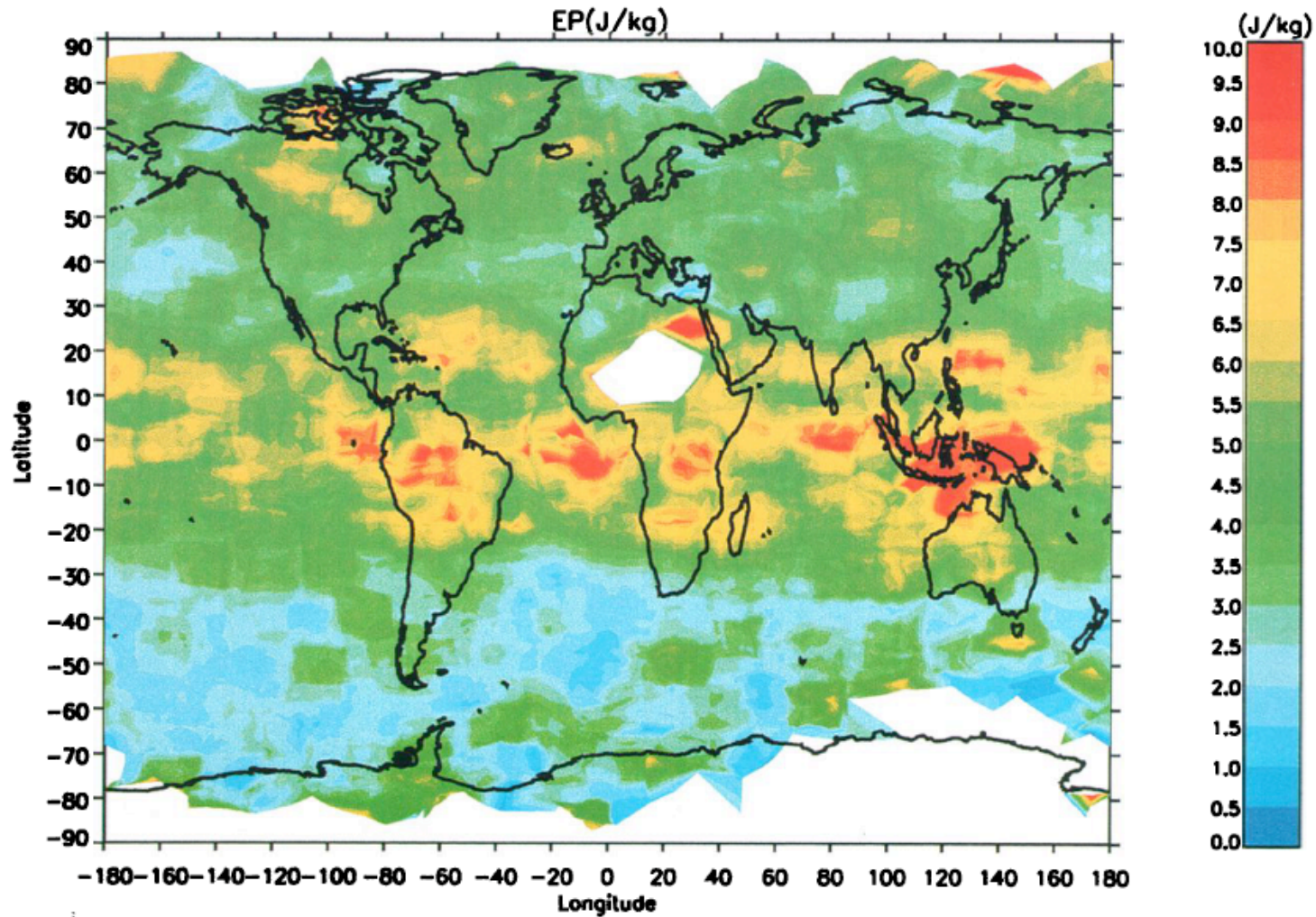
Spatial structures of stratospheric gravity waves derived from COSMIC GPS occultation data

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(RISH, Kyoto U)

GPS Radio Occultation (RO)

- GPS signals are received by low-earth-orbiting satellites
- refractivity profiles → Temperature (& humidity in lower-to-mid troposphere)

Tsuda et al (2000) with GPS/MET: 1st GPS RO mission

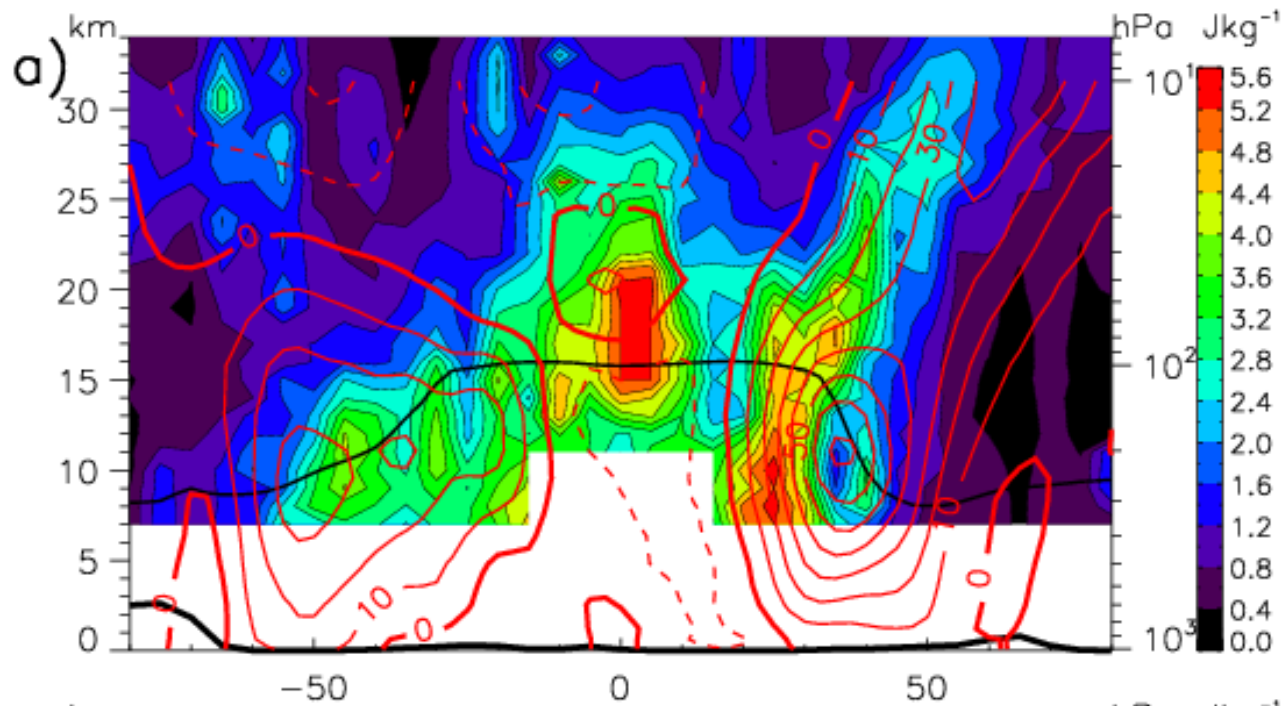


Global distribution of potential energy of stratospheric GWs

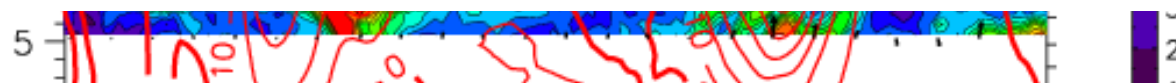
COSMIC / FORMOSAT-3

- by UCAR(US) & NSPO(Taiwan)
- Launch: Apr 2006
- 6 LEO satellites (recv 2 ways)
 - substantial increase of RO measurements (~2500 per day)

S. Alexander et al 2008



GW energy averaged between 130E-150E
23-28 Dec 2006 using COSMIC data



Purpose of this study

- To study spatial structures of gravity waves using multiple GPS RO profiles nearby
 - Earlier studies: profile-base
- Merit: Good vertical resolution (1km or better)
- Demerit: Horizontal resolution/sampling lower than conventional instruments (e.g. AIRS)

On the average, COSMIC sampling is still too sparse (1 per 500km*500km*1day) → Examine satellite orbits

- Launched by a single rocket (inclination 72°)
- Alt. raised 1 per ~2 months (500km → 800km)

*Liou et al
(2007)*

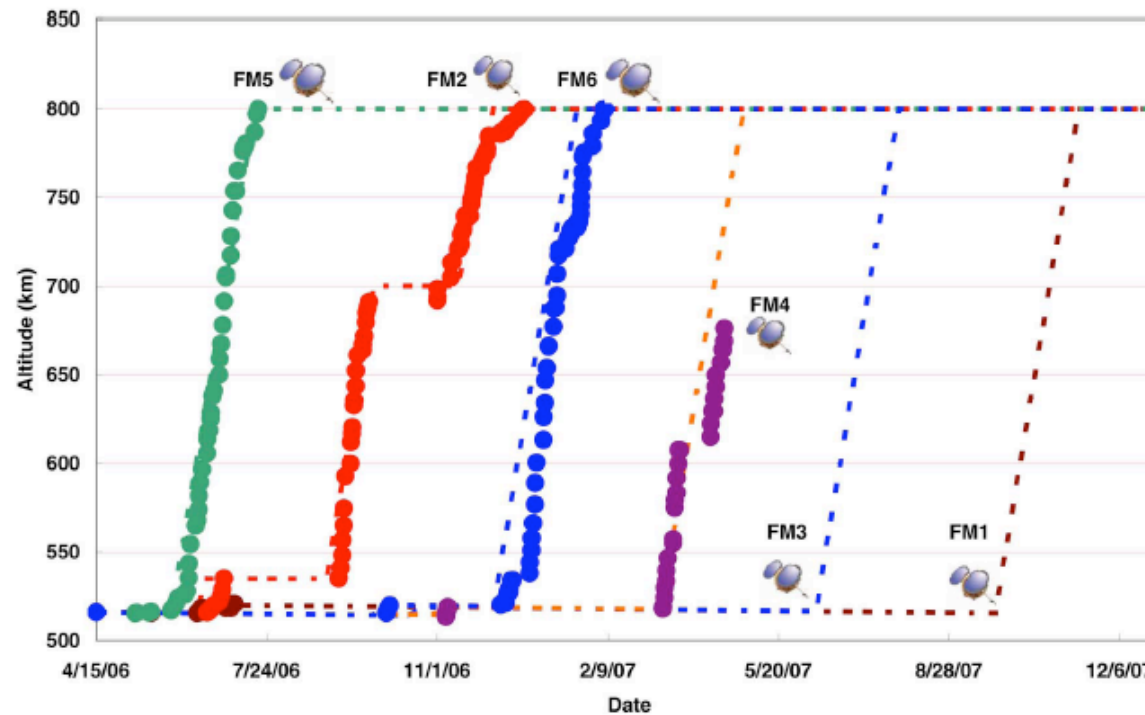
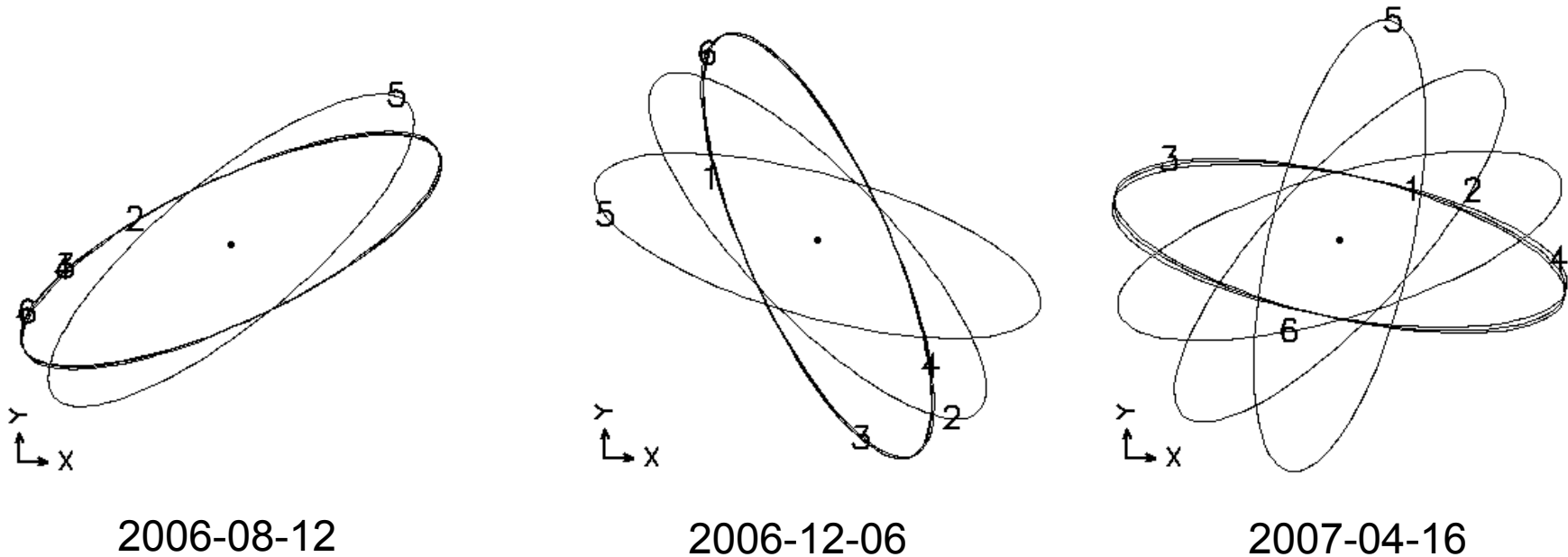


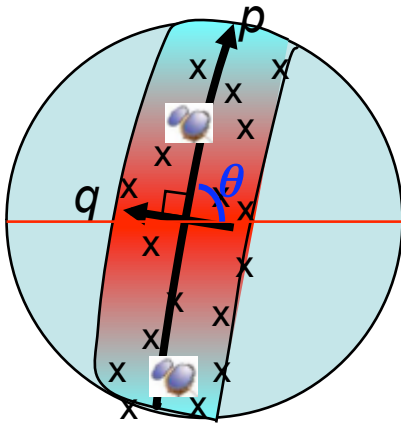
Fig. 4. FORMOSAT-3 constellation deployment profile versus time. The FORMOSAT-3 satellites flight model no. 2 (FM2), no. 5 (FM5), and no. 6 (FM6) have arrived to the 800-km orbit in early February 2007. The final mission constellation configuration will arrive to the 800-km orbit around November 30, 2007.

LEO orbits relative to the stars (ECI coordinate)



3 LEOs still on one orbit after 1 year since launch
→ High data acquisition rate around the orbit

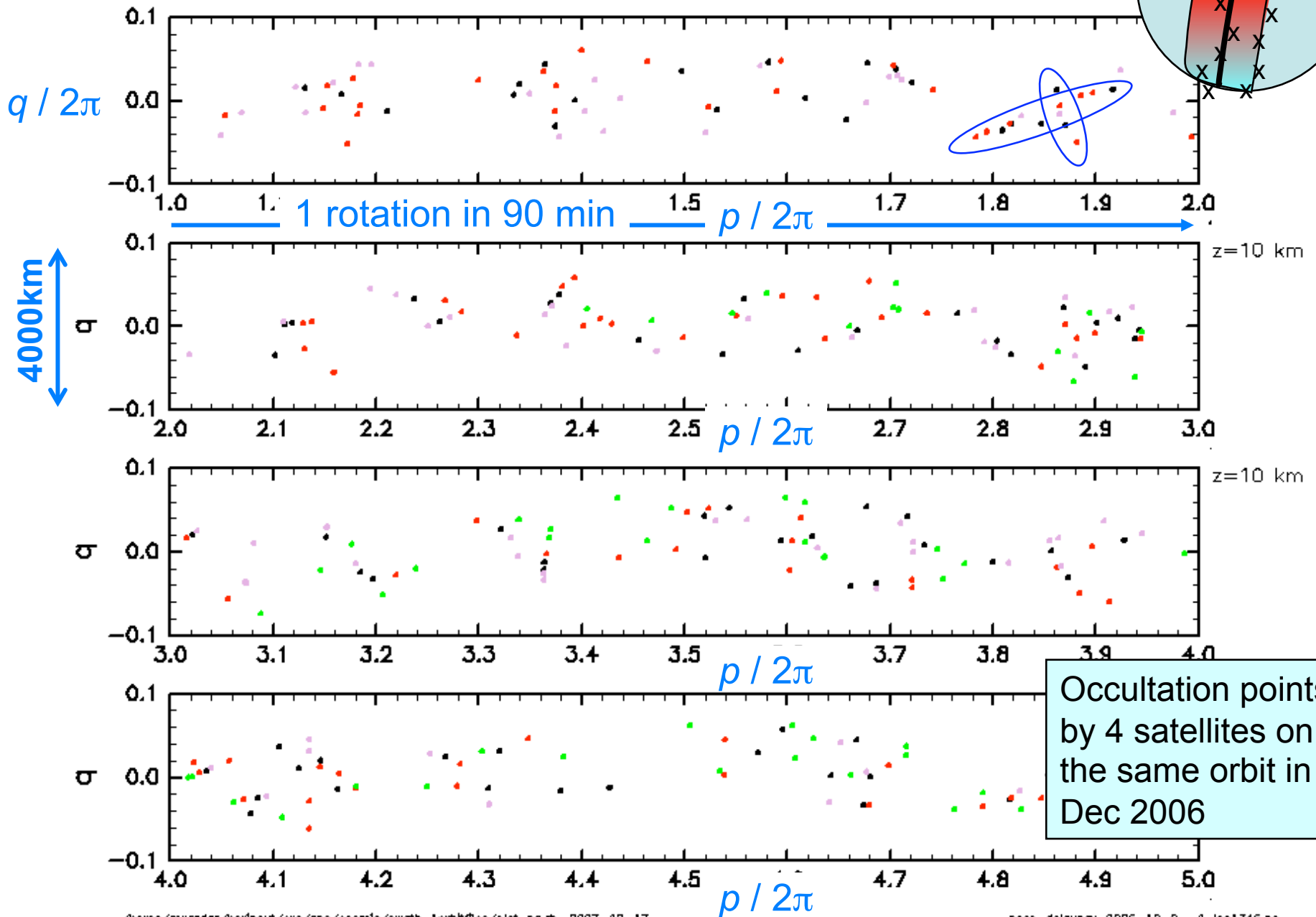
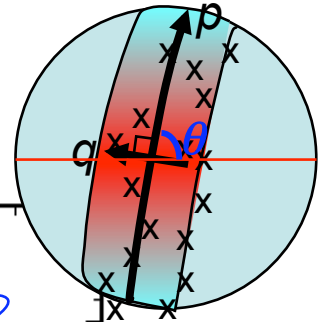
We focus on the neighborhood of the orbit
having multiple satellites



Coordinate transformation

- p : along the satellite track
- q : perp to p

Tangent points are found to be frequently organized in liner shapes



Occultation points by 4 satellites on the same orbit in Dec 2006

Strategy

- Extract groups of data points organized linearly → Obtain vertical cross sections
 - Snapshots within ~ 1 hours → Good to analyze short-lived disturbances with high vertical resolution

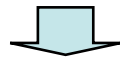
Dataset used

- COSMIC Level 2 by UCAR (dry temperature)
 - Used: Dec 2006 – Feb 2007

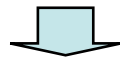
Heuristic Line Clustering

- Algorithm

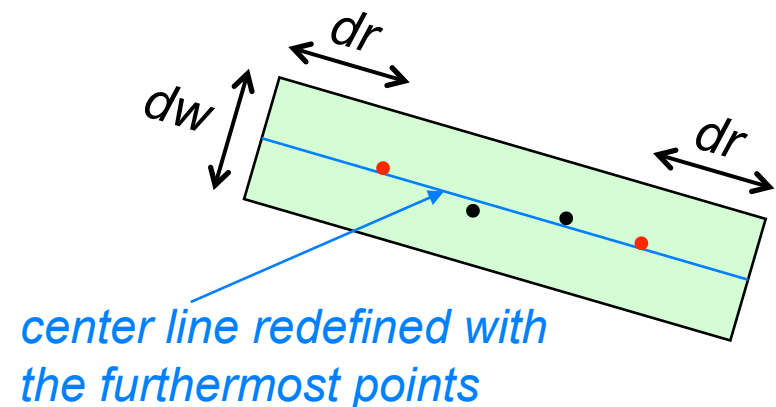
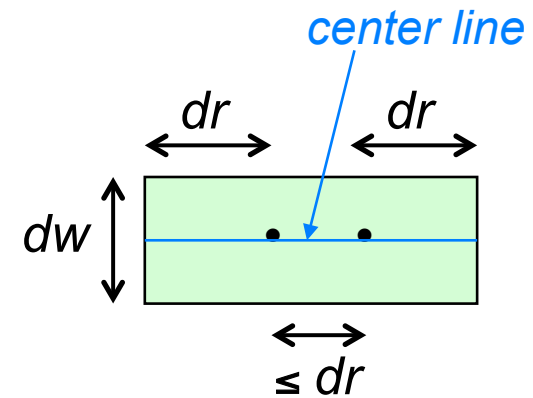
1. Initial pairs: nearest neighbors
(distance $\leq dr$, and \geq a threshold)



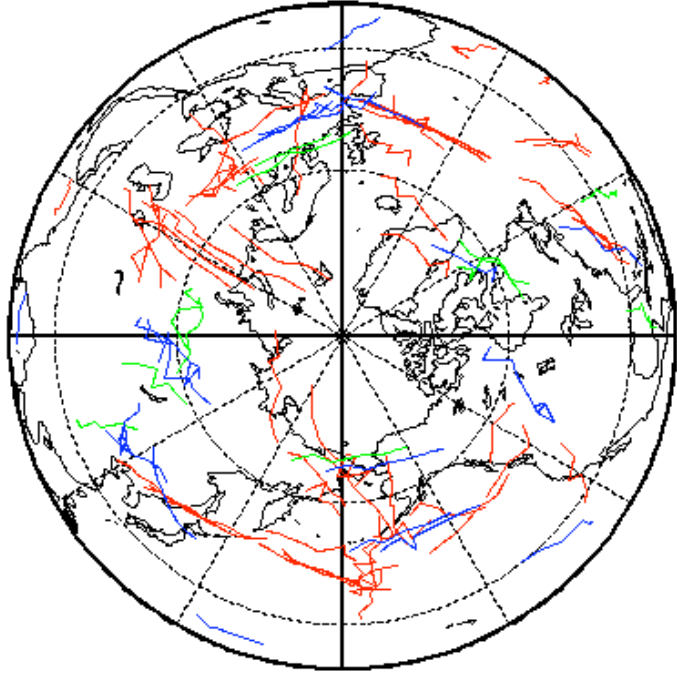
2. Find points in a rectangle like this:



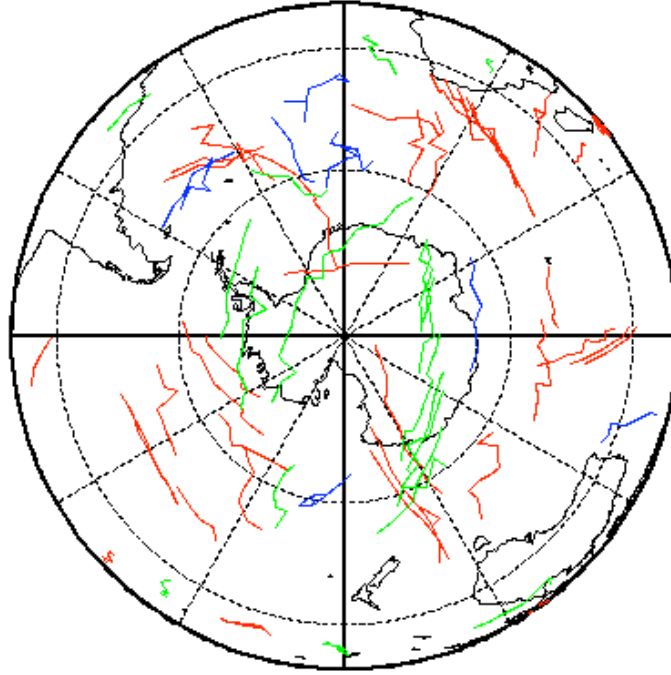
3. If found, add & extend the rectangle like this:
(back to 2)



NH

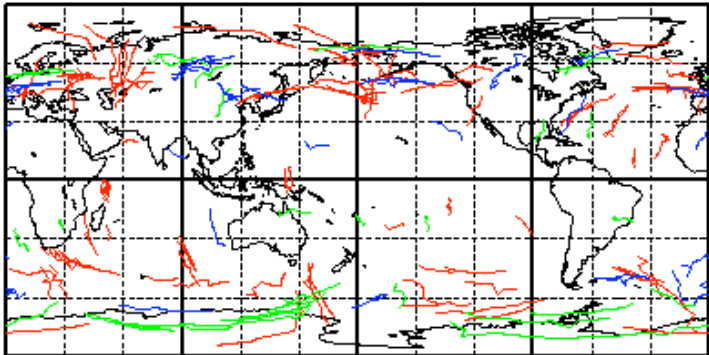


SH

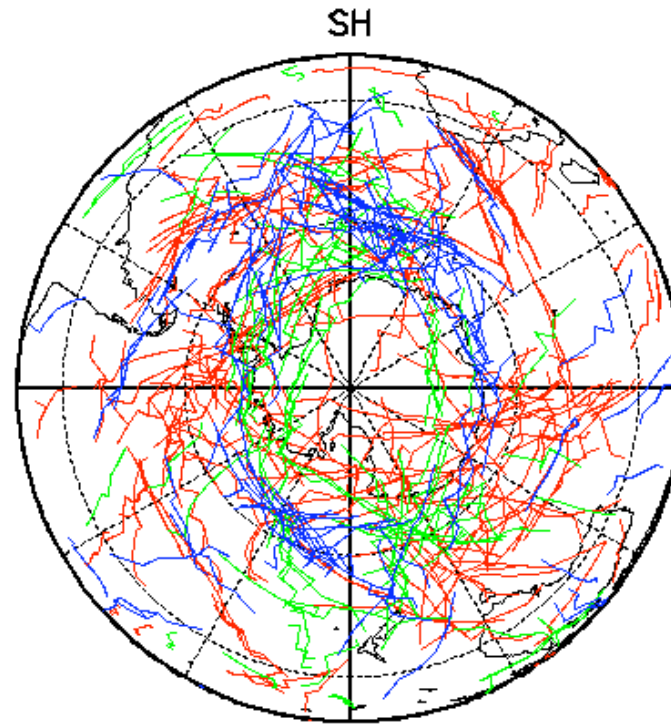
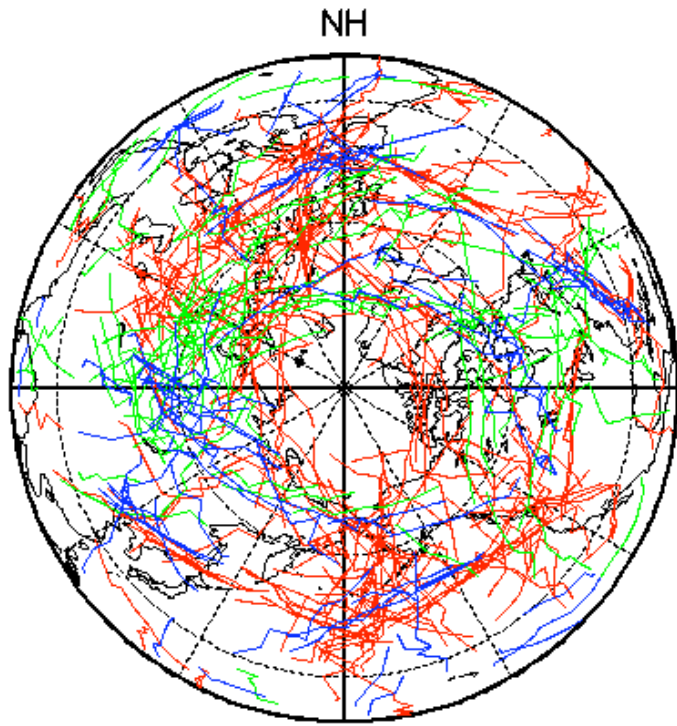


Extracted lines
 $dr = dw = 600$ km

Global



2006-12
2007-1
2007-2

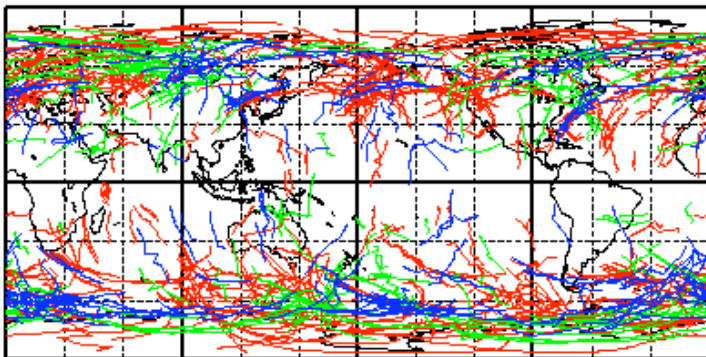


Extracted lines

$dr = 1000$ km

$dw = 600$ km

Global



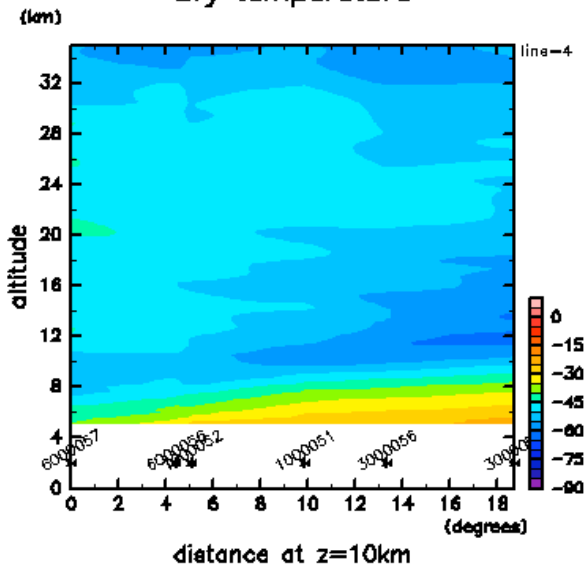
2006-12

2007-1

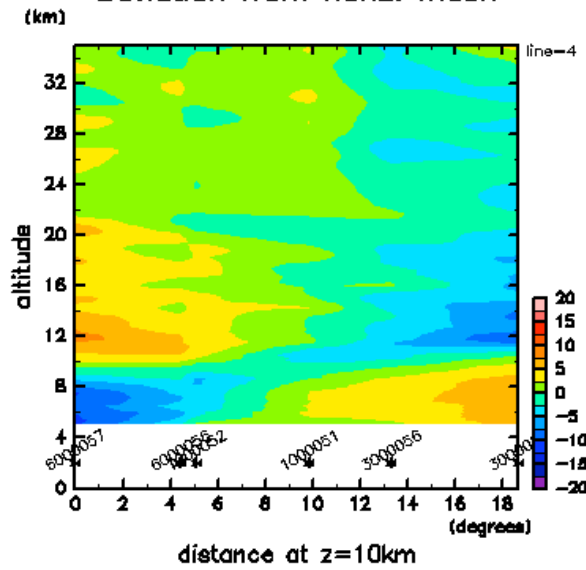
2007-2

- ✓ Dense in mid to high latitudes
(Rarely crossing the equator)
- ✓ Hemispheric, preferred longitudes moving eastward slowly, with wave number=2

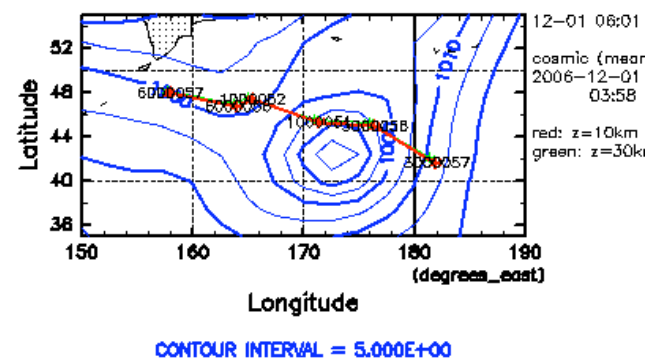
Dry temperature



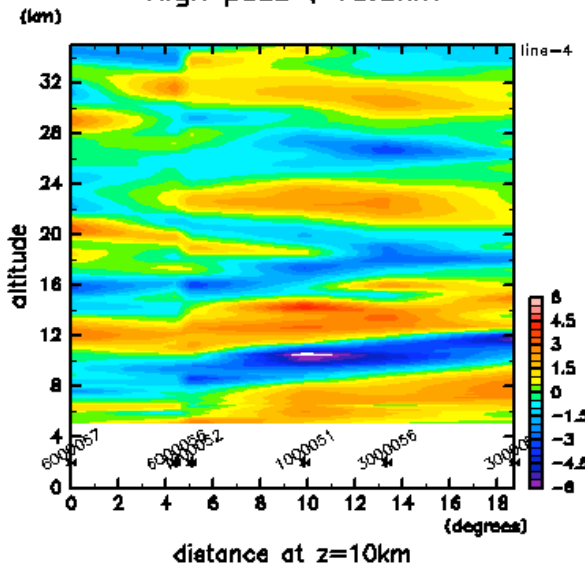
Deviation from horiz. mean



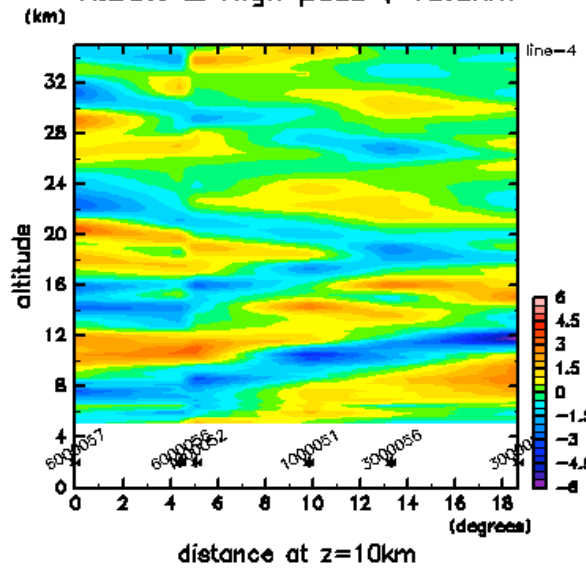
Line 4 & NCEP MSLP



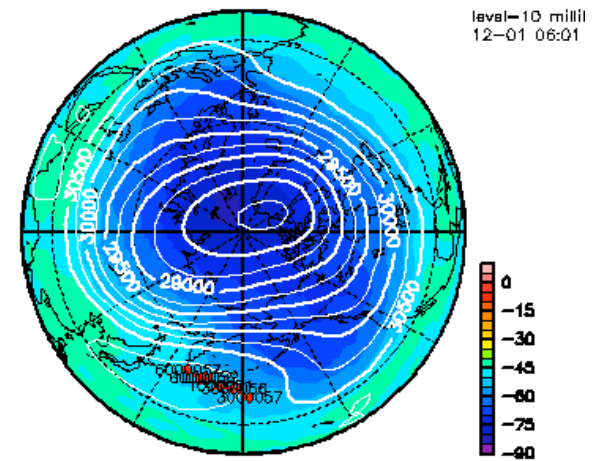
High pass < 10.0km



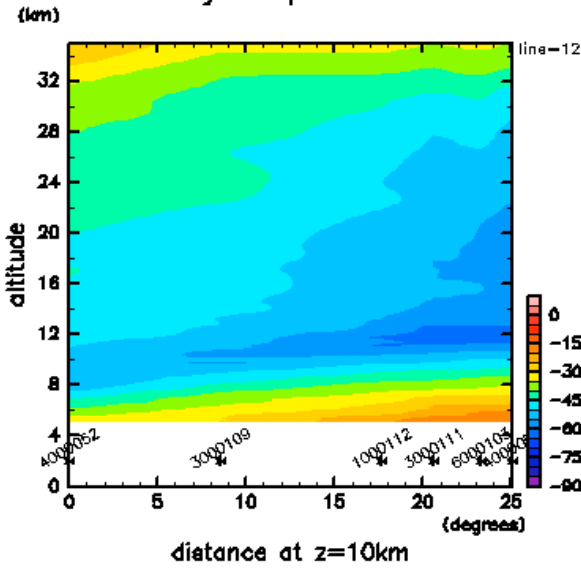
H.Dev. & High pass < 10.0km



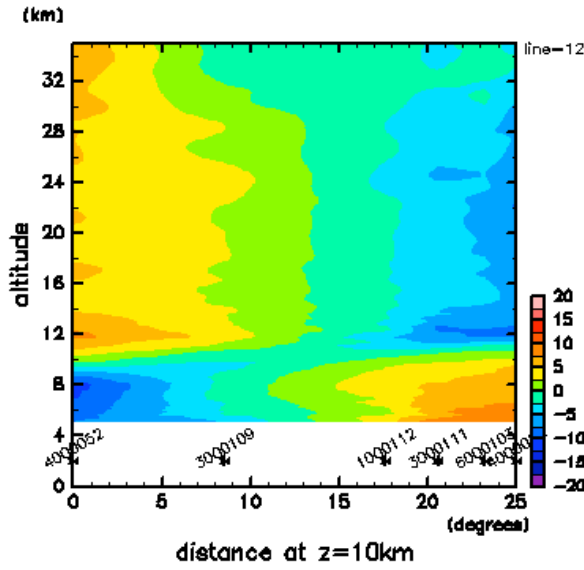
Line 4 & NCEP T&Z



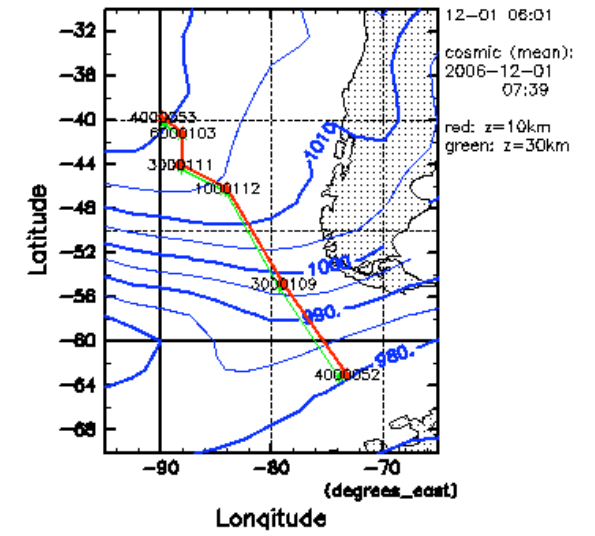
Dry temperature



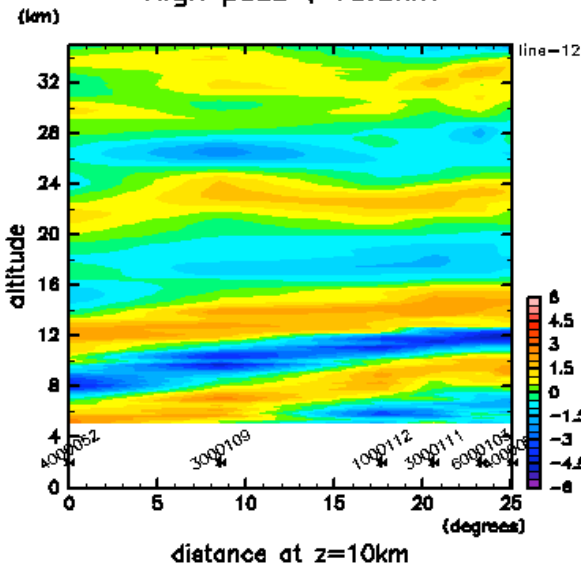
Deviation from hariz. mean



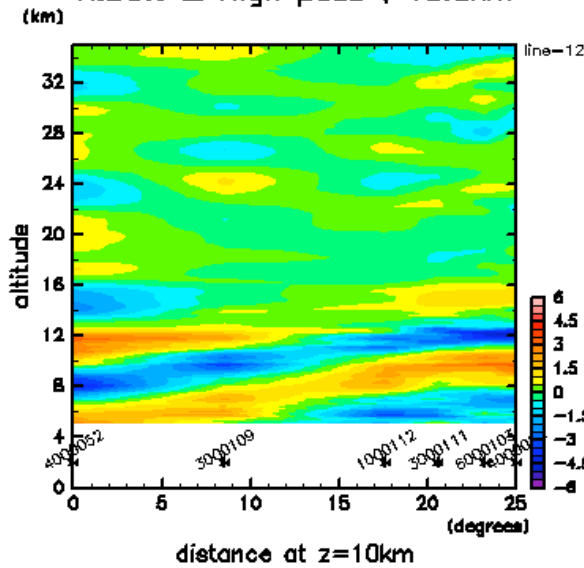
Line 12 & NCEP MSLP



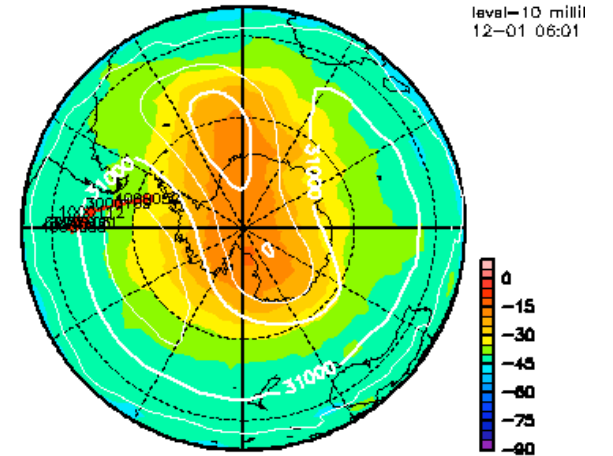
High pass < 10.0km

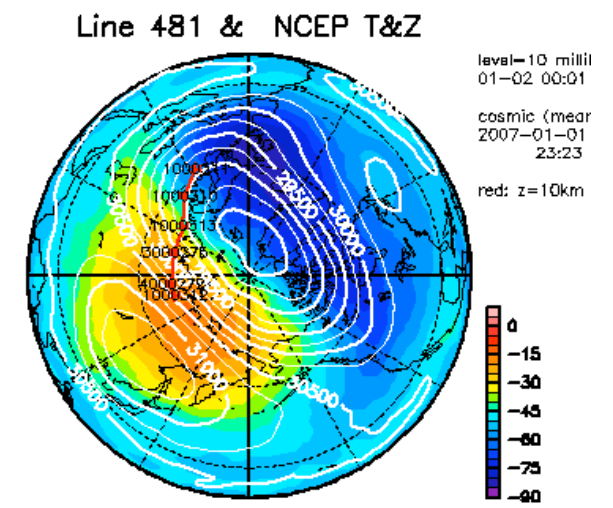
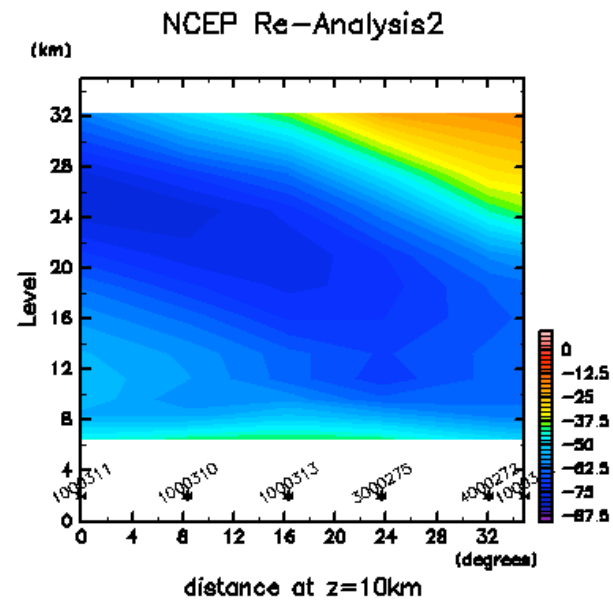
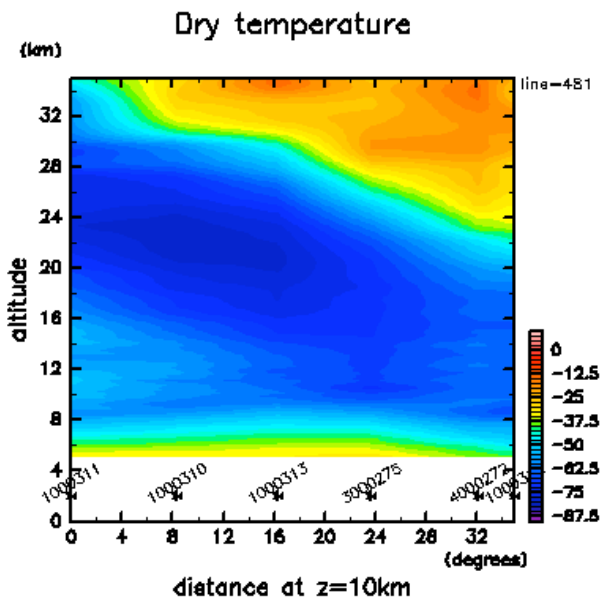


H.Dev. & High pass < 10.0km

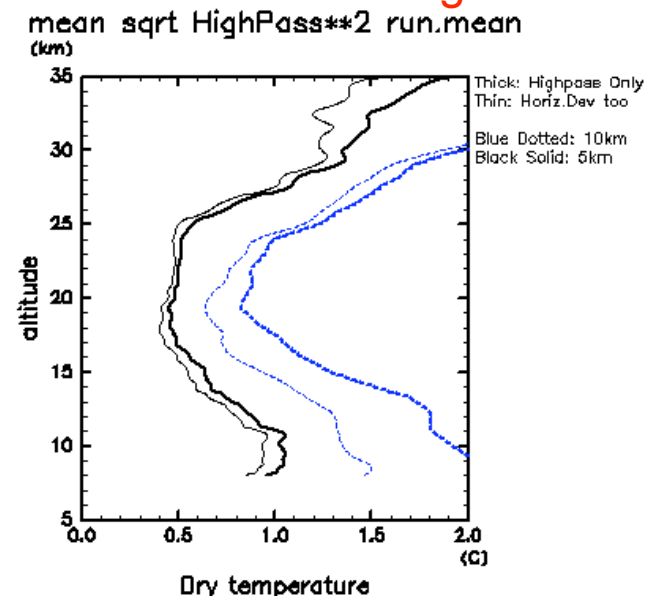
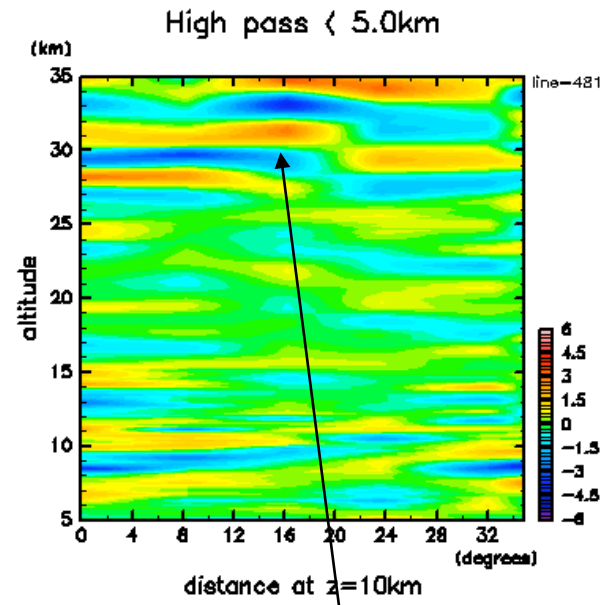
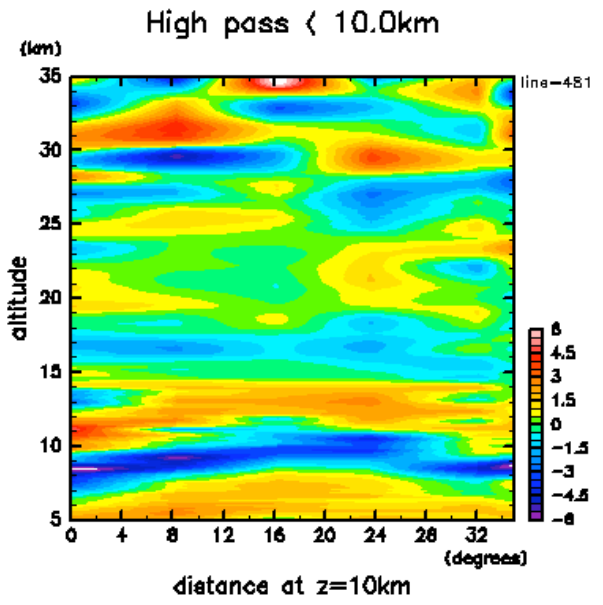


Line 12 & NCEP T&Z

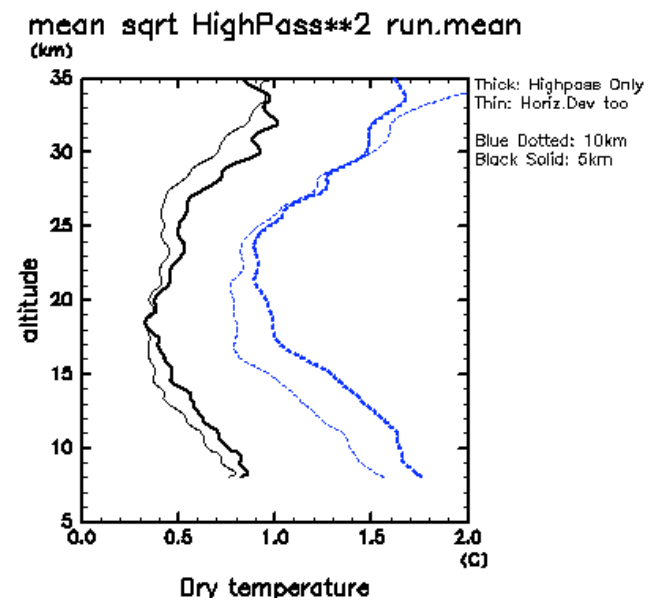
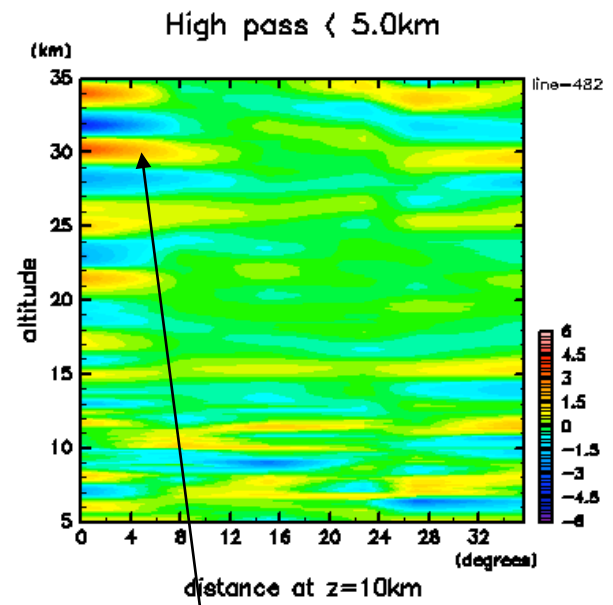
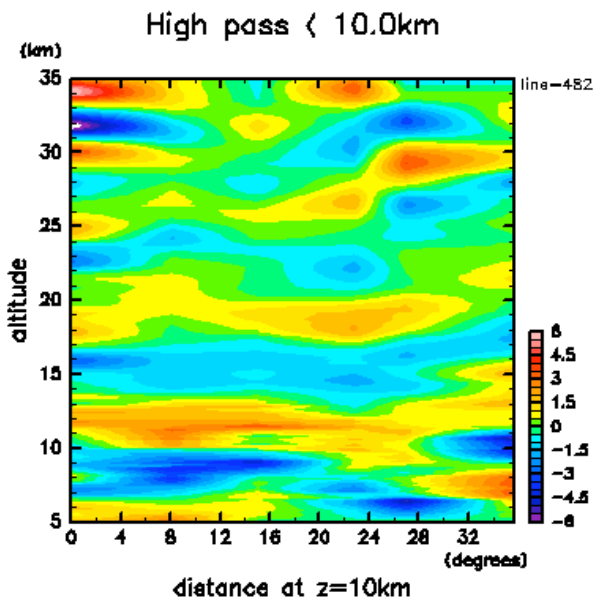
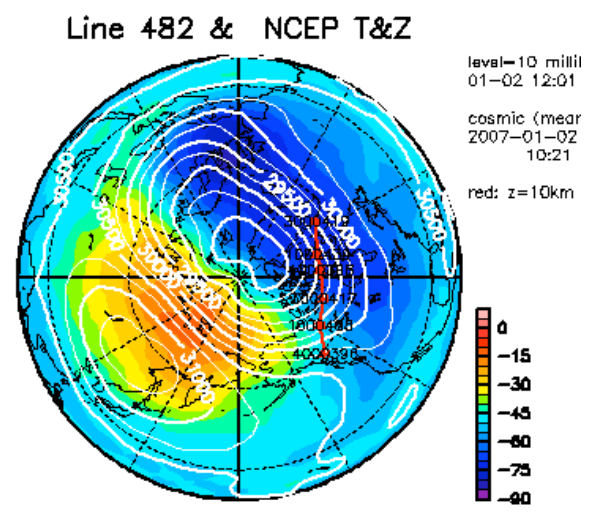
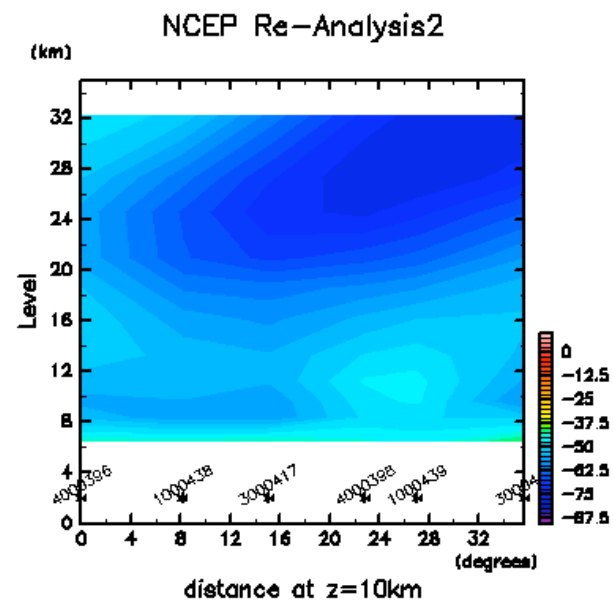
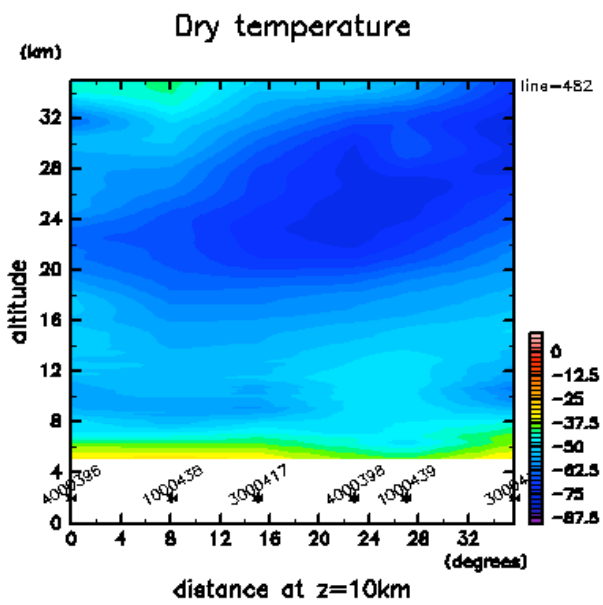




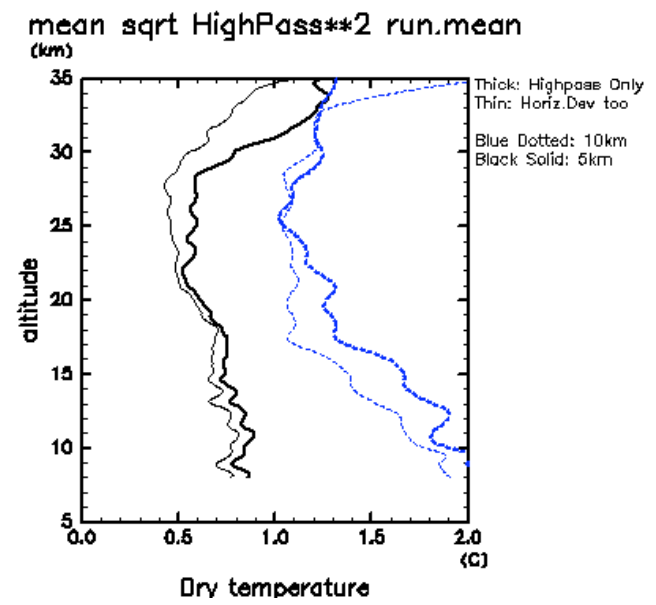
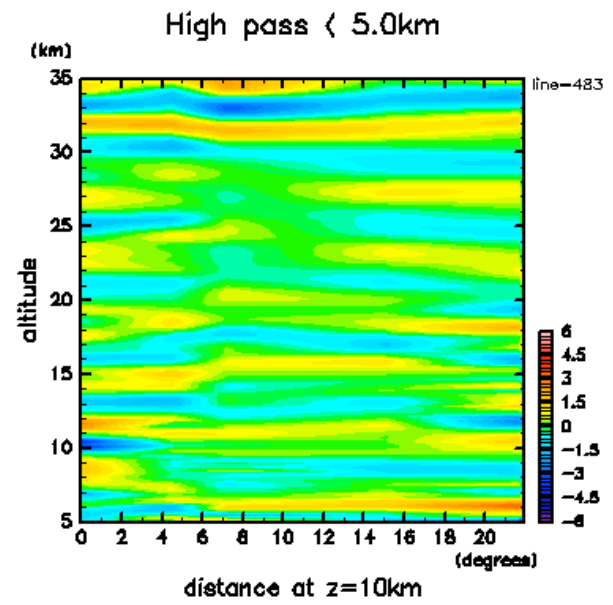
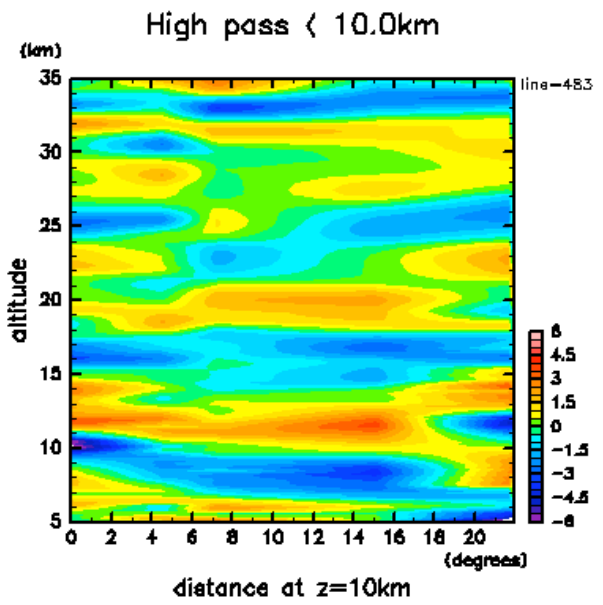
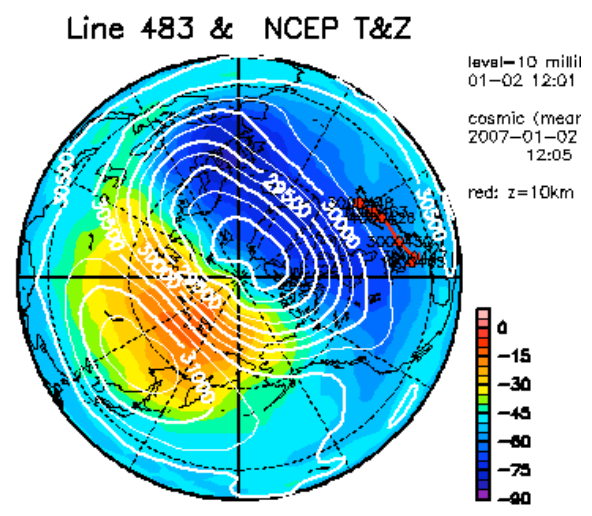
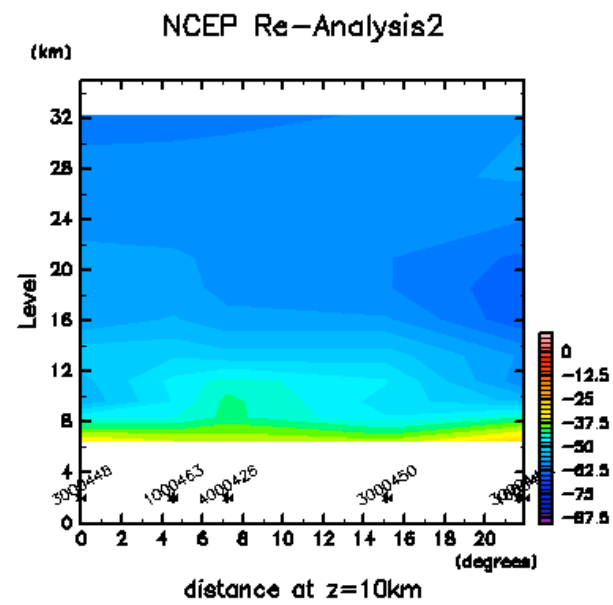
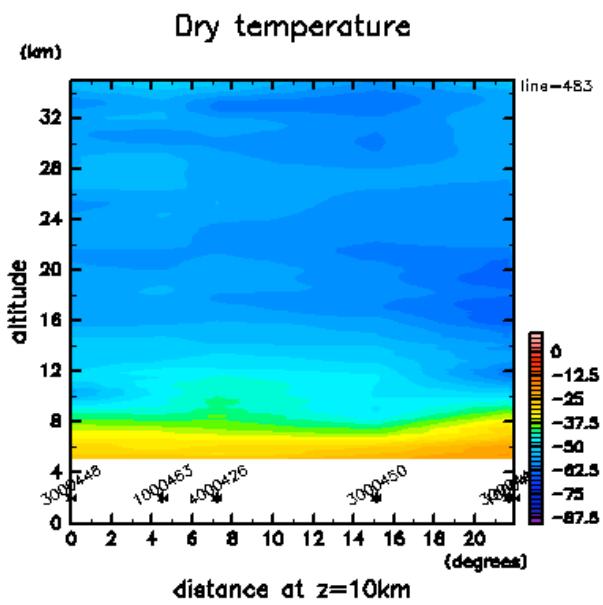
Now in stratospheric sudden warming

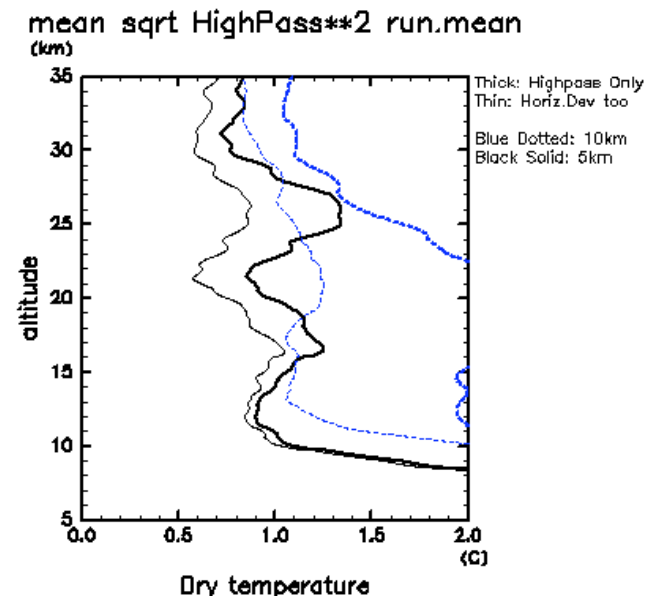
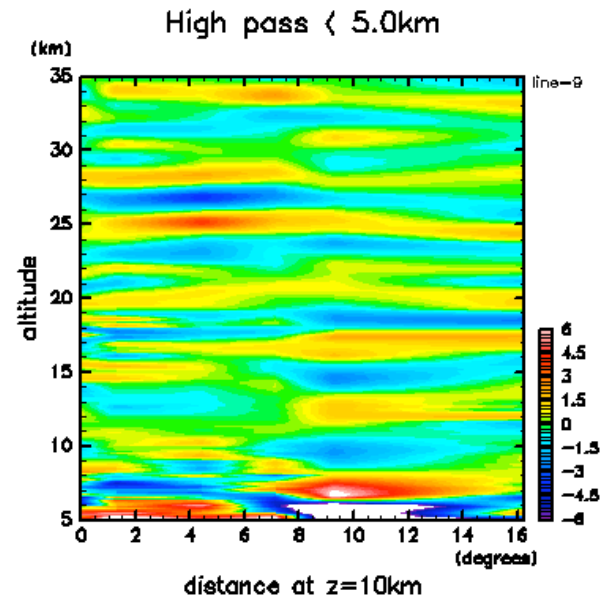
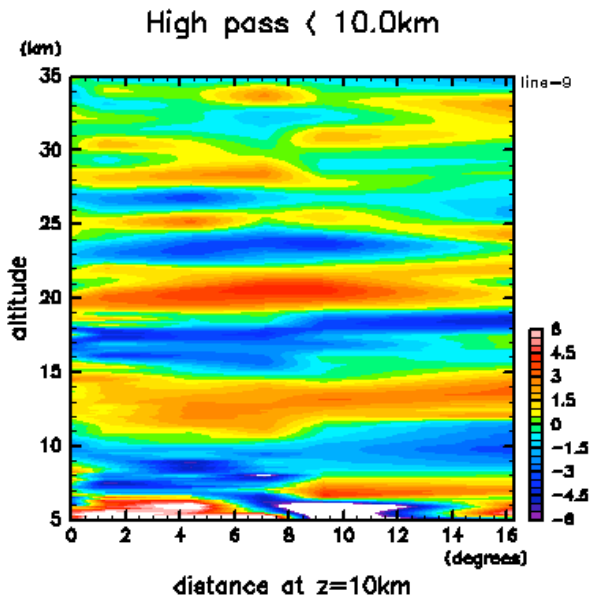
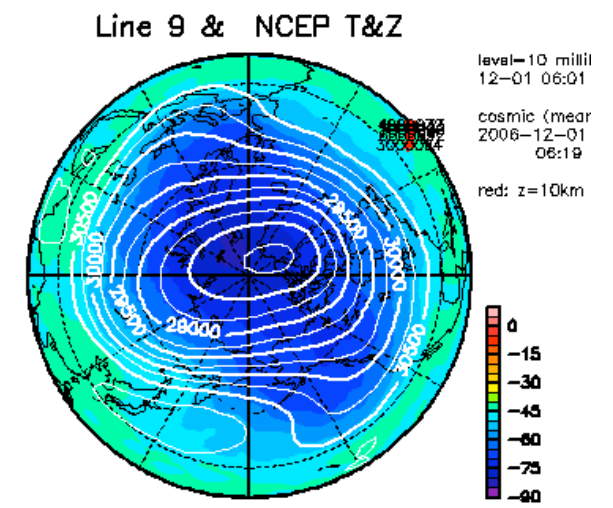
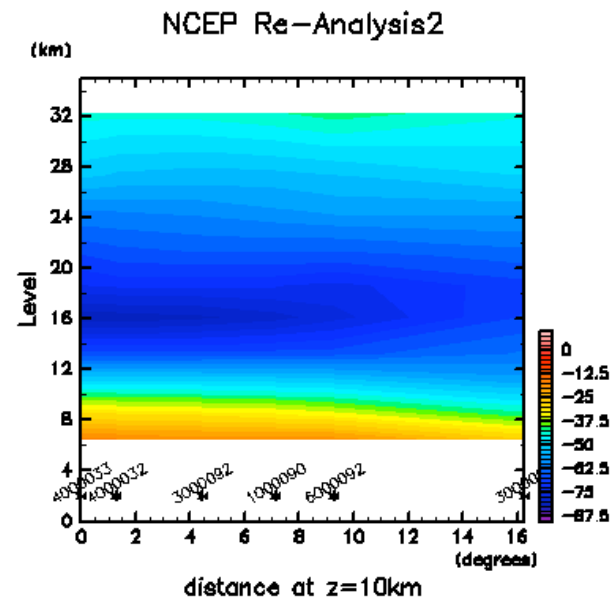
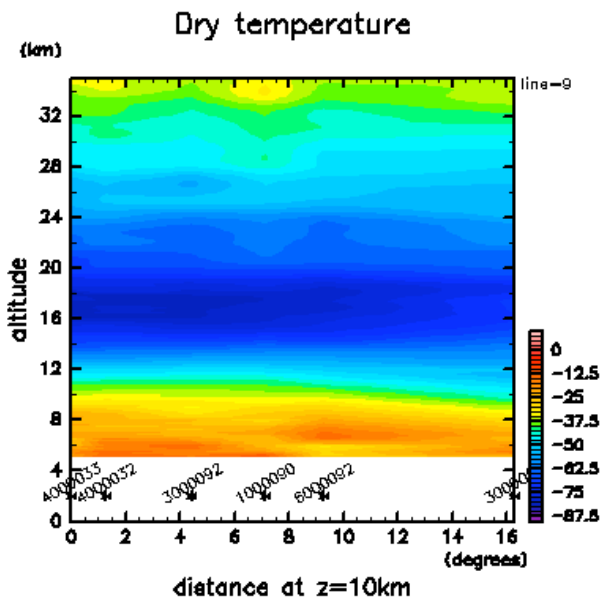


near the polar vortex edge



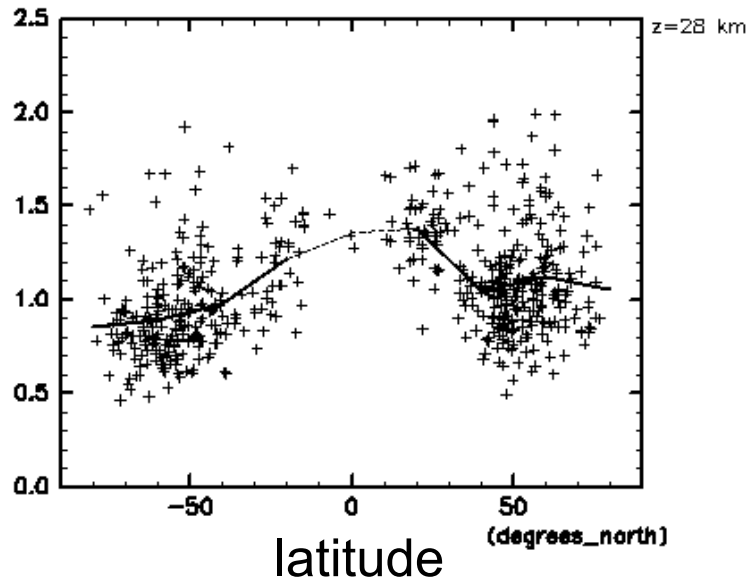
near the polar vortex edge



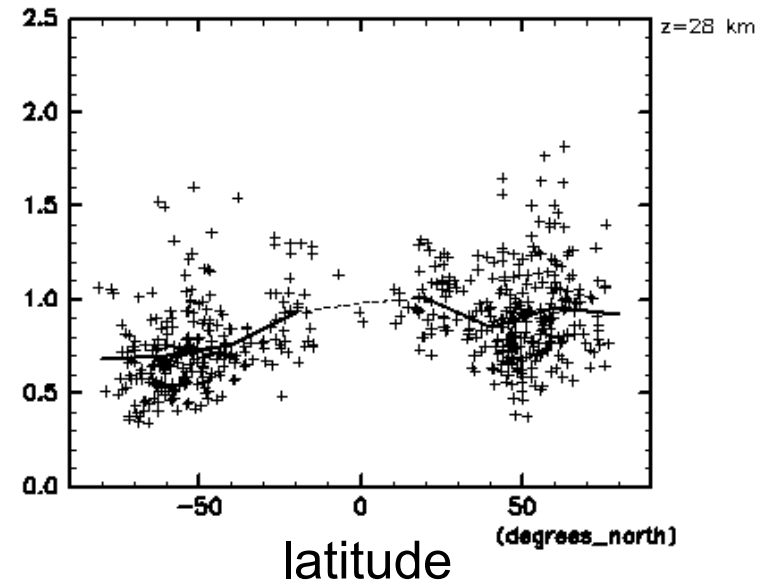


Amplitudes at 28km using lines with lengths 1500 ~ 3000 km

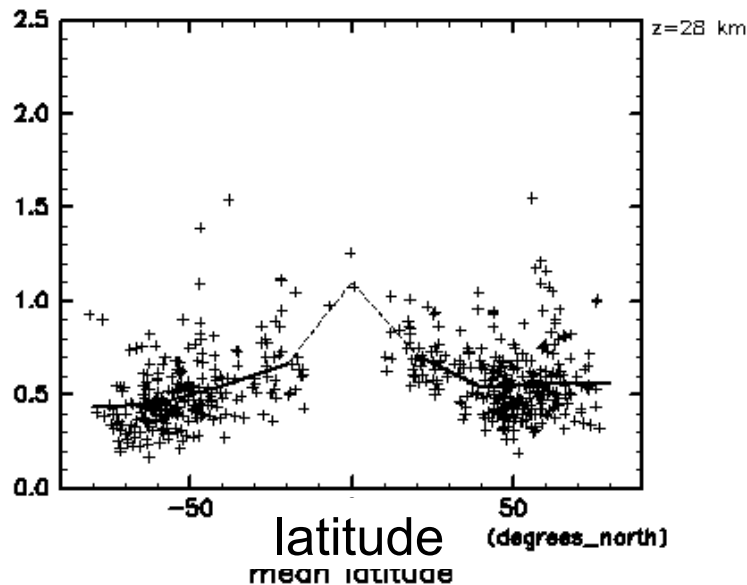
(a) Amplitude (cutoff 10km)



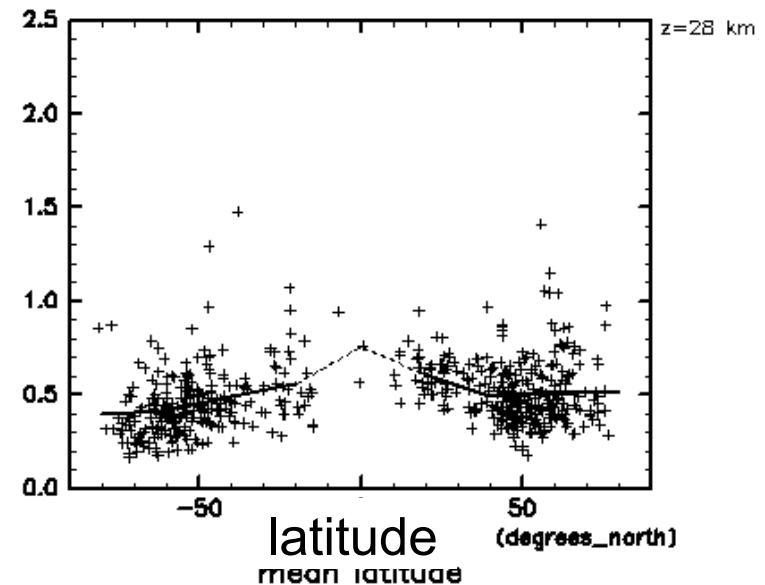
(b) Amplitude (horiz.dev & cutoff 10km)



(c) Amplitude (cutoff 5km)



(d) Amplitude (horiz.dev & cutoff 5km)



Preusse 2006

- Mean horizontal wavelength estimated for Aug 1997 & Aug 2003 from CRISTA data (infrared limb sounder) using adjacent profile pairs

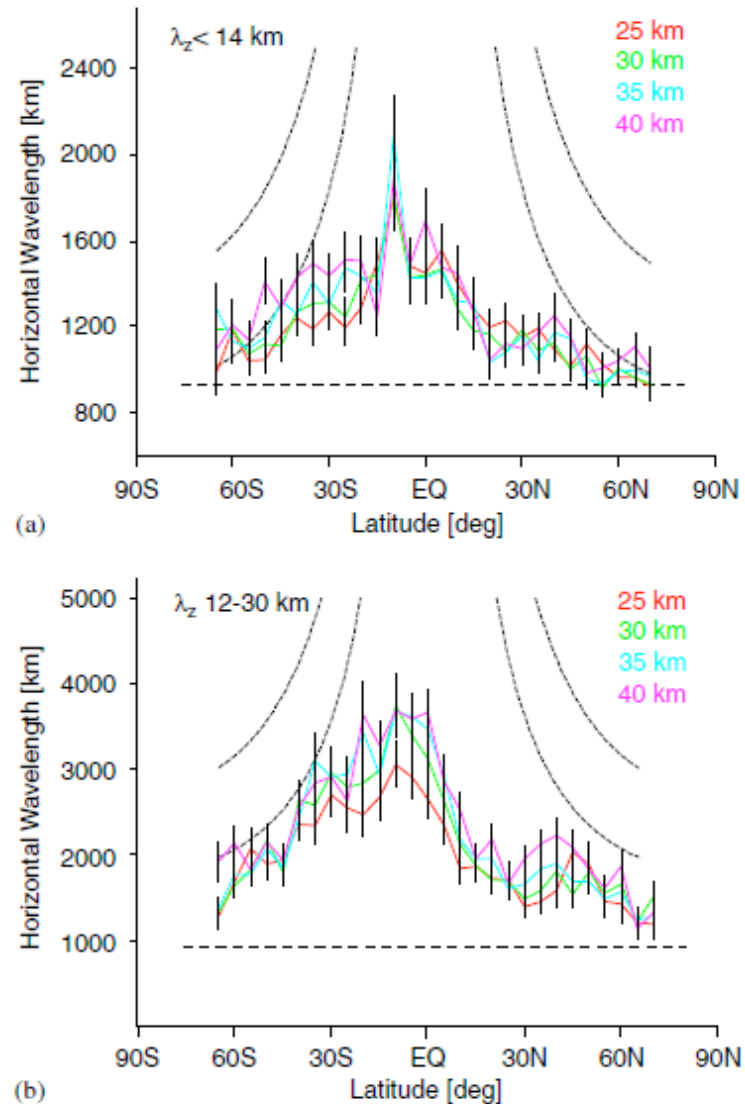
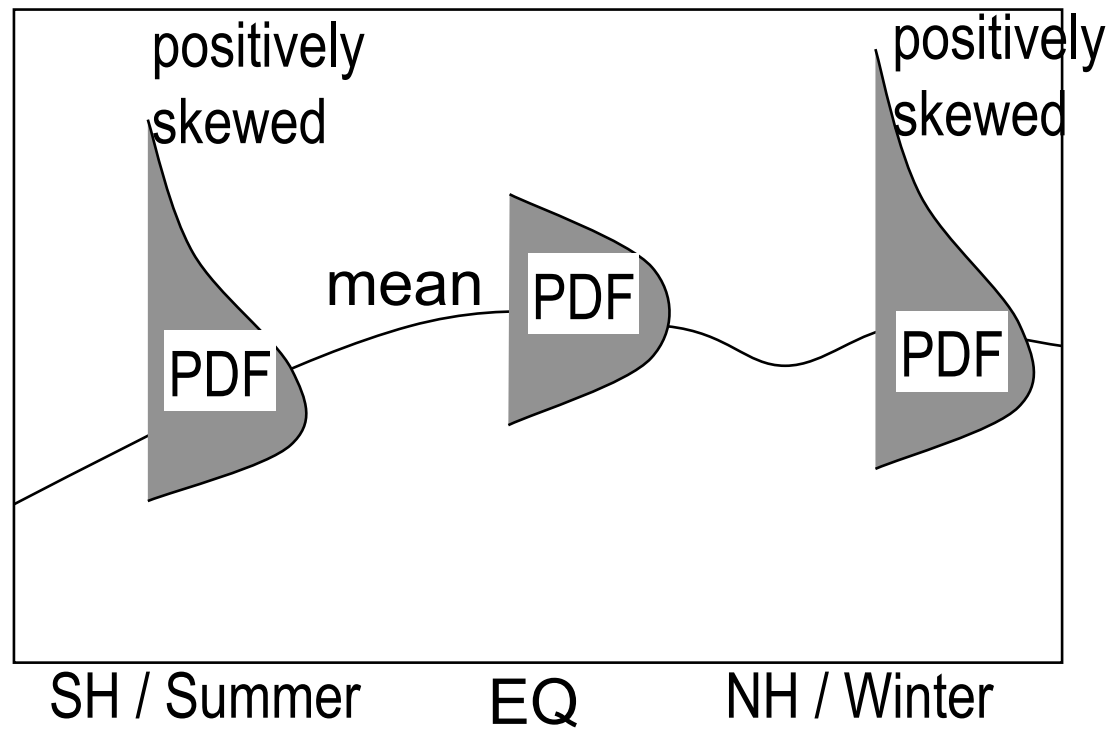
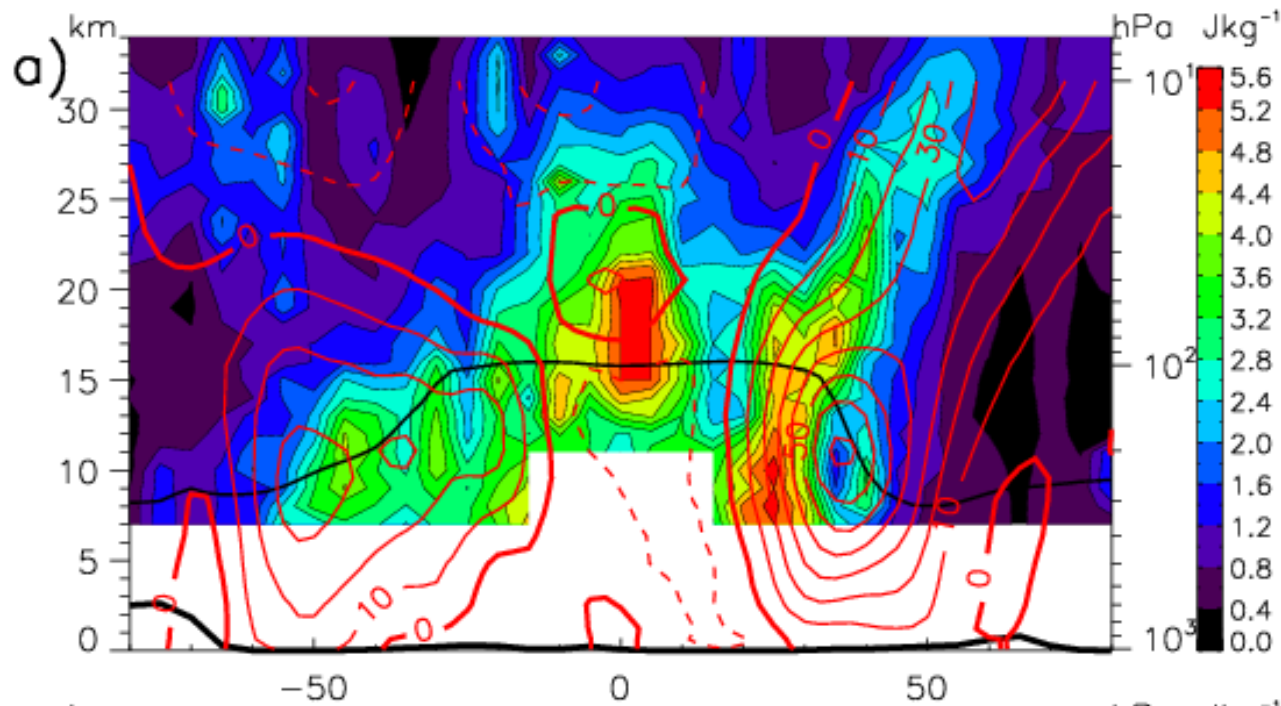


Fig. 3. Zonal mean horizontal wavelength $\bar{\lambda}_h = 2\pi/\bar{k}$ measured by CRISTA-2 (color code gives altitude), horizontal wavelength ensemble average lower limit (horizontal line), and upper limit for two values of constant ω/f (black dashed). The upper panel shows short vertical wavelengths, the lower panel shows long vertical wavelengths. For details see text.

Features of GW amplitudes in T (cutoff: 10 km)



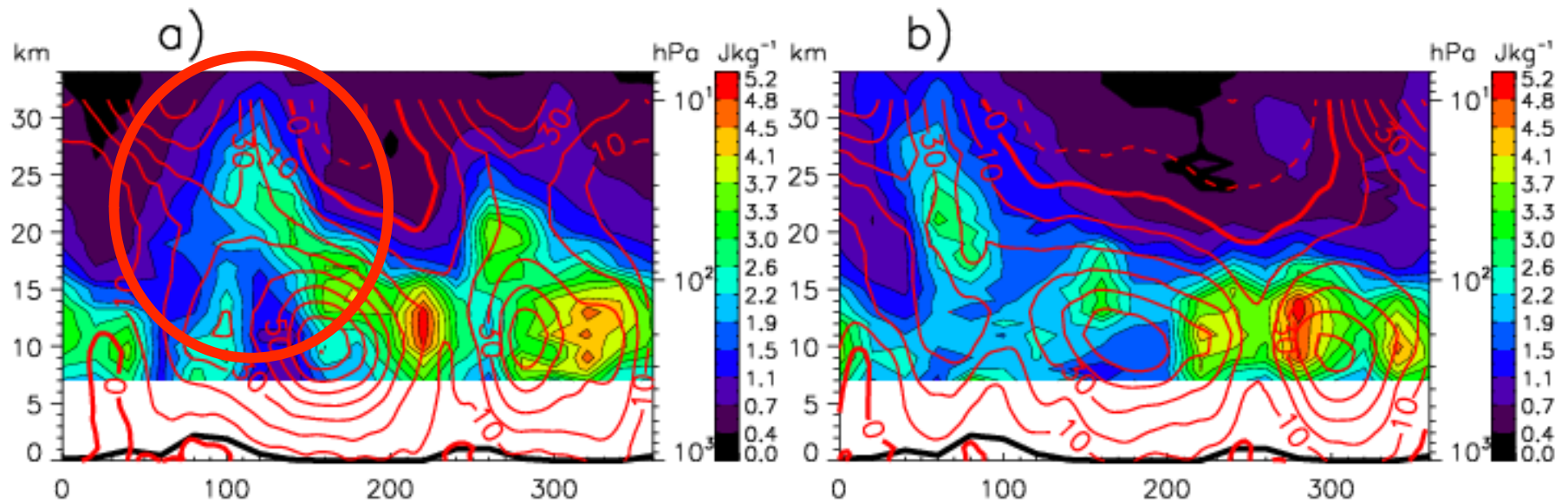
S. Alexander et al 2008 (again)



GW energy averaged between 130E-150E
23-28 Dec 2006 using COSMIC data

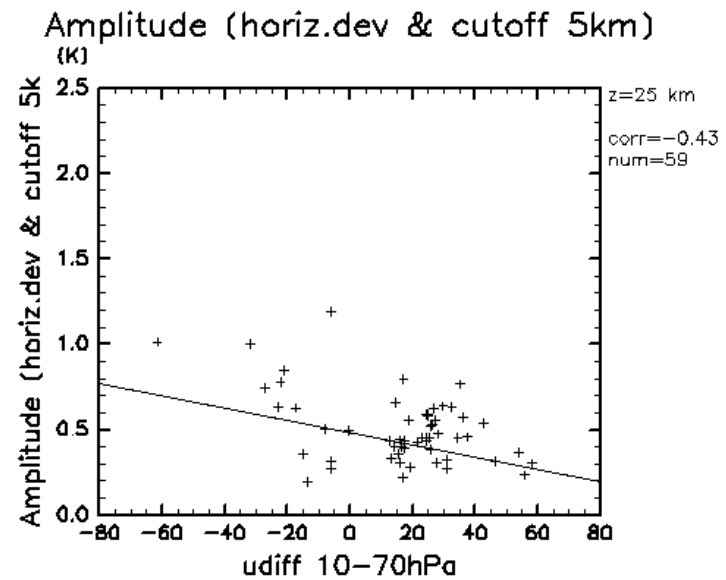
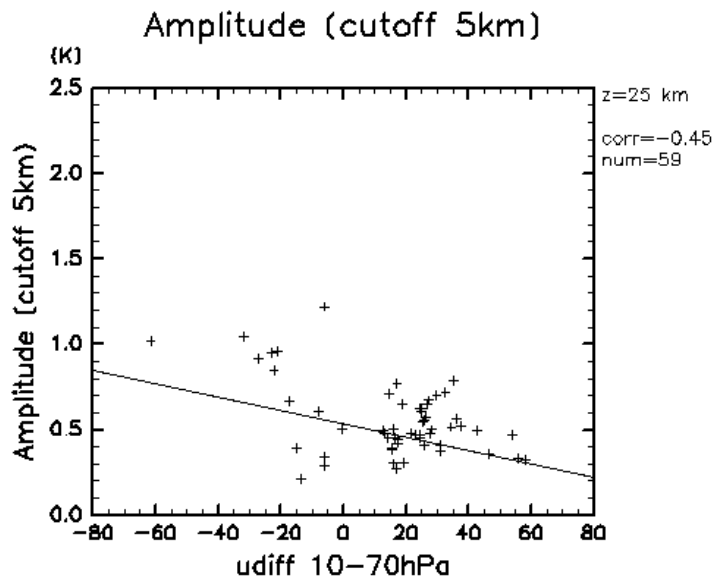
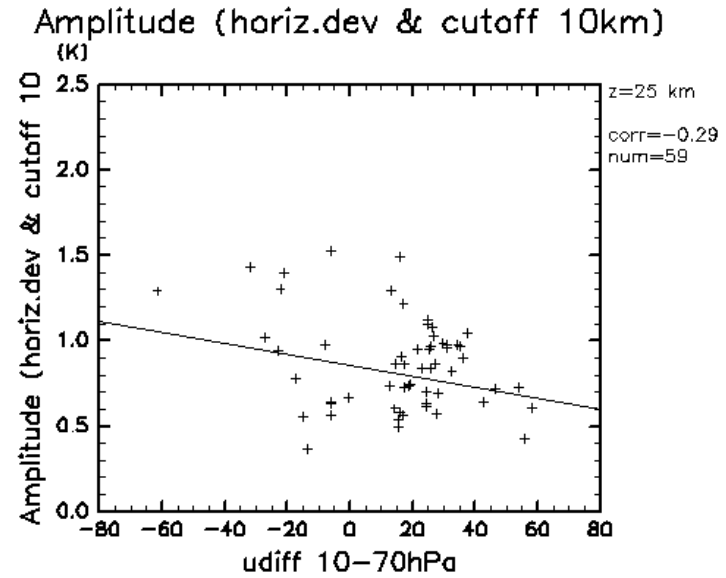
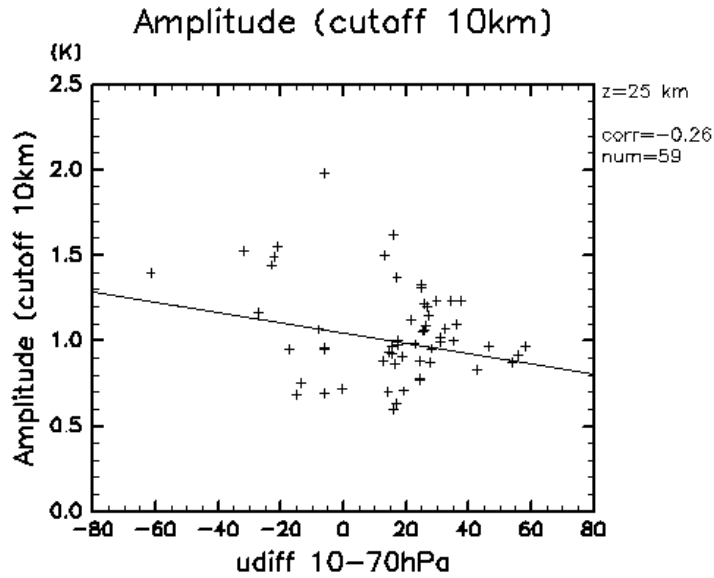


Alexander et al 2008

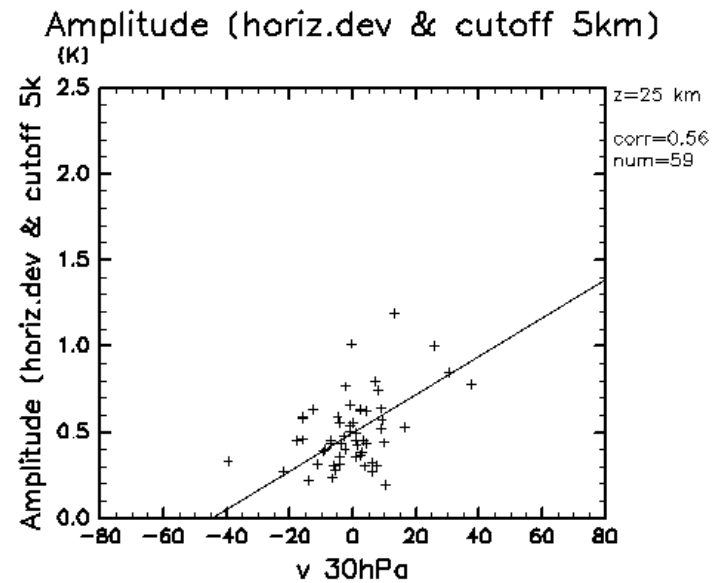
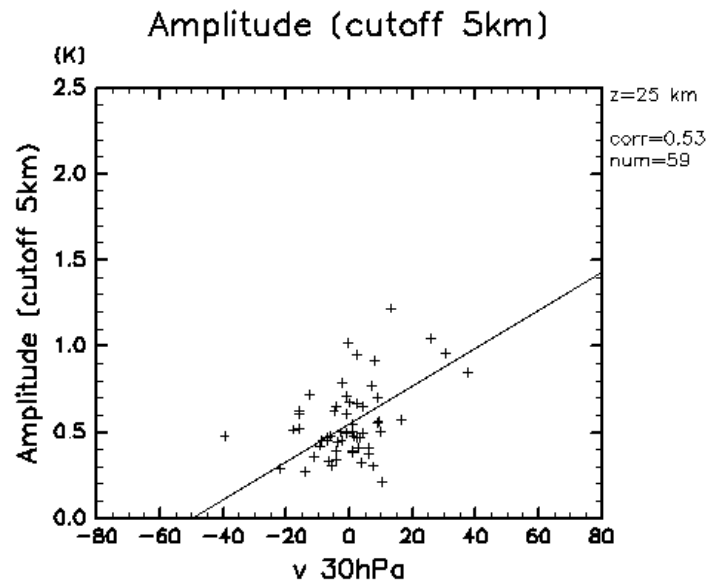
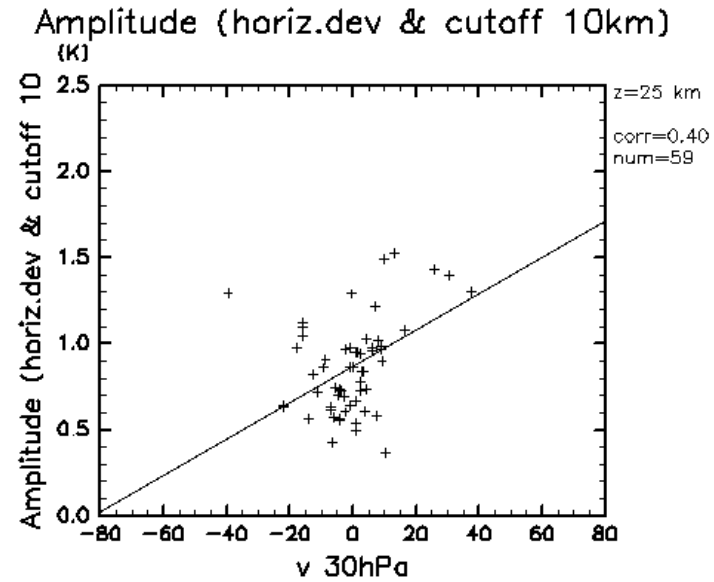
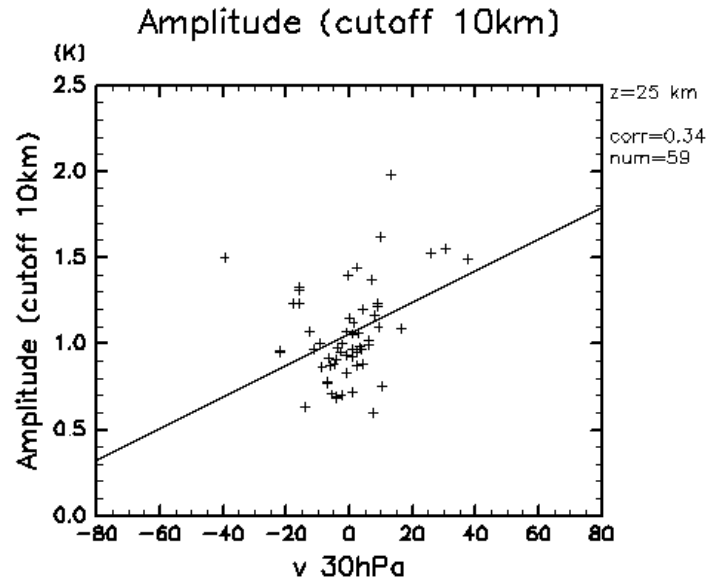


GW energy averaged at 40N using COSMIC data
(a) 29 Nov-5 Dec 2006 (b) 24-30 Dec 2006

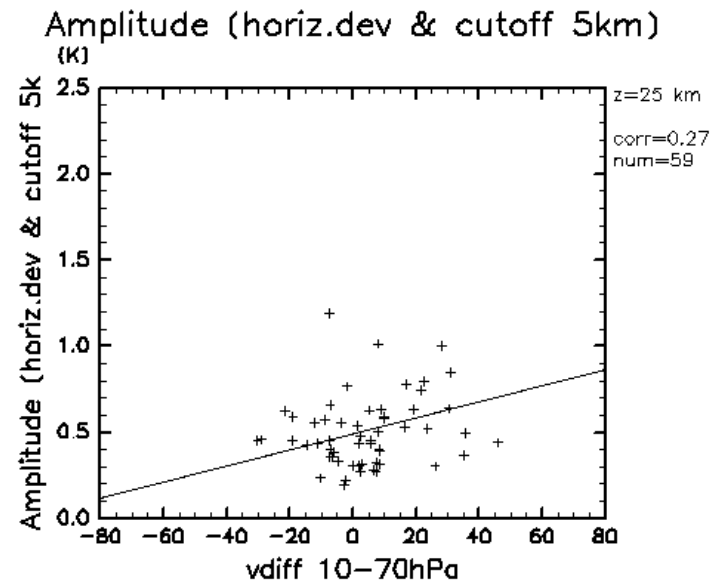
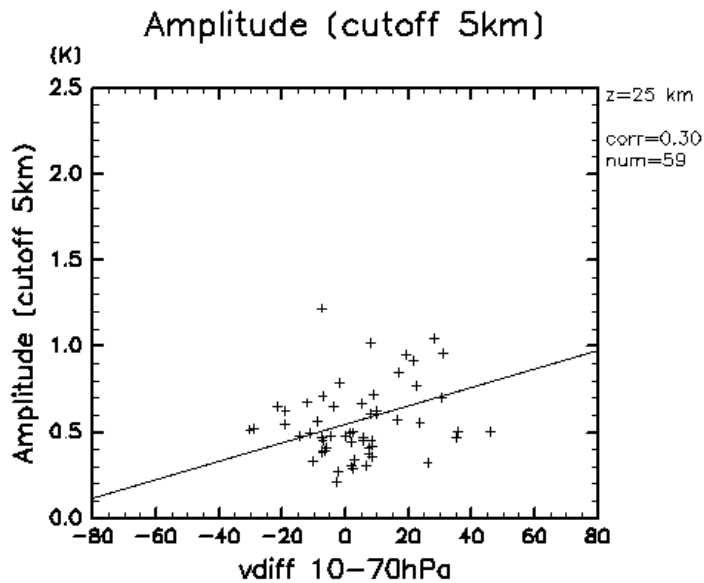
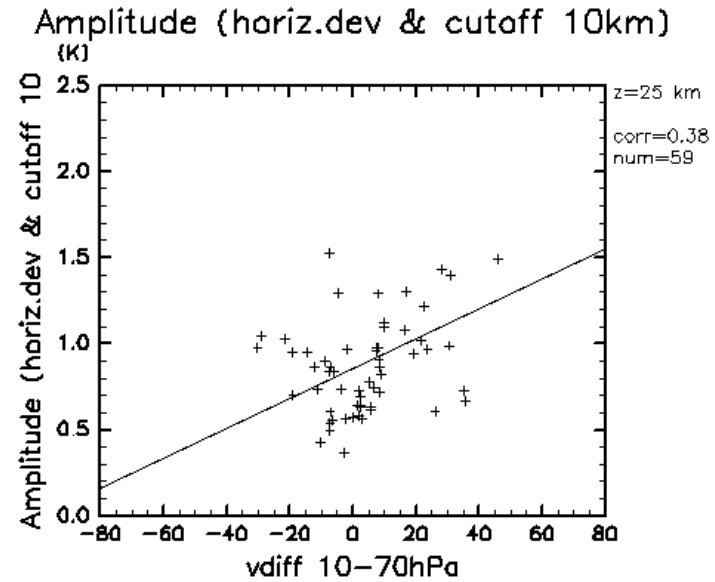
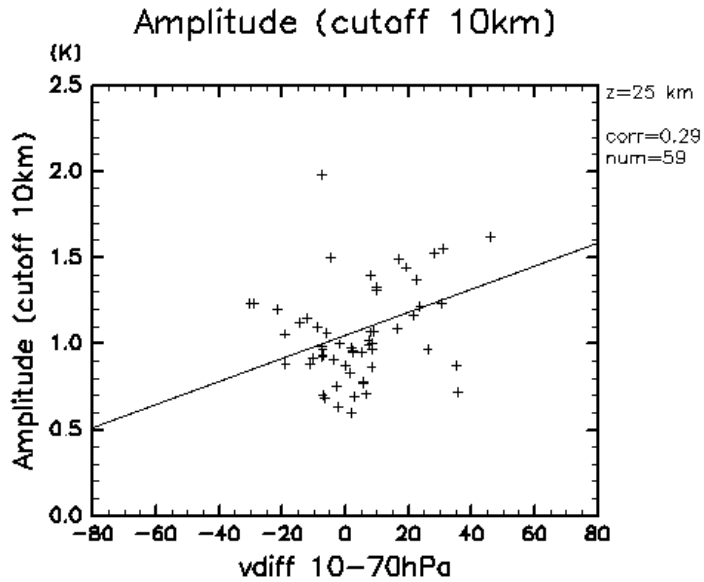
T amp (25km) vs U shear (10 – 70hPa) 40-65N



T amp (25km) vs V (30hPa) 40-65N

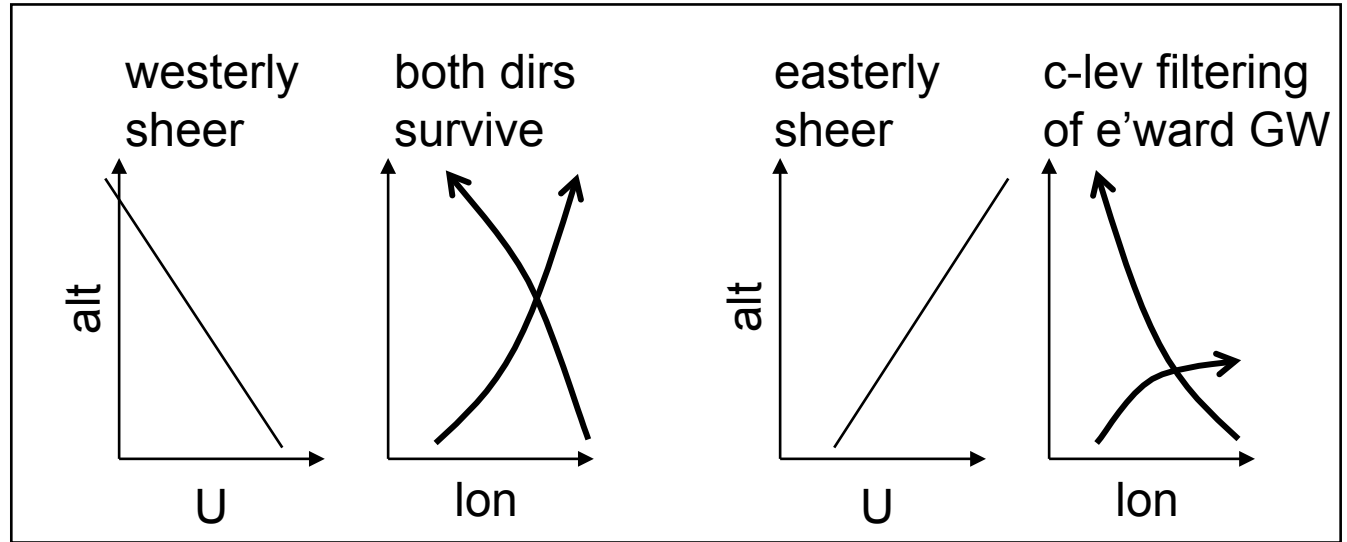


T amp (25km) vs V shear (10 – 70hPa) 40-65N

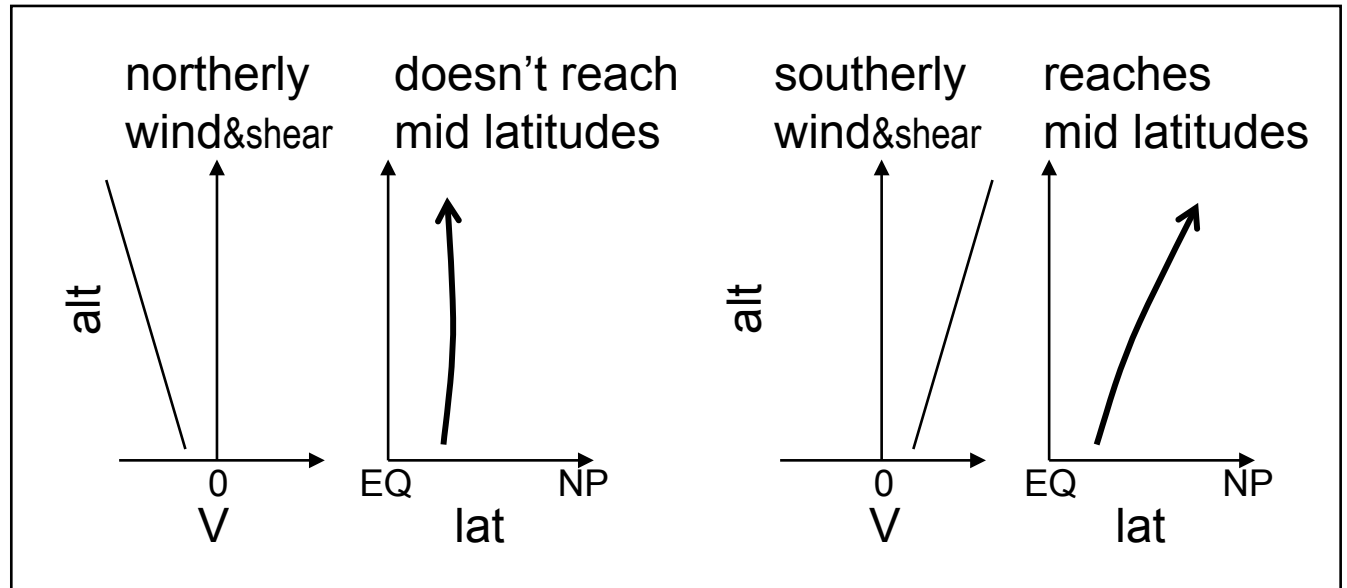


GW propag directions and mean winds

Zonal propagation



Meridional propagation

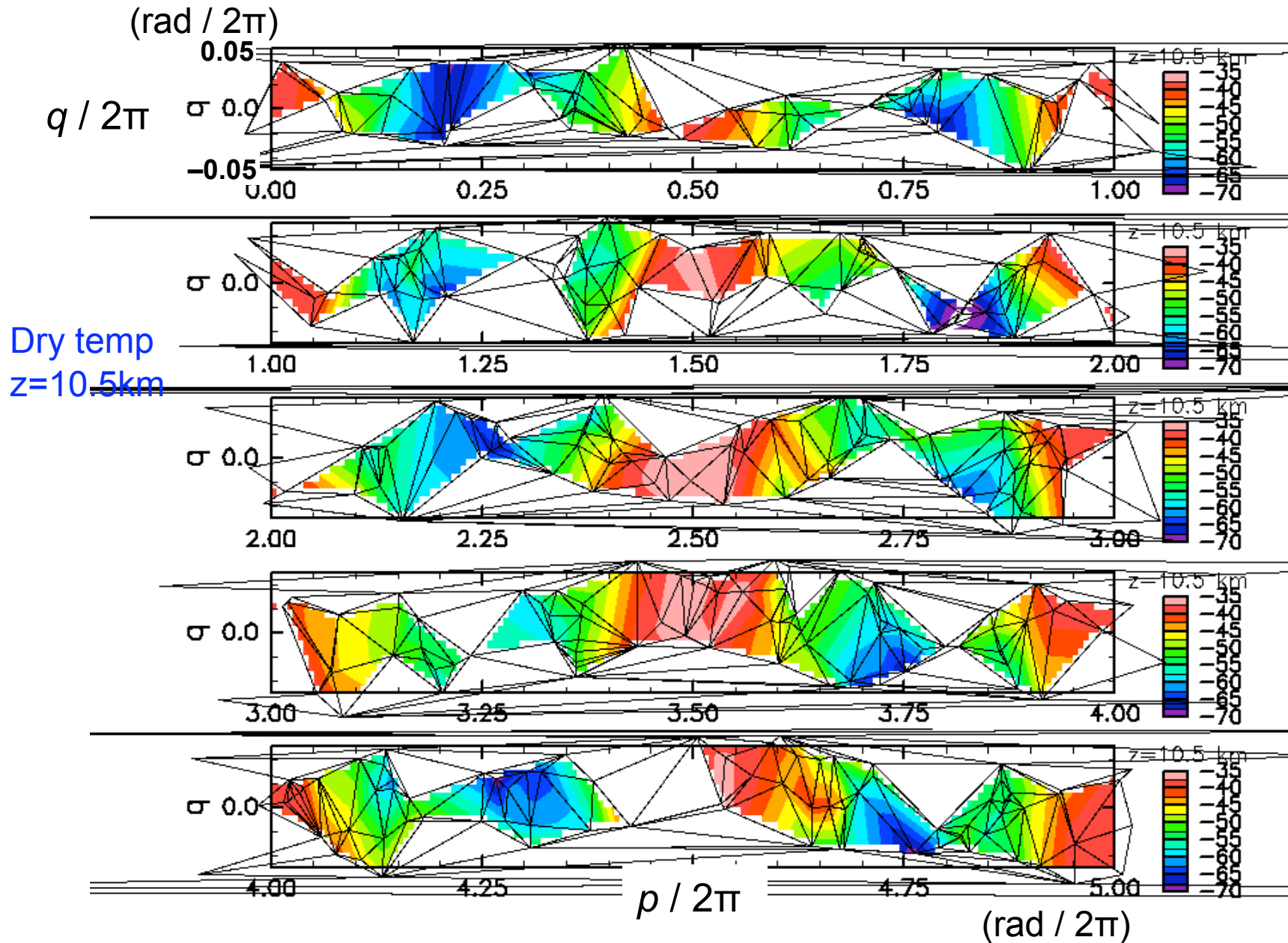


Summary

- Snapshot “lines” were extracted from COSMIC RO data
 - Found: GPS RO events are frequently form lines (esp. mid-high latitudes)
 - Longitudinally concentrated, which moves eastward
- Vertical cross sections → Spatial structures of GWs
 - In many cases, horizontal structures indicating GWs are identified over multiple RO events
 - Statistics → GWs of $5 < \lambda_z < 10$ km tend to have relatively large λ_H in the equatorial region & SH.
 - GW in NH (winter) mid-to-high lat
 - Propagation: zonally both Eward&Wward, meridionally northward
 - Source: mainly in subtropics
 - GW amp correlated negatively with U shear and positively with V wind; Critical-level filtering for zonal amp fluctuation

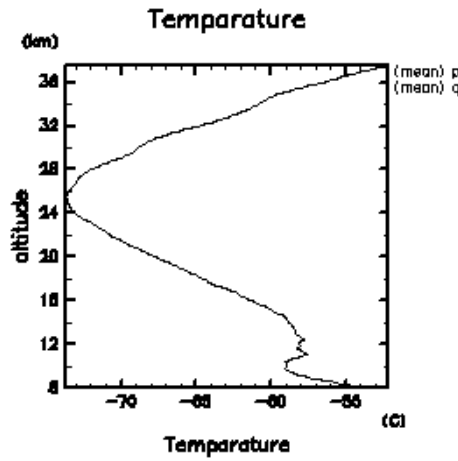
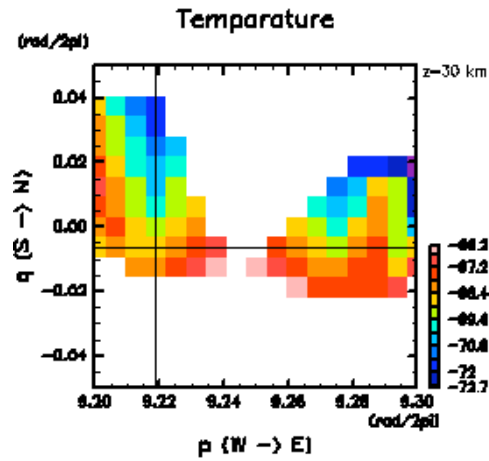
Gridding by triangulation (2006/12/01 4LEOs $\Delta p_{\max} = \Delta q_{\max} = 0.1 \text{ rad} / 2\pi$)

Work done by Y. Azuma, Horinouchi, and Tsuda



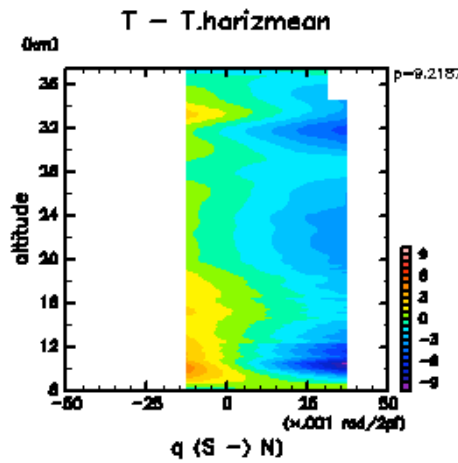
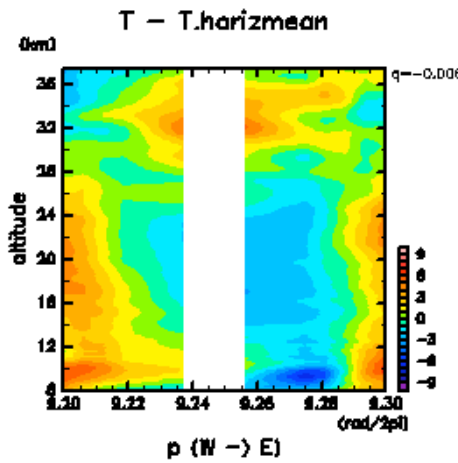
Polar case

Horizontal x-sec
 $z=30\text{km}$
 (2000 km x
 2000 km)



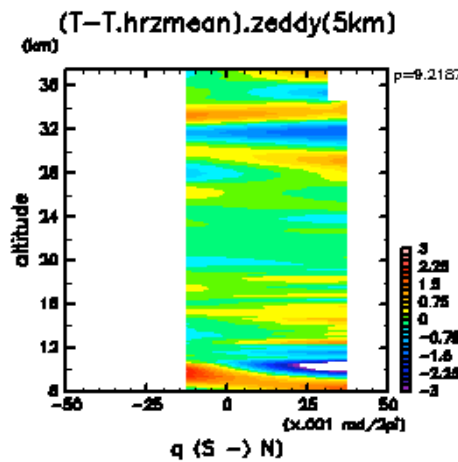
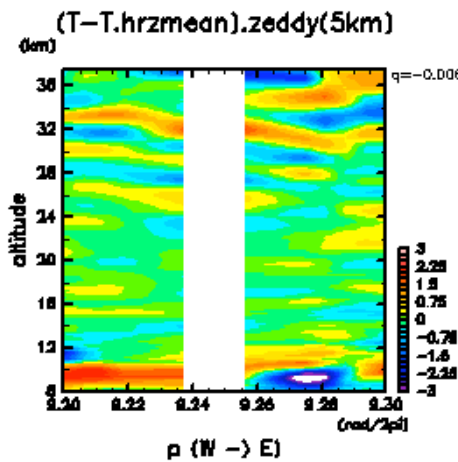
Horizontal mean

x-z sec
 (dev. from horiz.
 mean)



y-z sec
 (dev. from horiz.
 mean)

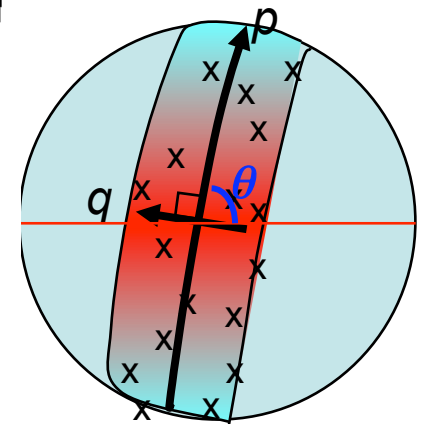
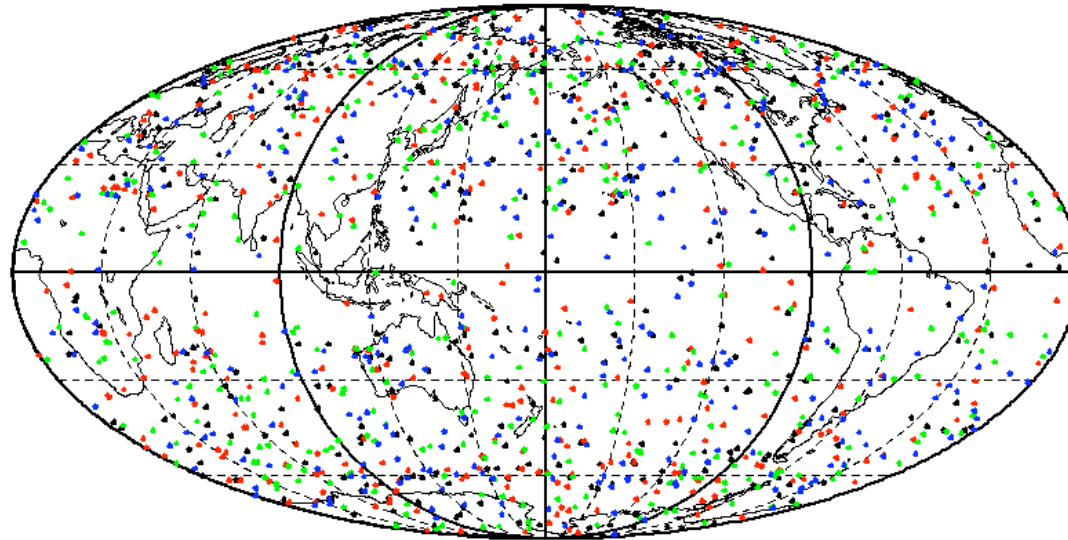
z high pass
 (< 5km)



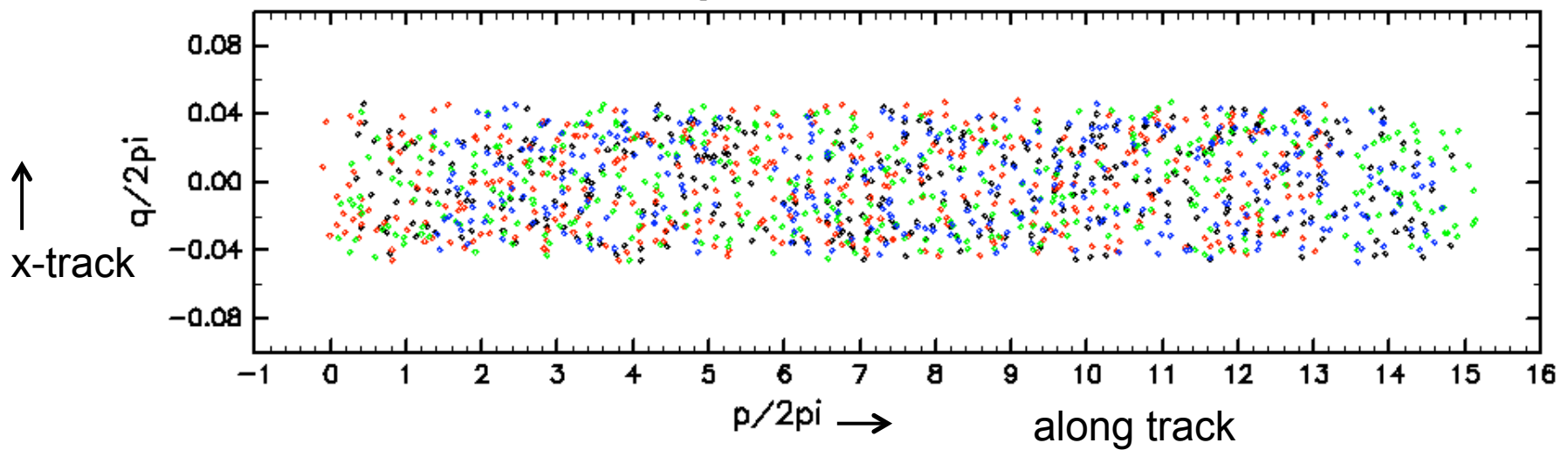
z high pass
 (< 5km)

Sample transformation results (Nov 16, 2006 leo:1,3,4,6)

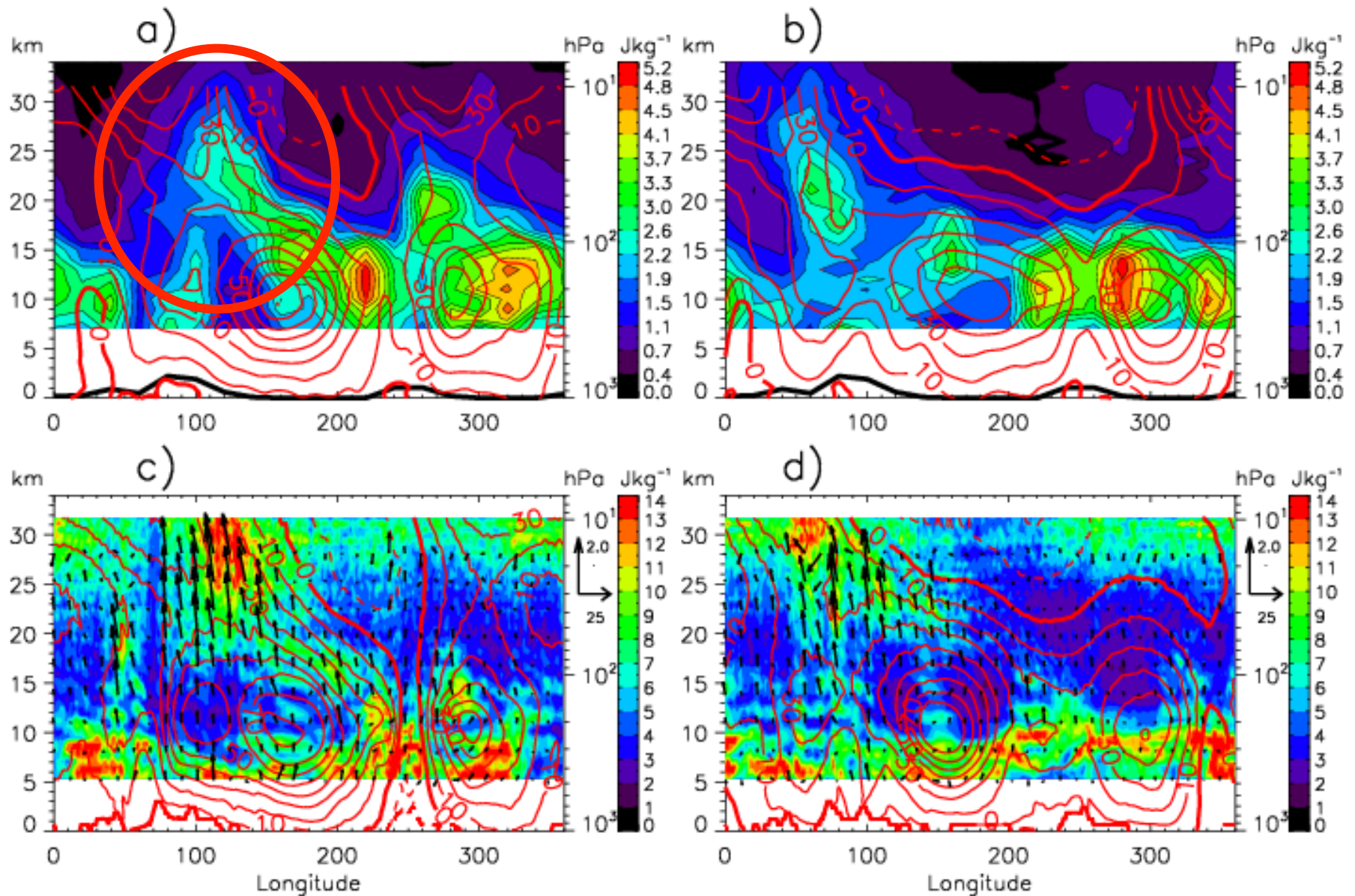
day:2006.320 leo ids:1346

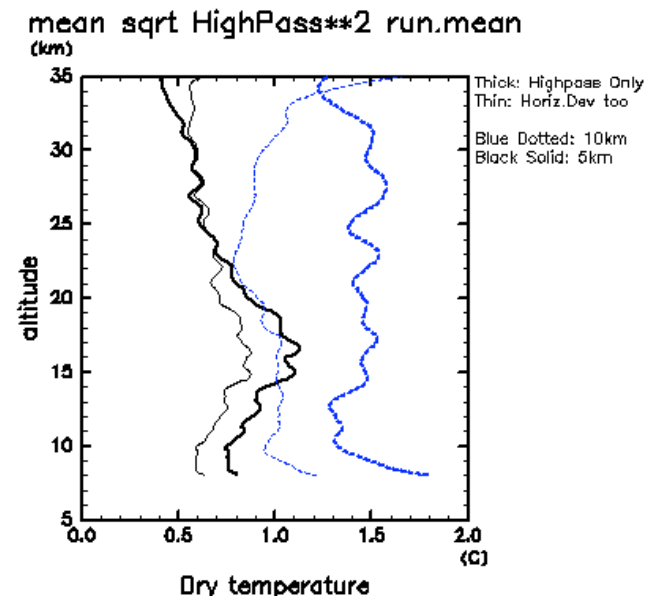
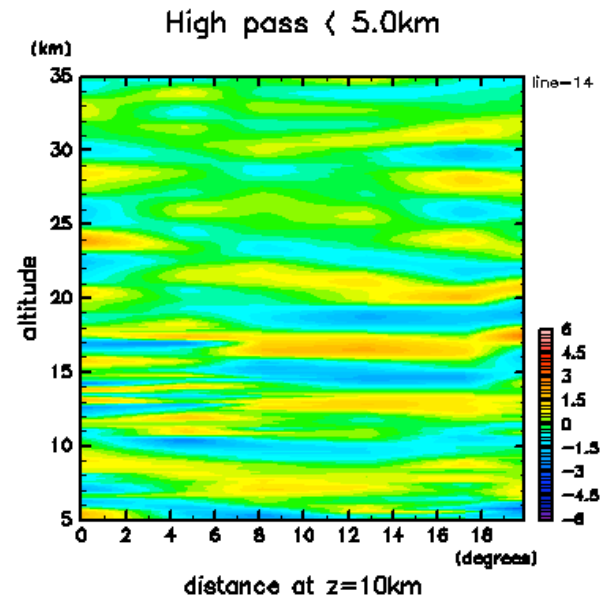
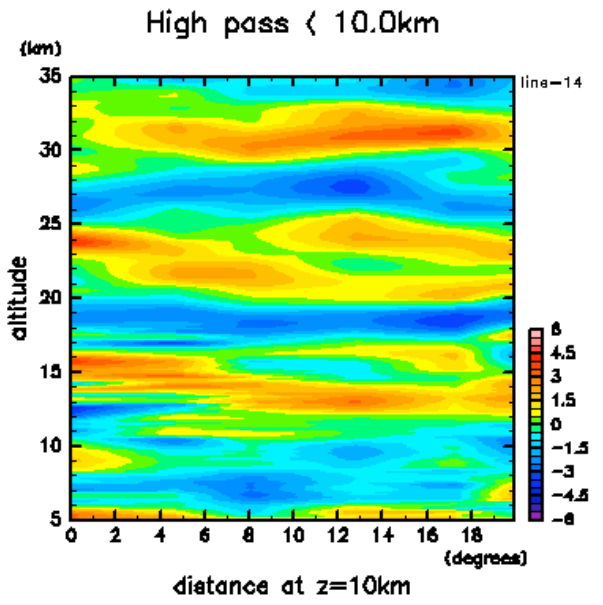
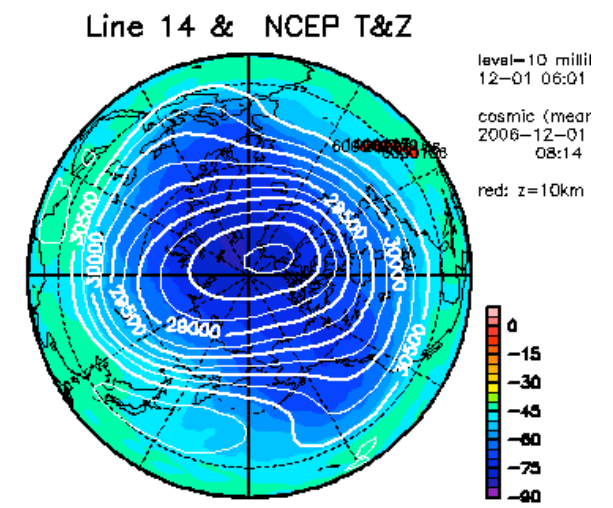
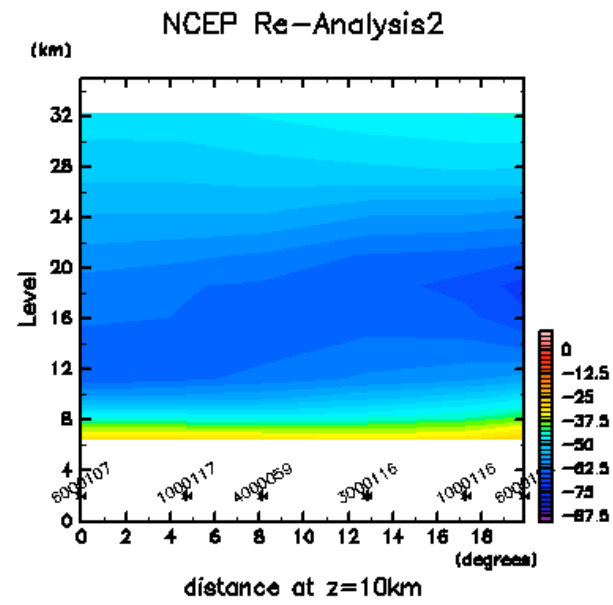
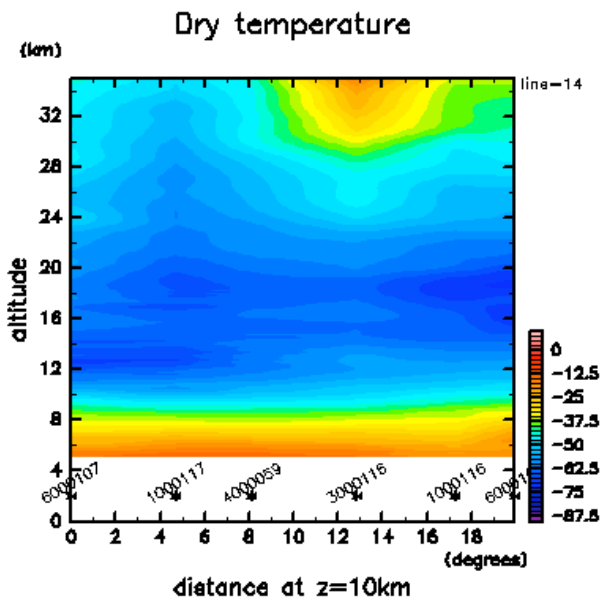


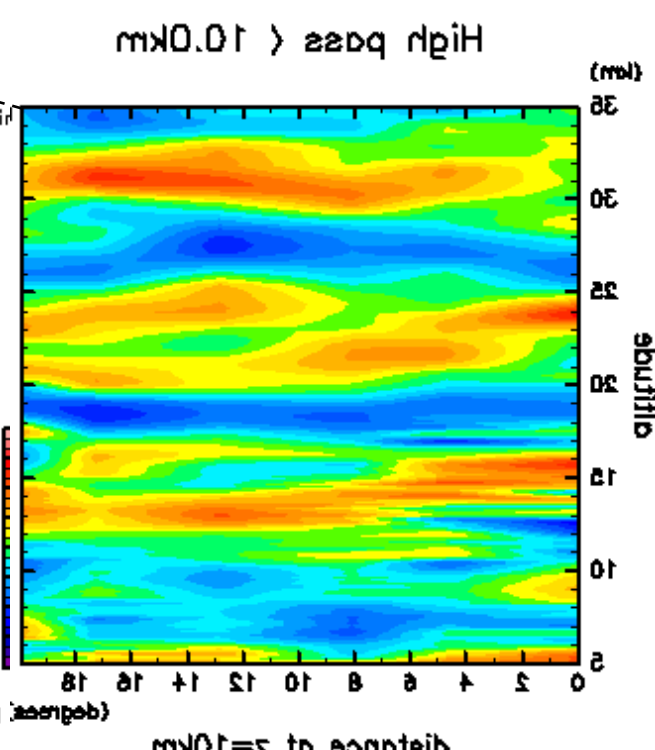
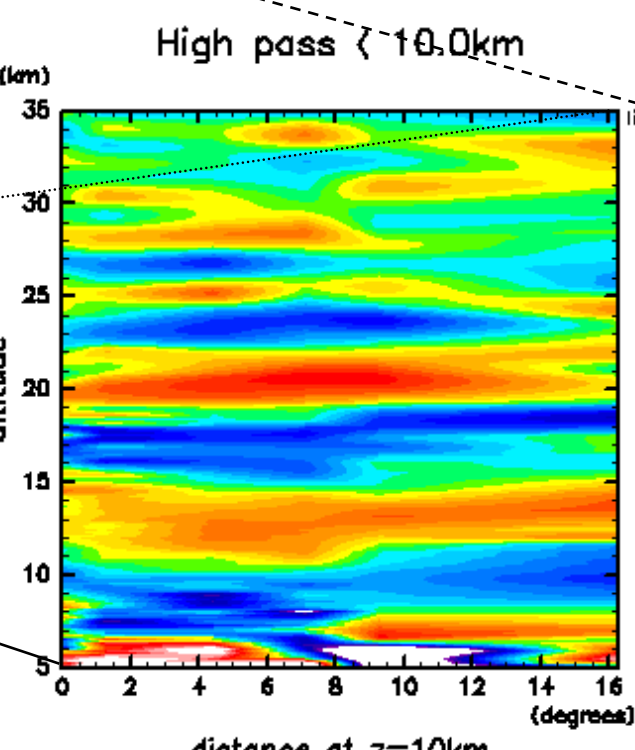
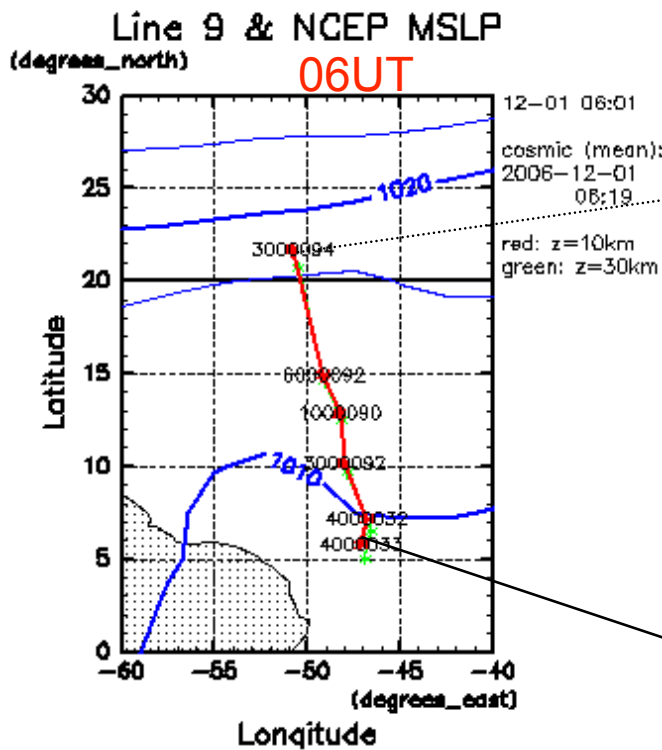
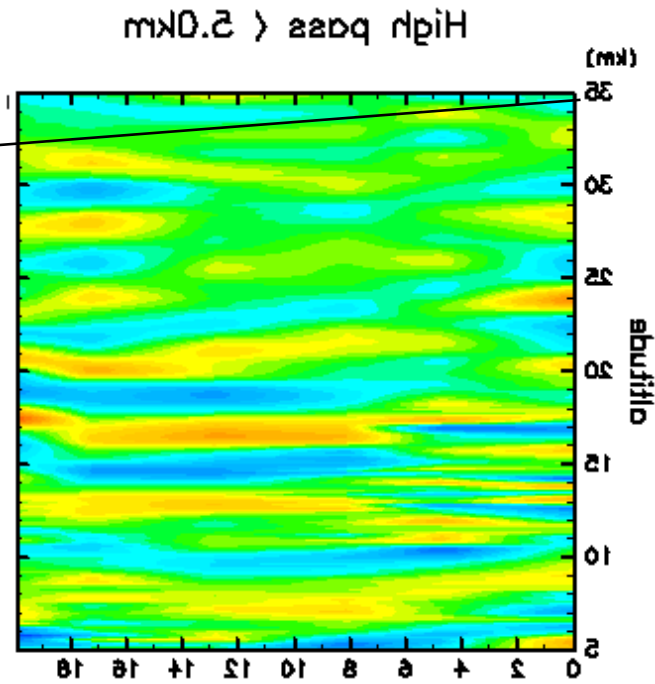
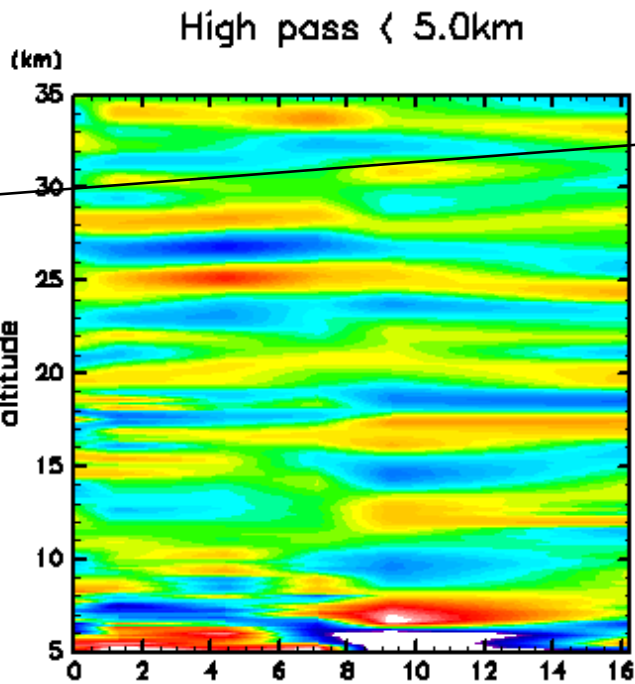
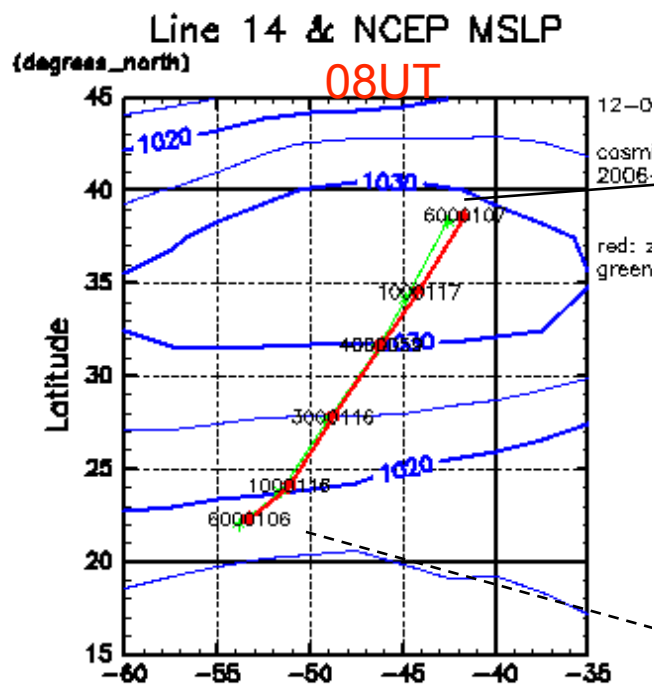
day:2006.320 leo ids:1346



Alexander et al 2008

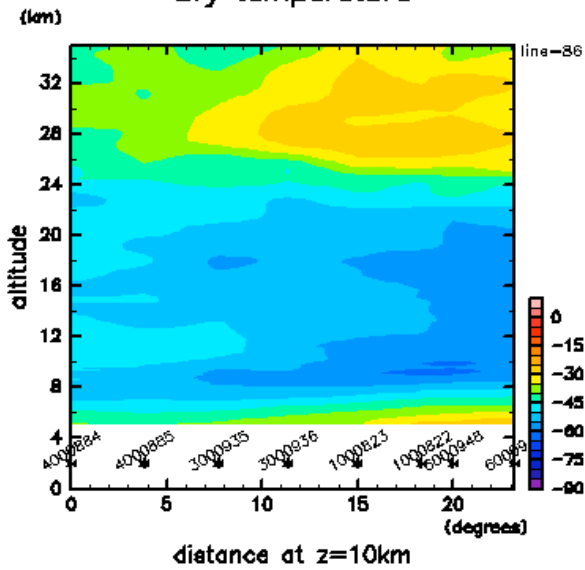




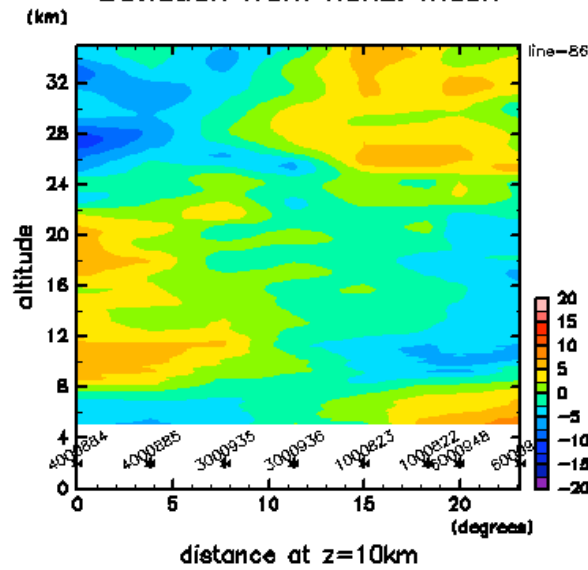


East \longleftrightarrow West

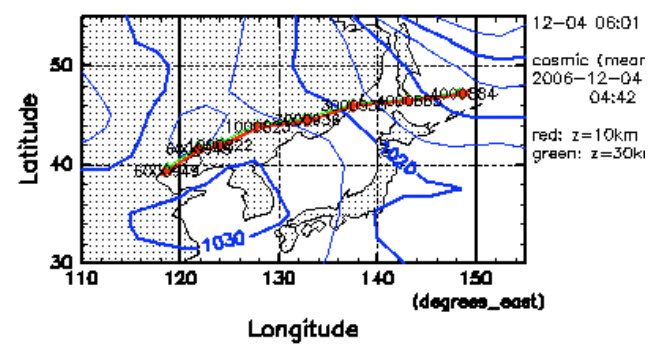
Dry temperature



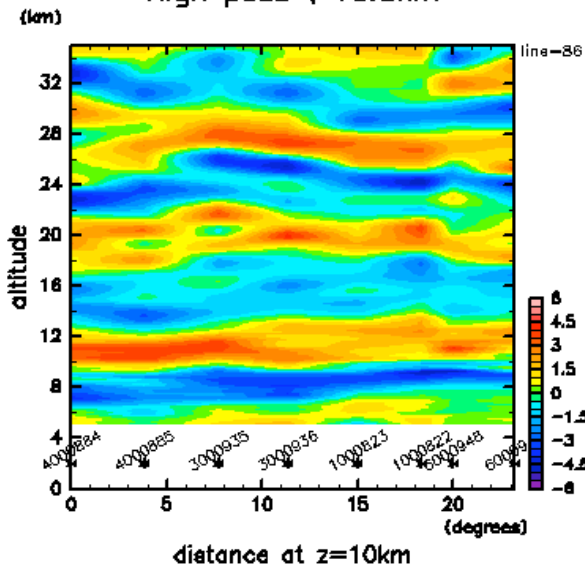
Deviation from hariz. mean



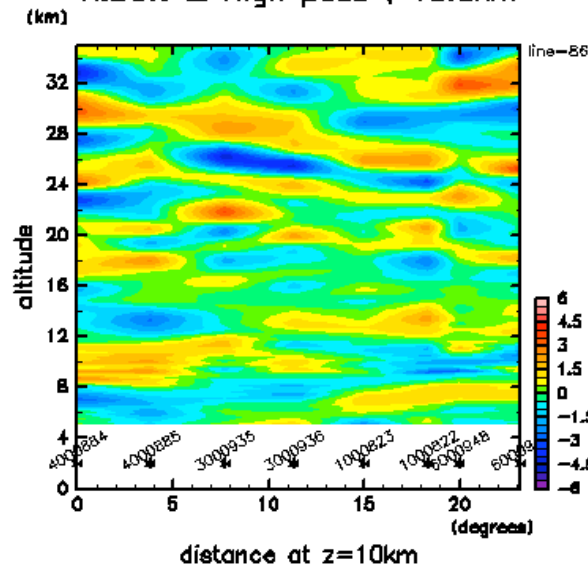
Line 86 & NCEP MSLP



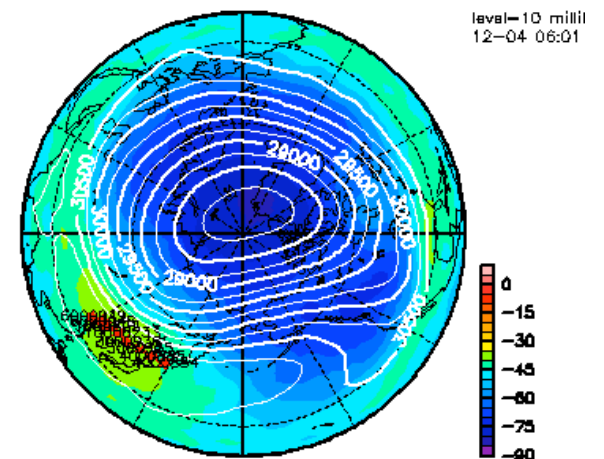
High pass < 10.0km



H.Dev. & High pass < 10.0km

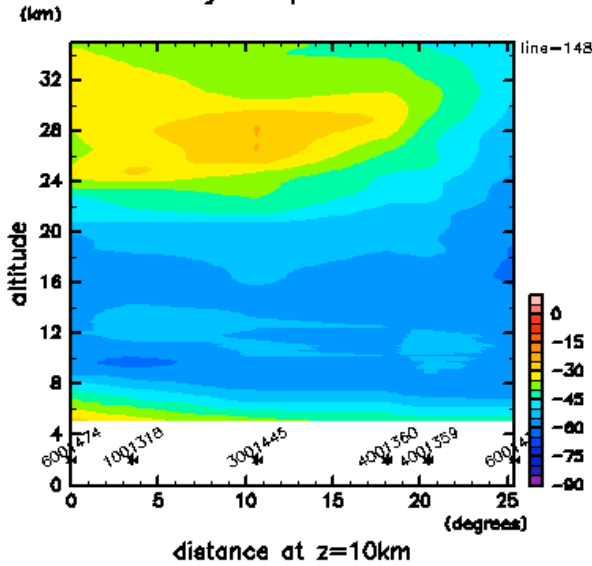


Line 86 & NCEP T&Z

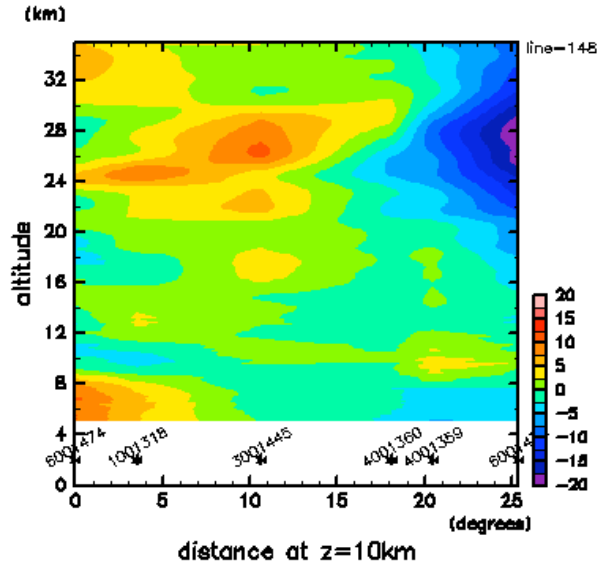


SSW ← → NNE

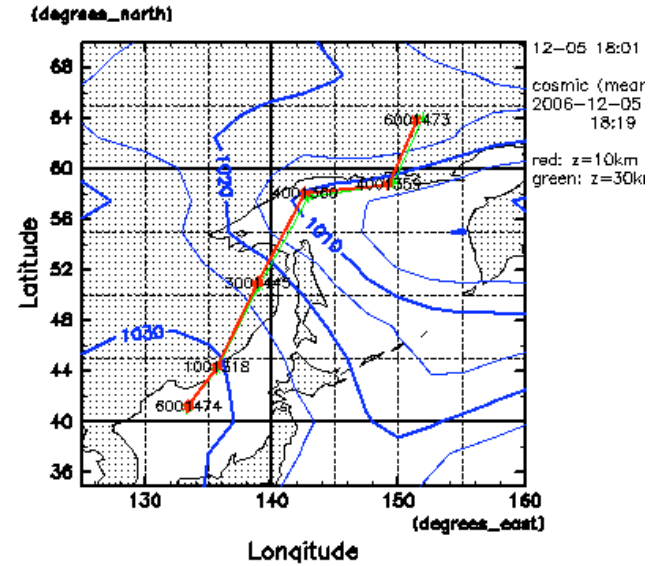
Dry temperature



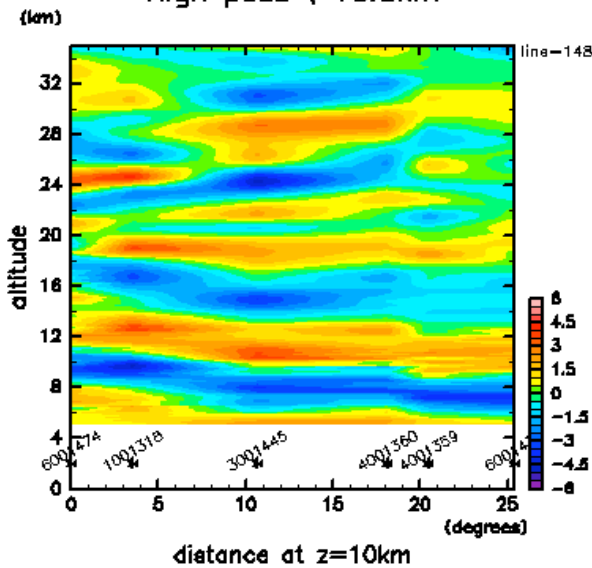
Deviation from horiz. mean



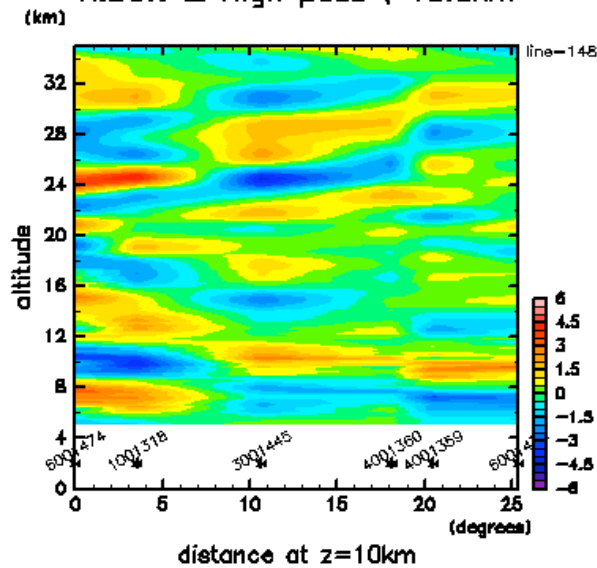
Line 148 & NCEP MSLP



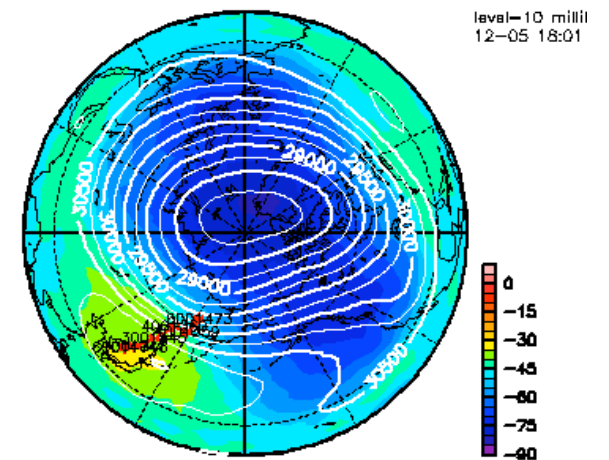
High pass < 10.0km



H.Dev. & High pass < 10.0km

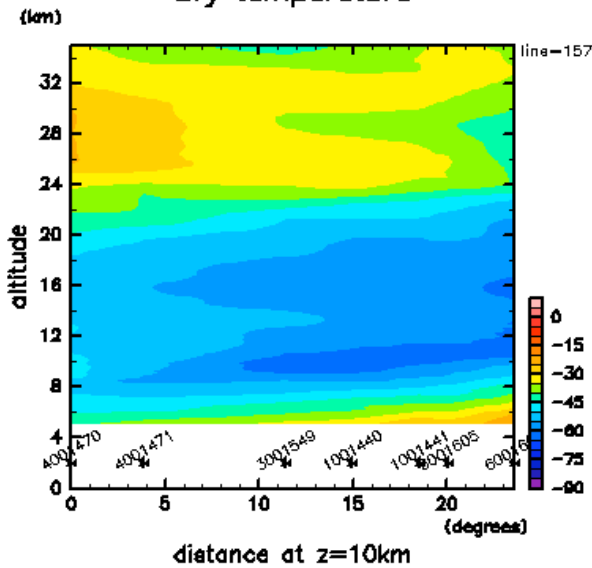


Line 148 & NCEP T&Z

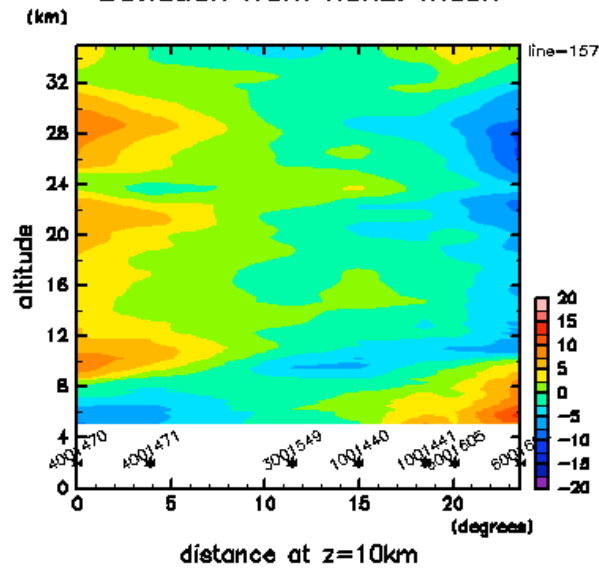


East ← → West

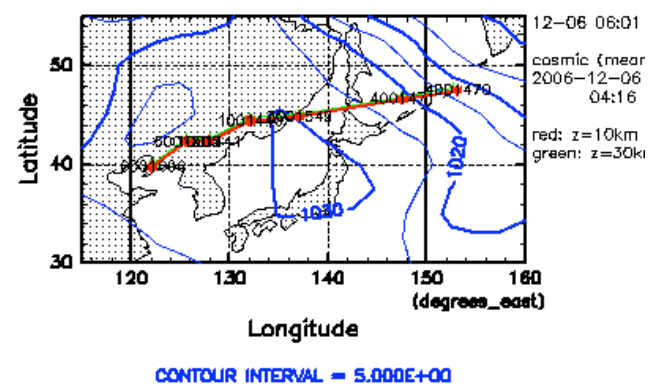
Dry temperature



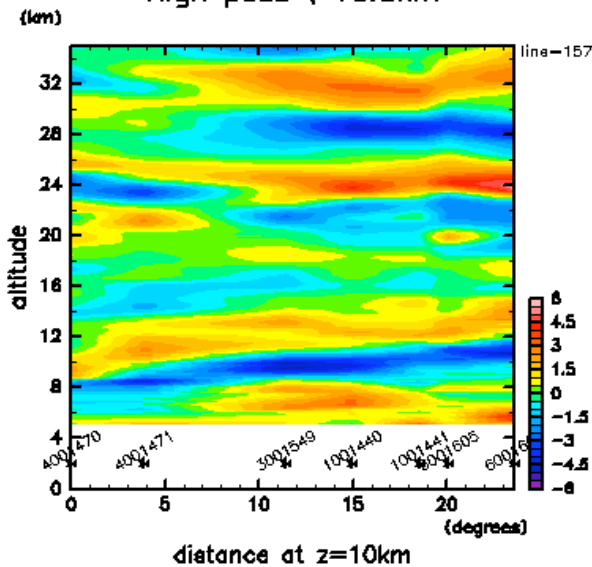
Deviation from hariz. mean



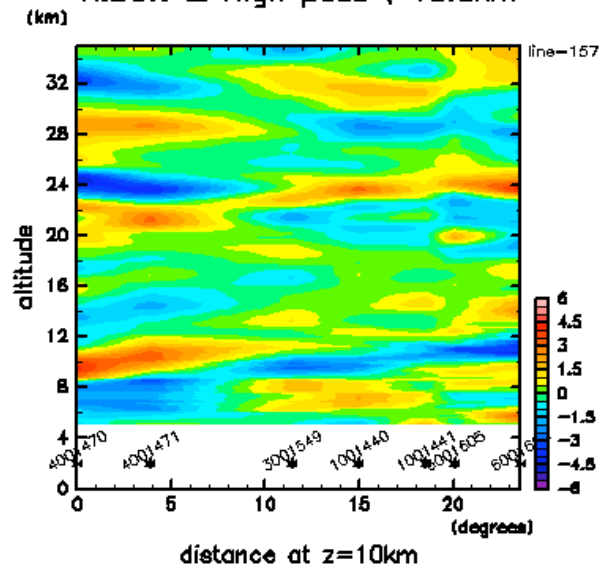
Line 157 & NCEP MSLP



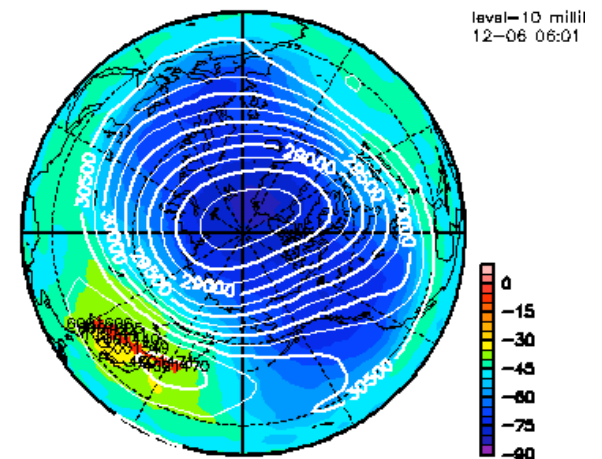
High pass < 10.0km



H.Dev. & High pass < 10.0km

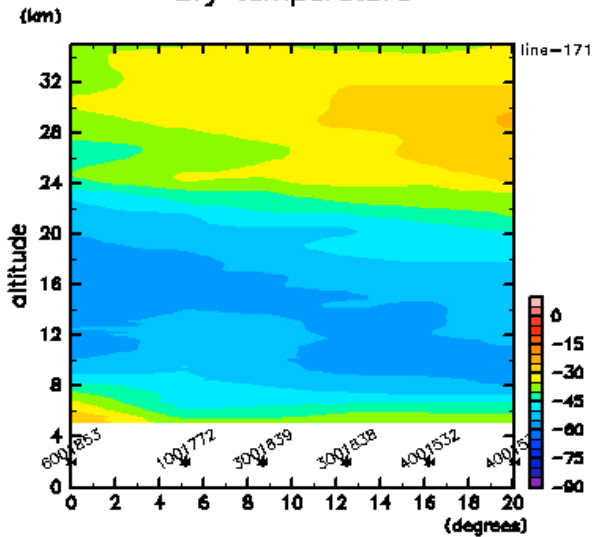


Line 157 & NCEP T&Z



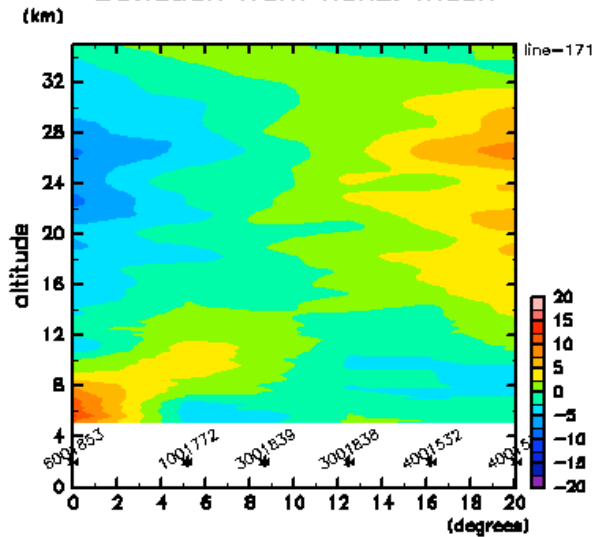
West \longleftrightarrow East

Dry temperature



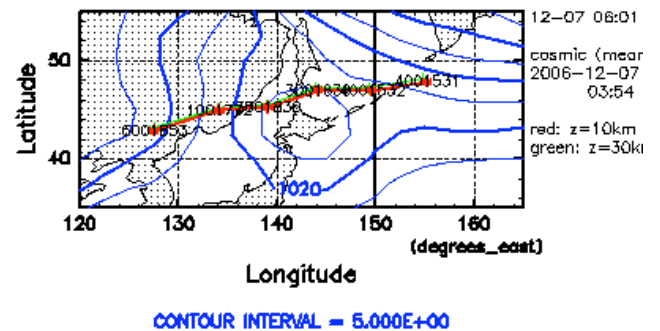
distance at z=10km

Deviation from hariz. mean

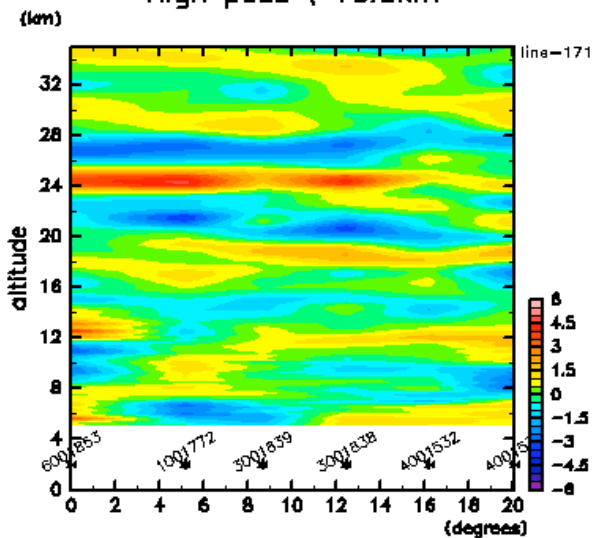


distance at z=10km

Line 171 & NCEP MSLP

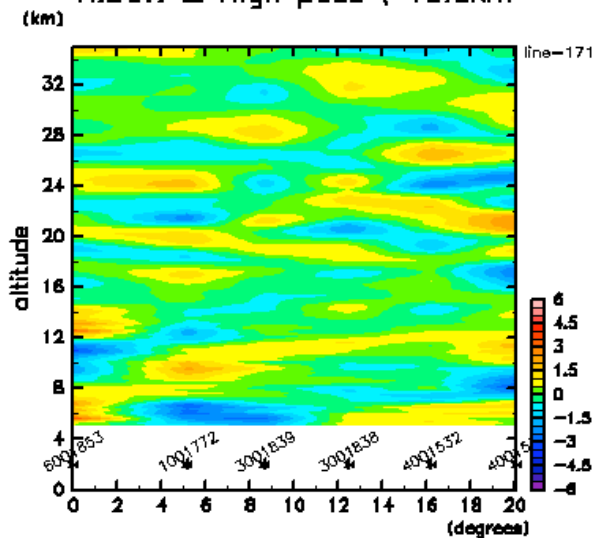


High pass < 10.0km



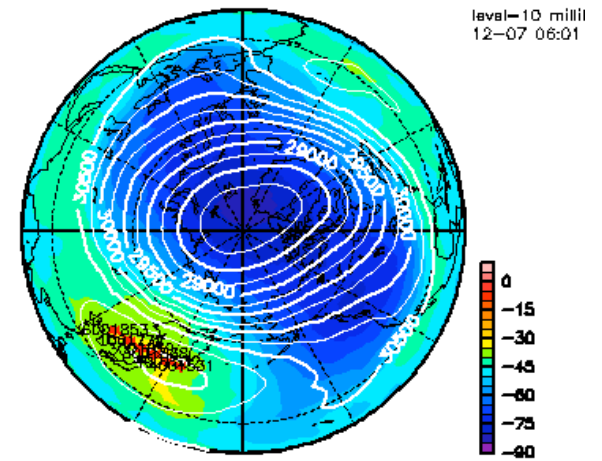
distance at z=10km

H.Dev. & High pass < 10.0km



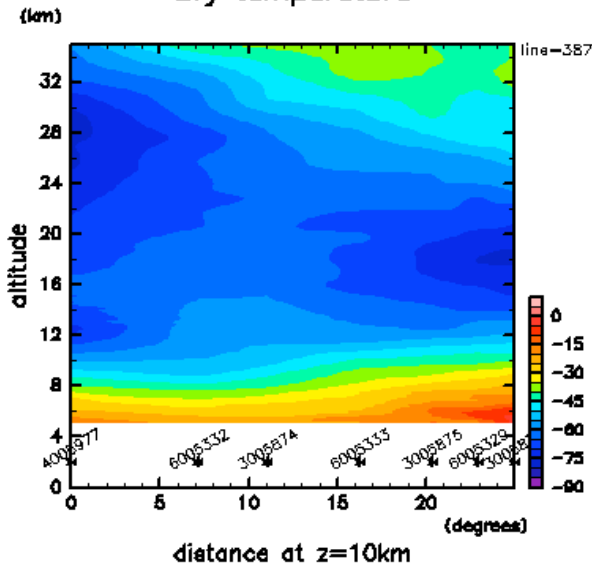
distance at z=10km

Line 171 & NCEP T&Z

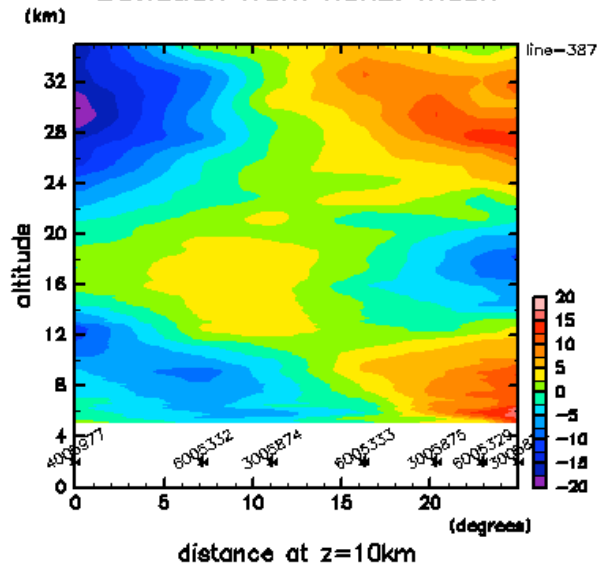


NNW ← → SSE

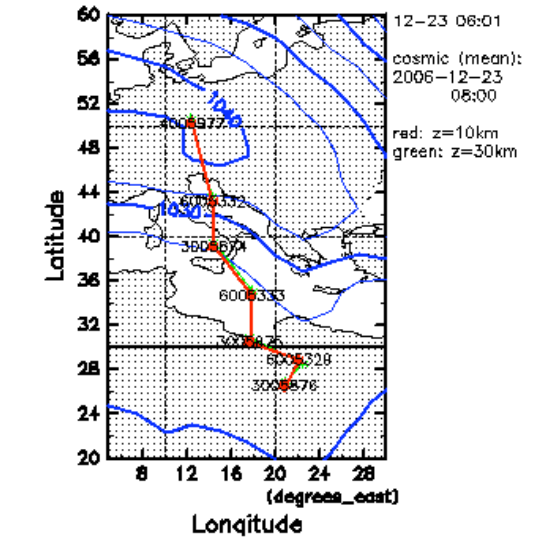
Dry temperature



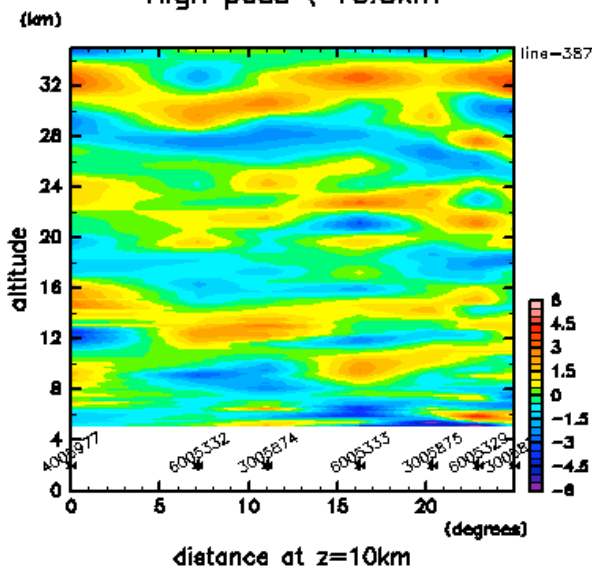
Deviation from horiz. mean



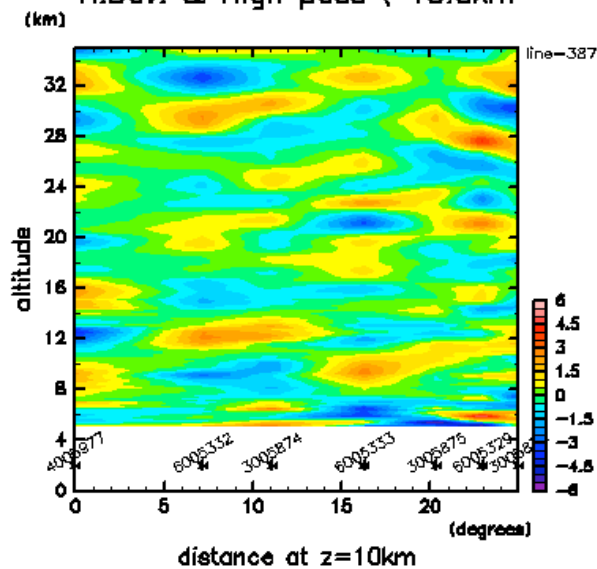
Line 387 & NCEP MSLP



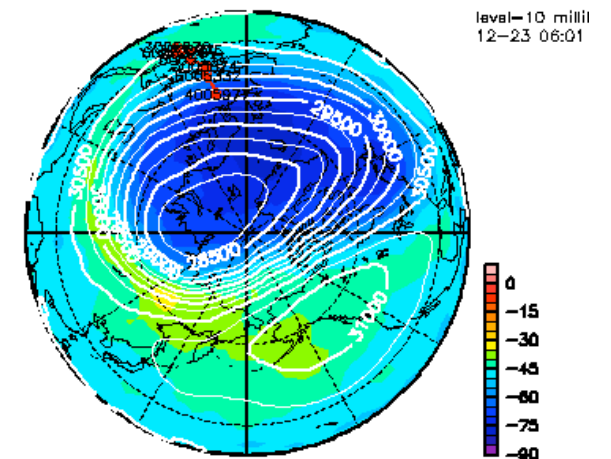
High pass < 10.0km



H.Dev. & High pass < 10.0km

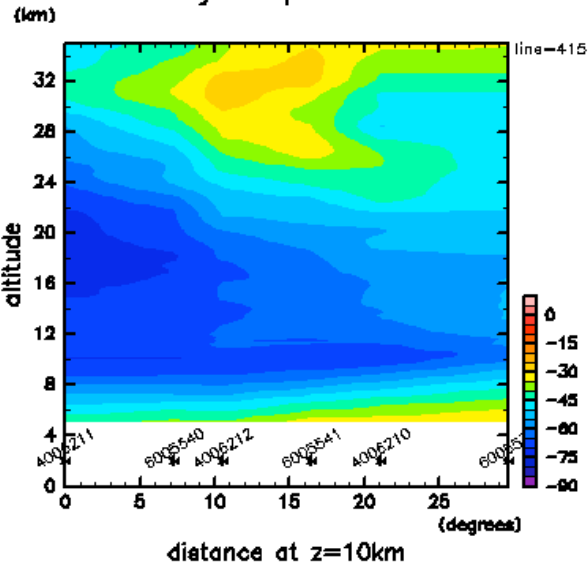


Line 387 & NCEP T&Z

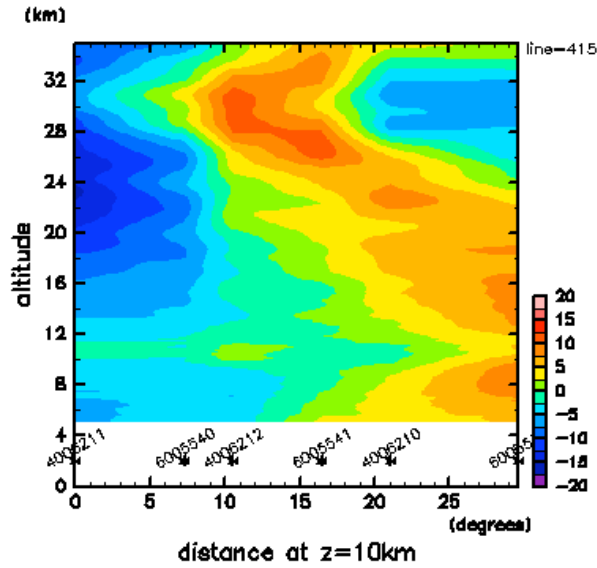




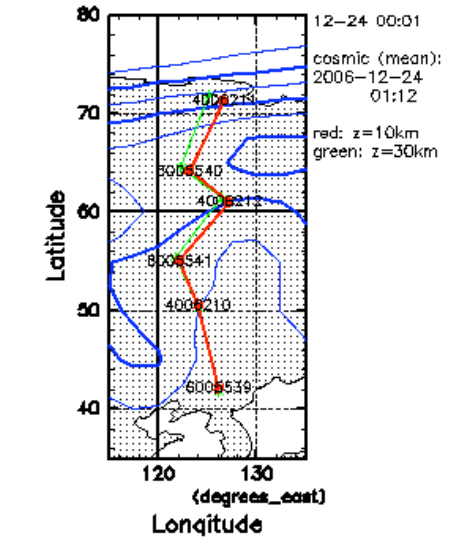
Dry temperature



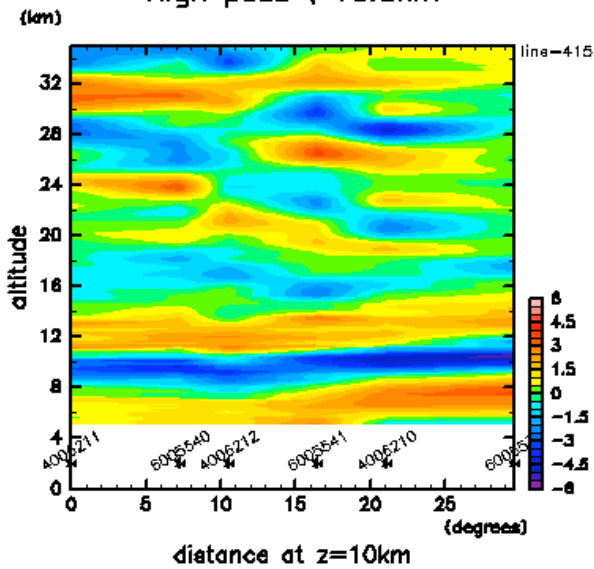
Deviation from horiz. mean



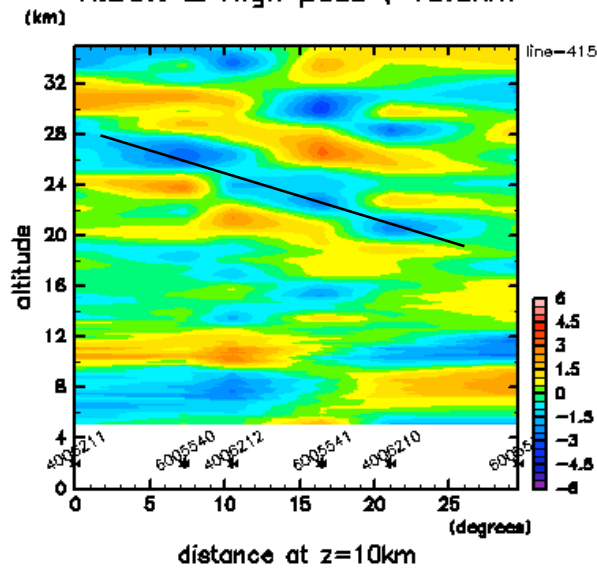
Line 415 & NCEP MSLP
(degrees_north)



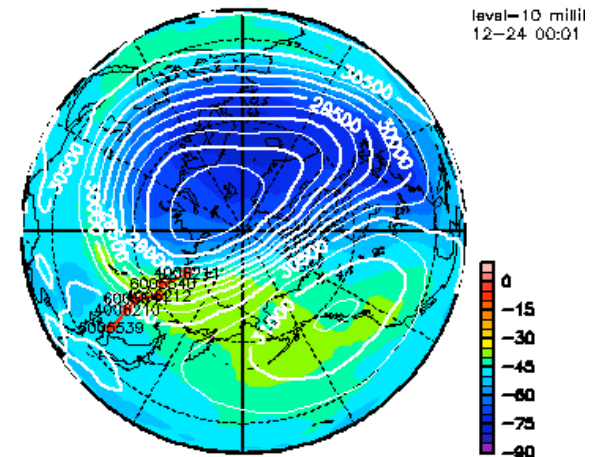
High pass < 10.0km

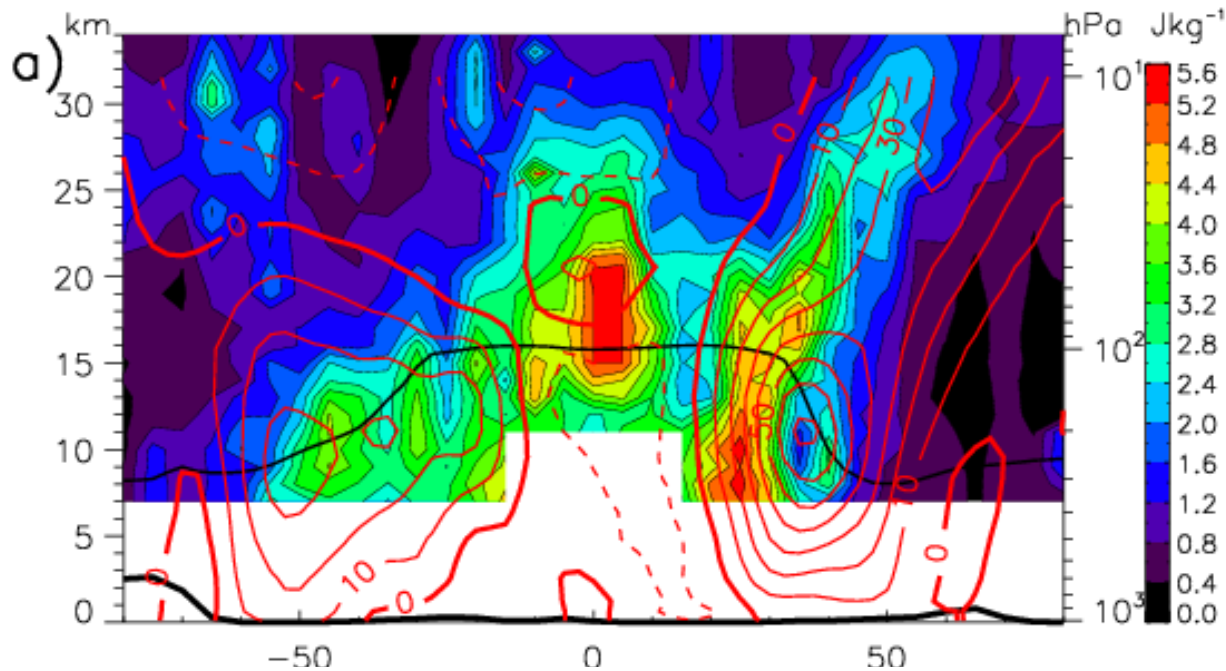


H.Dev. & High pass < 10.0km



Line 415 & NCEP T&Z





COSMIC PE 140E, 12 – 18 Dec 2006 (as before)

- Strong winter time sub-tropical jet
- Large PE from mid-troposphere up to PNJ

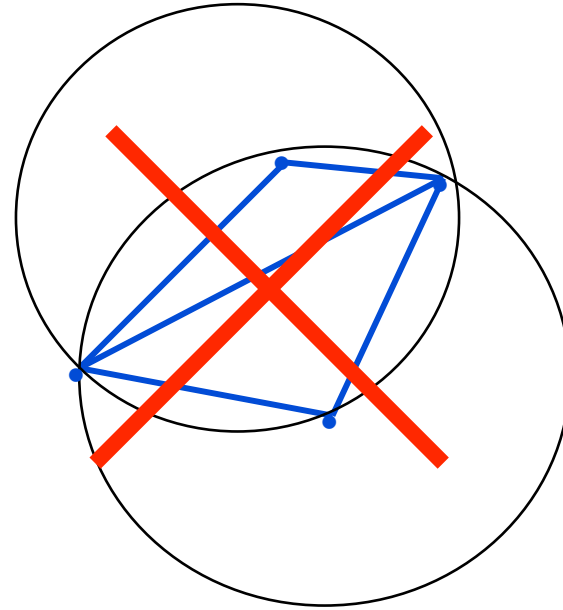
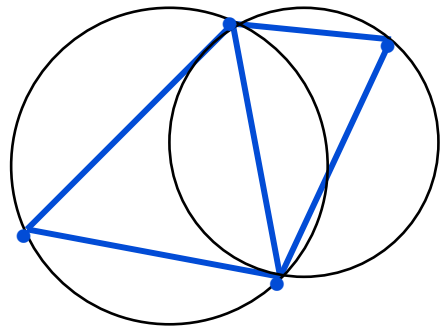
AGCM PE 140E, 1 – 7 Jan (similar wind conditions to COSMIC)

- PE from waves with periods 6hr – 1 month, $\lambda_z < 7\text{km}$, $380 < \lambda_x < 40,000\text{km}$
- Note different colour scale
- Vectors show meridional and vertical energy fluxes due to $\lambda_z < 7\text{km}$
- Most mid-lat stratospheric PE due to the jet (upward vectors)

by Alexander and Tsuda

予備スライド

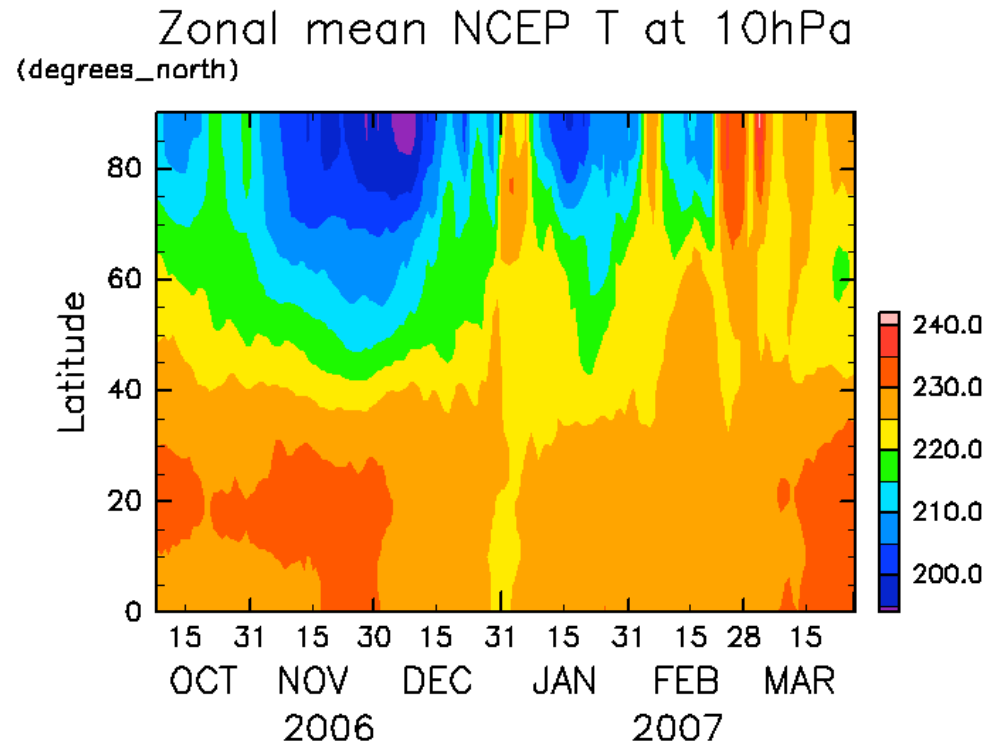
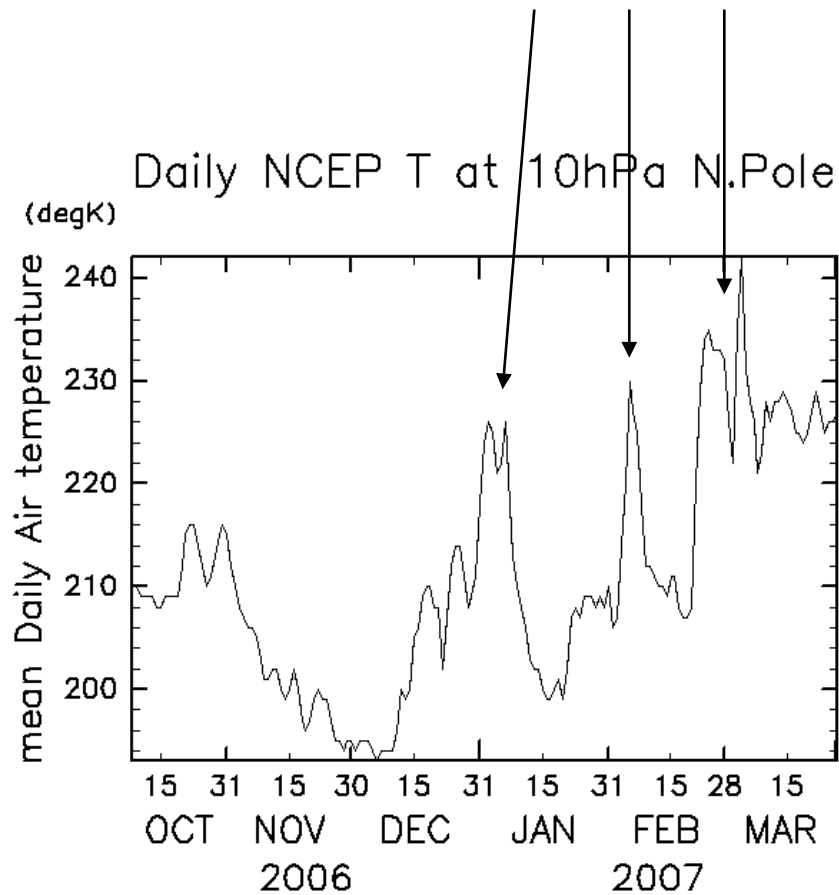
Delaunay triangulation

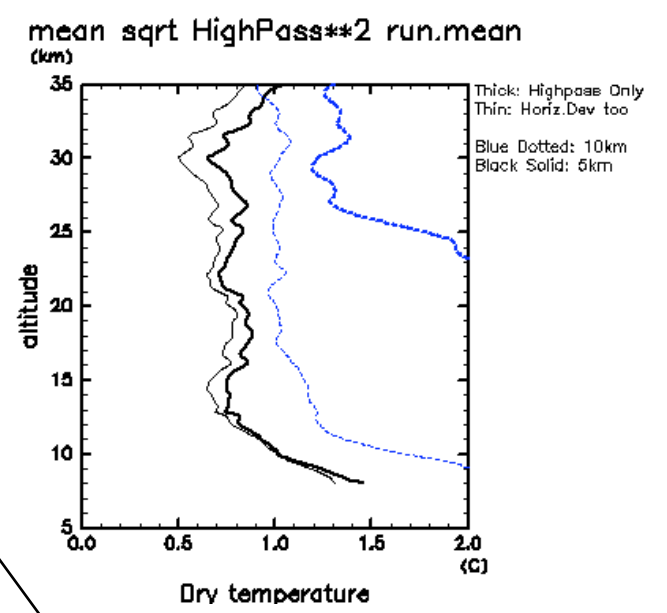
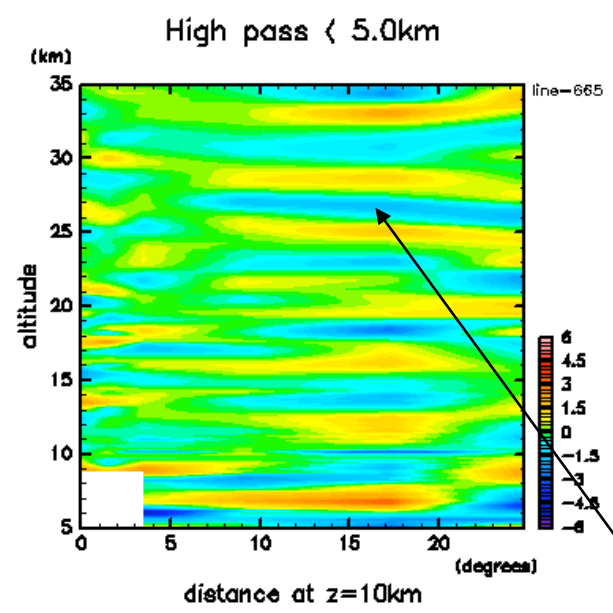
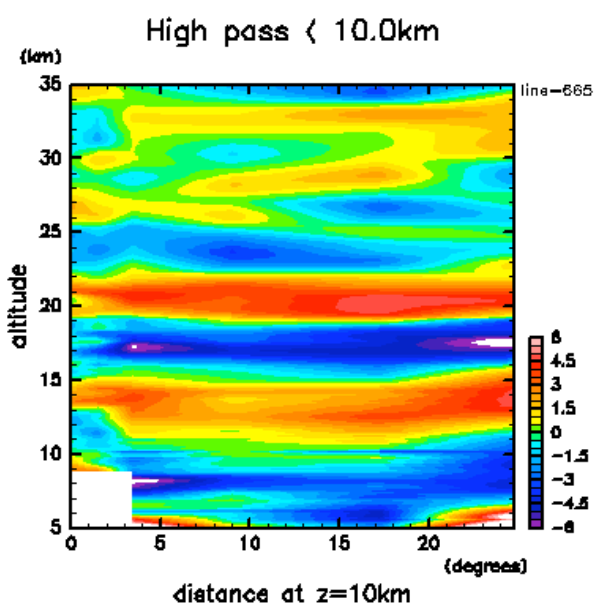
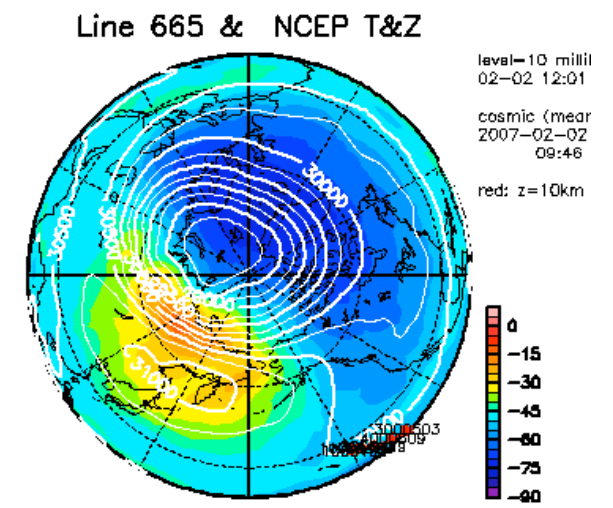
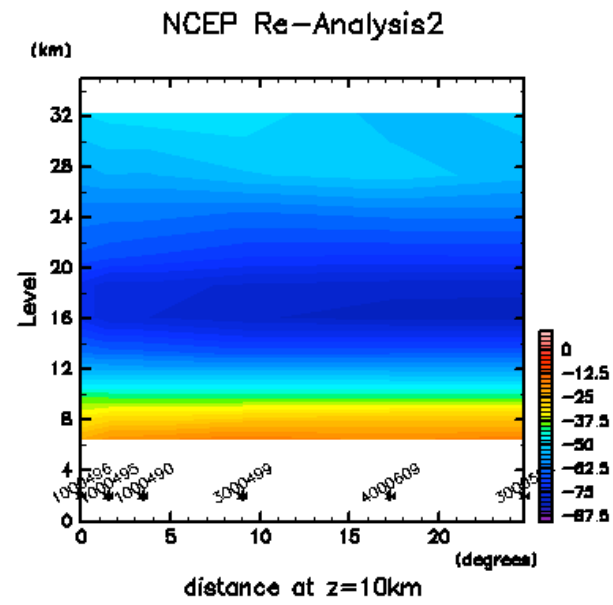
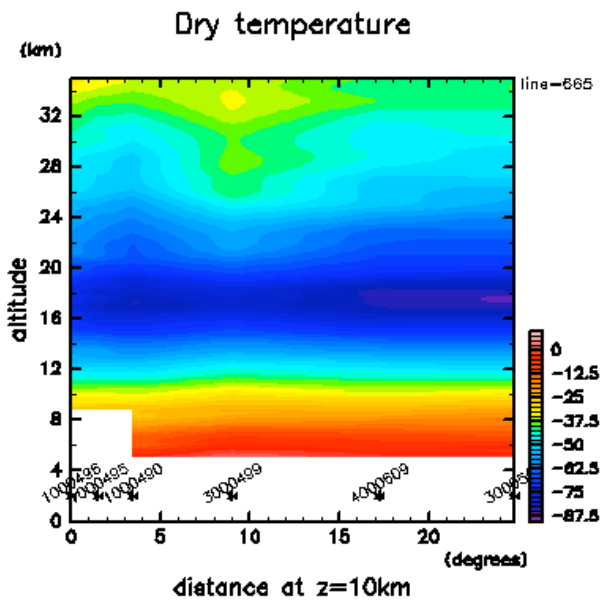


Allow no other point in
the outer circles of all
triangles

N.H. Stratosphere of the period

- 3 major sudden warmings





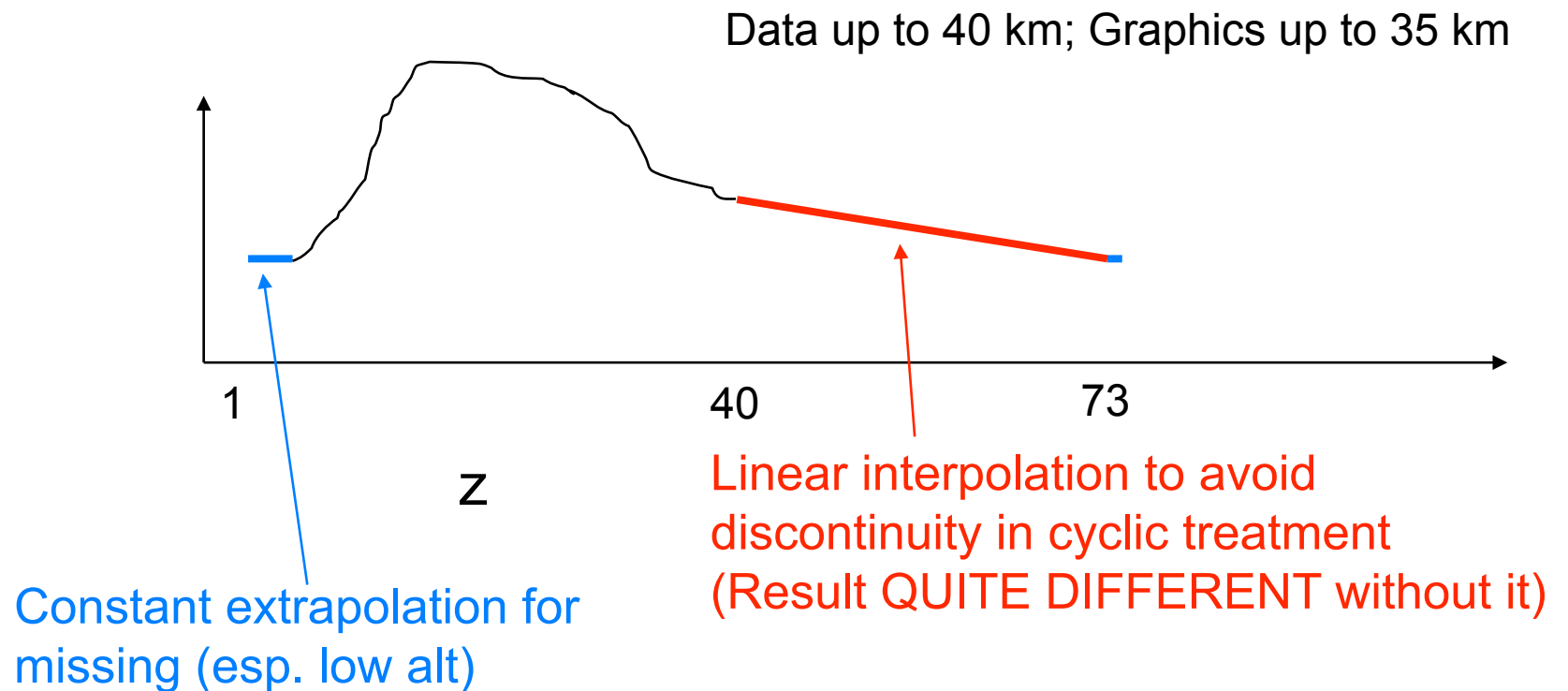
20N

5N

Short λ_z wave

Vertical high-pass filtering

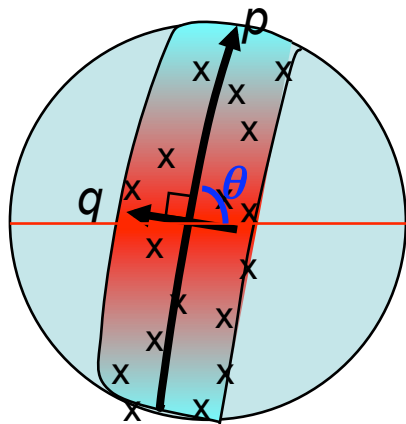
- Based on FFT
- Padding & Extending to avoid leak from cyclic discontinuity



座標変換

- 衛星進行方向(p とする)とそれに直交する方向(q)に軸を取る。
 - 通常のswathデータ同様
 - ただし、軌道情報がなくても、**掩蔽点の位置だけから求まるよう工夫**(\because 衛星が複数なので、軌道パラメタも一意でない。利点:解析的。軌道データがない場合でも適用できる)。

$$\begin{aligned}\sin q &= \sin \phi \cos \theta - \sin \lambda' \cos \phi \sin \theta \\ \cos p &= \frac{\sin \lambda' \cos \phi}{\sqrt{1 - \sin 2q}} \\ \sin p &= \frac{\sin \lambda' \cos \phi \cos \theta + \sin \phi \sin \theta}{\sqrt{1 - \sin 2q}}\end{aligned}$$



ここで $\lambda' = \lambda - \lambda_0 - \alpha\omega^{-1}p$ (衛星軌道面と共に回転する座標系での経度)

p について再帰的なので、反復法。

λ_0 も $\overline{p \cos q}$ をゼロに近づける反復より求める。